

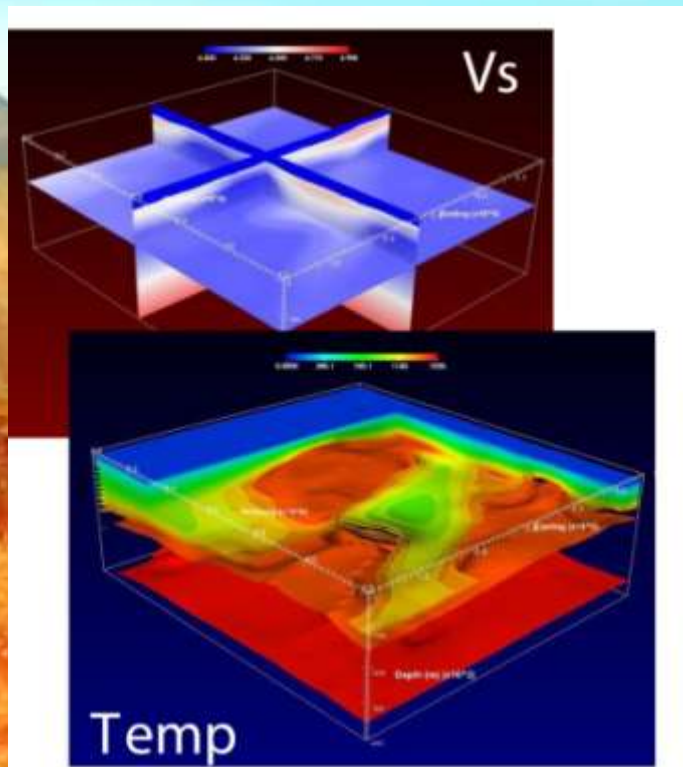


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Integrated geophysical and petrological modelling of the uppermost mantle: forward and inversion approaches



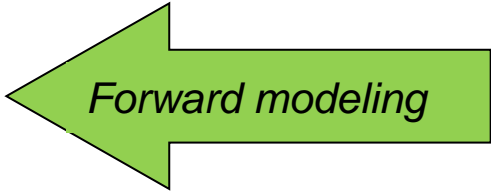
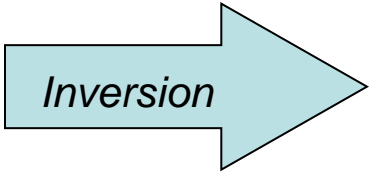
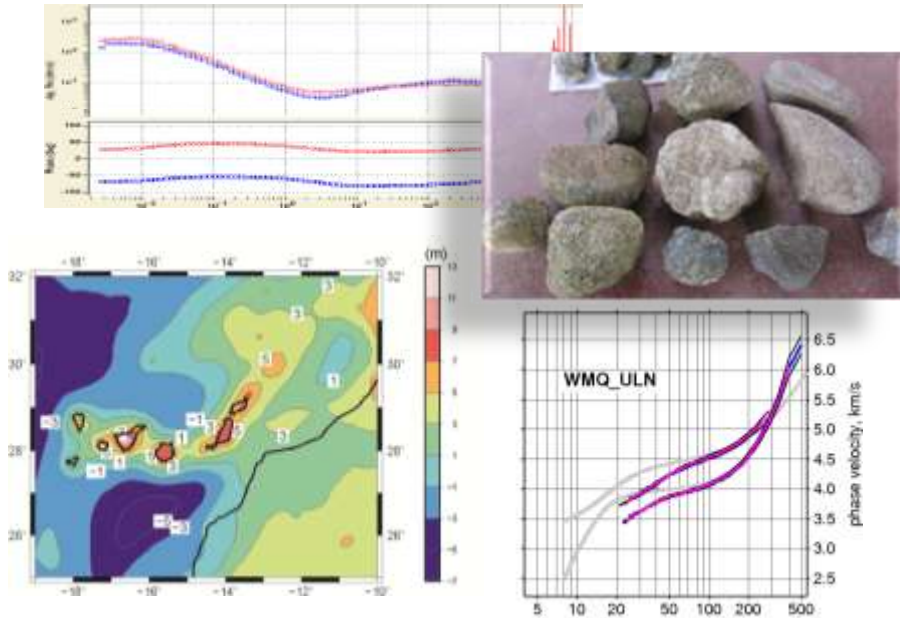
J. Fullea

