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# rthquakes: Observations eling



# reat Earthquakes c Radiation

rre Vilotte,

Collège de France, November 30th 2017







## Seismic radiation and earthquake size

Since the early days of Seismology (1930-1940), it has been well known that the greater the size of an earthquake, the more efficiently longer-period waves are generated.





# Far field radiation from a dislocation source









$$u(f) \propto \frac{\Omega_0}{1 + \left(\frac{f}{f_c}\right)^2}$$

 $\beta =$ S-wave speed a =source radius  $f_c = 0.3724 \frac{\beta}{-1}$ 

"k-squared" models: self-similar slip on the fault at high wavenumbers k. (Herrero and Bernard, 1994; Causse et al., 2010; Ruiz et al., 2011;)

Slip [m] 0.5 1.0







PEAK GRO 0% PROBABILITY OF EXCEED



PEAK GROUND ACCELERATION (m/s<sup>2</sup>)

10% PROBABILITY OF EXCEEDANCE IN 50 YEARS, 475-year return period

.6	2	.4	3	.2	4	.0	0 4.8		
ERATE		HIGH				VERY HIGH			
ZARD		HAZARD				HAZARD			

# The 2011 Tohoku earthquake (Mw 9.1): slip from kinematic modeling













# Tohoku: a complex rupture when seen through strong motion



40<sup>°</sup>

39°

38

37

36

35

### Fault slip imaged by teleseismic data A compact rupture, close to the trench.

Satriano et al., EPSL (2014)



### Ground acceleration at regional scale

A complex rupture: at least 5 sub-events, close to the coast.

Lee et al., 2011 (redrawn)



# Teleseismic imaging of high-frequency radiation: back projection

![](_page_8_Picture_1.jpeg)

![](_page_8_Figure_2.jpeg)

![](_page_8_Figure_5.jpeg)

![](_page_8_Figure_6.jpeg)

![](_page_8_Figure_7.jpeg)

# High-frequency sources and strong ground motion

![](_page_9_Figure_1.jpeg)

## Fault slip hand high-frequency sources imaged by teleseismic data

Satriano et al., EPSL (2014)

Lee et al., 2011 (redrawn)

![](_page_9_Figure_6.jpeg)

## Northwest Japan: Structural heterogeneity and seismic asperities

![](_page_10_Figure_1.jpeg)

![](_page_10_Picture_2.jpeg)

Satriano et al., EPSL (2014)

![](_page_10_Figure_4.jpeg)

Background seismicity and large earthquake ruptures (M 7-8) inform us on the mechanical state of the plate interface.

This information is useful to forecast the rupture and radiation properties of future large earthquakes.

145°

![](_page_10_Figure_8.jpeg)

# Other examples: downdip HF emission for large subduction earthquakes

### Maule (Chile) 2010 Mw 8.8

![](_page_11_Figure_2.jpeg)

# The 2015 Mw 7.8 Gorkha earthquake

![](_page_12_Figure_1.jpeg)

Grandin, Vallée, Satriano et al., GRL (2015)

![](_page_12_Picture_3.jpeg)

# 04-25-2015 Sat 115-56-28

यही नव ग्रह

धूप पाईन्छ 8 सन

![](_page_13_Figure_1.jpeg)

# DABALI

![](_page_13_Picture_3.jpeg)

# Gorkha earthquake: intensity and ground acceleration above the rupture

![](_page_14_Figure_1.jpeg)

![](_page_14_Figure_2.jpeg)

![](_page_14_Figure_3.jpeg)

![](_page_14_Figure_4.jpeg)

![](_page_14_Figure_5.jpeg)

![](_page_14_Figure_6.jpeg)

# Gorkha earthquake: kinematic slip model

![](_page_15_Figure_1.jpeg)

![](_page_15_Figure_3.jpeg)

![](_page_15_Figure_6.jpeg)

## 2. Static GPS displacement

![](_page_15_Figure_8.jpeg)

## 2. Local ground motion (up to 0.1 Hz)

![](_page_15_Figure_10.jpeg)

### 3. Teleseismic waveforms (up to 0.125 Hz)

![](_page_15_Figure_12.jpeg)

![](_page_15_Figure_13.jpeg)

# LF earthquake in Kathmandu, HF at the base of the rupture

![](_page_16_Figure_1.jpeg)

![](_page_16_Figure_2.jpeg)

![](_page_16_Figure_3.jpeg)

High-Frequency peaks (1-4 Hz) from back projection are at the down dip edge of the rupture

The kinematic model (inversion of data up to 0.1Hz), well explains near-field data at frequencies up to 1Hz

Grandin, Vallée, Satriano et al., GRL (2015)

![](_page_16_Picture_9.jpeg)

## **Rupture properties and fault zone heterogeneity**

![](_page_17_Figure_1.jpeg)

Rupture abruptly stops down-dip, with highfrequency radiation, where the plate interface strongly decouples.

Almost all the seismicity before 2015 Gorkha earthquake is in this zone: stress concentration?

Rupture gradually slows down up-dip, with no high-frequency radiation.

No seismicity in this zone before 2015: fully locked and/or strengthening friction?

Larger ruptures (M 8+) coud reach the surface but only with low-frequency radiation?

Grandin, Vallée, Satriano et al., GRL (2015)

![](_page_17_Picture_9.jpeg)

![](_page_17_Picture_10.jpeg)

![](_page_18_Figure_1.jpeg)

# Fault heterogeneity and high-frequency radiation: synthetic tests

![](_page_19_Picture_1.jpeg)

Marina Corradini: PhD project

Homogenous: "Haskell" fault

![](_page_19_Figure_4.jpeg)

hypocenter

# Synthetic test: homogenous model

![](_page_20_Figure_1.jpeg)

![](_page_20_Figure_2.jpeg)

![](_page_20_Figure_3.jpeg)

# Synthetic test: variable rise time

![](_page_21_Figure_1.jpeg)

hypocenter

![](_page_21_Figure_3.jpeg)

slip acceleration Time (s)

**Coherent HF radiation** 0.0 0.5 slip-rate 60 -40 -20 -0 --200 100 -100200 0 slip-rate EU d(slip-rate)/dx Along-strike direction (km) High-frequency radiation linked to space-time variability of slip rate on the fault

![](_page_21_Picture_6.jpeg)

![](_page_21_Picture_7.jpeg)

# Conclusions

Variability of fault-zone mechanical properties (e.g., coupling, asperities, geometry) controls the variability of earthquake radiation (high-frequency - low-frequency).

Study of seismicity and deformation before large earthquakes is necessary to characterize those properties and define future large earthquake scenarios.

New imaging techniques like back projection can help in illuminating the complex nature of earthquake rupture. Still need to better understand the link between high-frequency images and rupture process.

![](_page_22_Figure_4.jpeg)

Along-strike direction (km)

![](_page_22_Figure_6.jpeg)

![](_page_22_Figure_7.jpeg)