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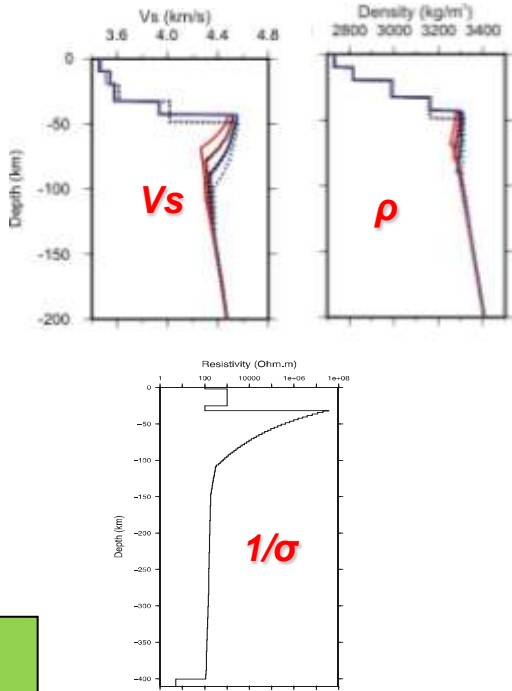
J. C. Afonso, A. G. Jones, S. Lebedev, J. Connolly, Y. Yang, N. Rawlinson, W. Griffin, S. O'Reilly

- **Integrating (self-consistently) geophysical and petrological data to image the lithosphere/uppermost mantle**

Geophysical & petrological data



Geophysical parameters



- **Forward approach**
- **Non-linear probabilistic inversion**

Integrated modelling

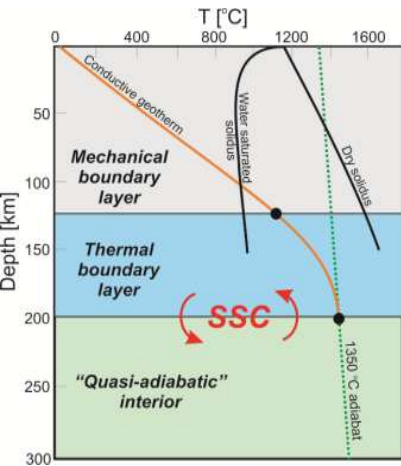


Temperature, Pressure, Composition

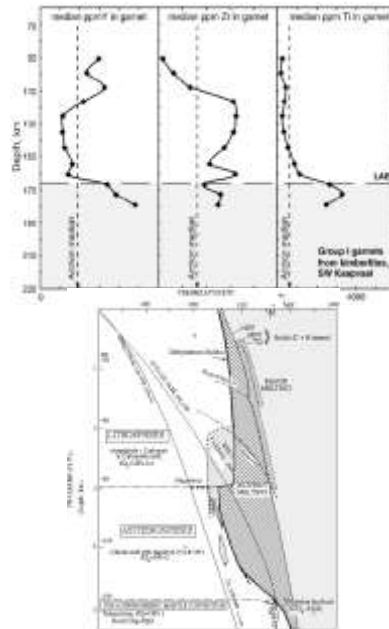
Why integrated modelling...?

A fundamental upper mantle discontinuity for plate tectonics: the lithosphere-asthenosphere boundary (LAB) can be defined as **thermal, mechanical and chemical boundary**

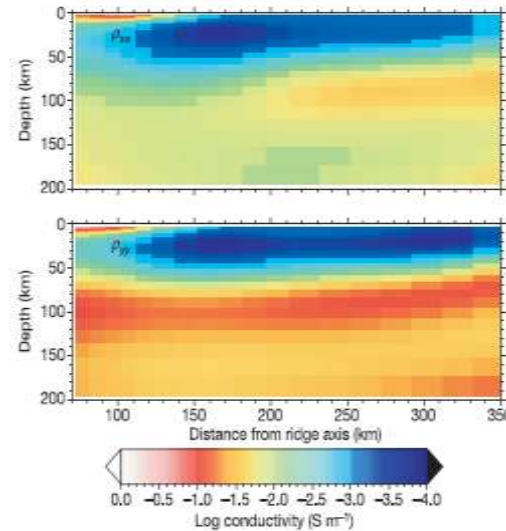
According to the property we focus on (e.g., **temperature, composition, Vs, Vp, anisotropy, electrical conductivity...**) there are many possible “LABs” (e.g. Eaton 2009):



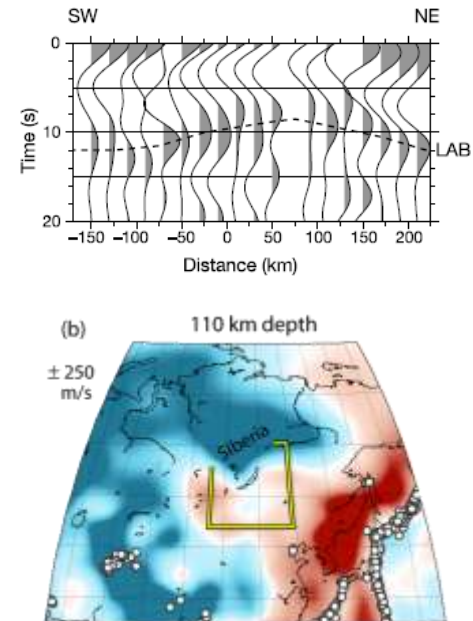
Thermal LAB



**Geochemical/
petrological
LAB**



Electric LAB

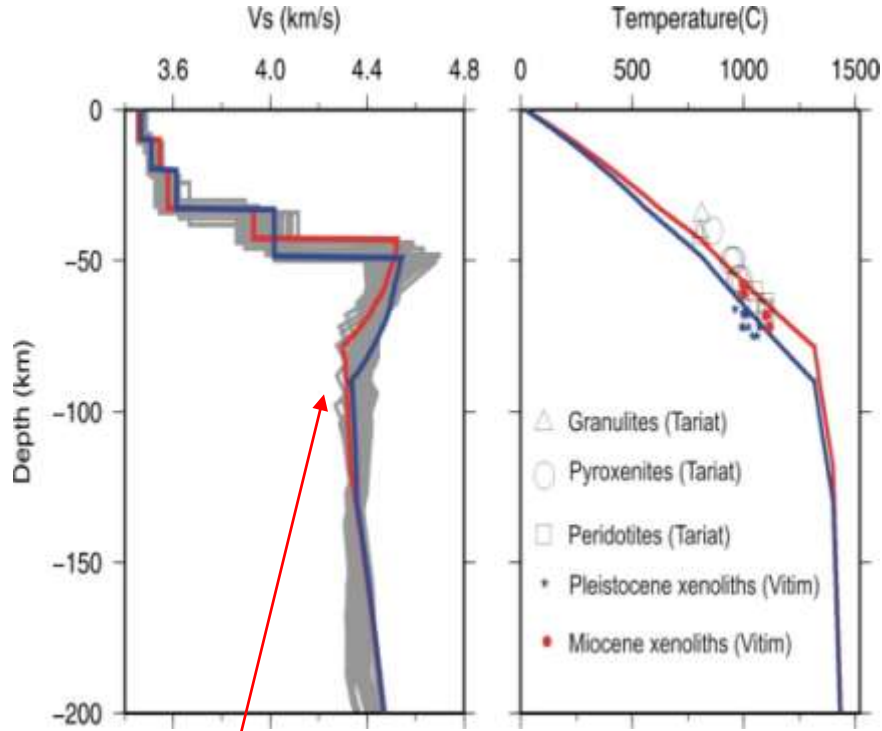


Seismic LAB

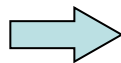
Why integrated modelling...?

Ultimately most of the **LAB** definitions depend upon **temperature/pressure and composition**

Surface waves

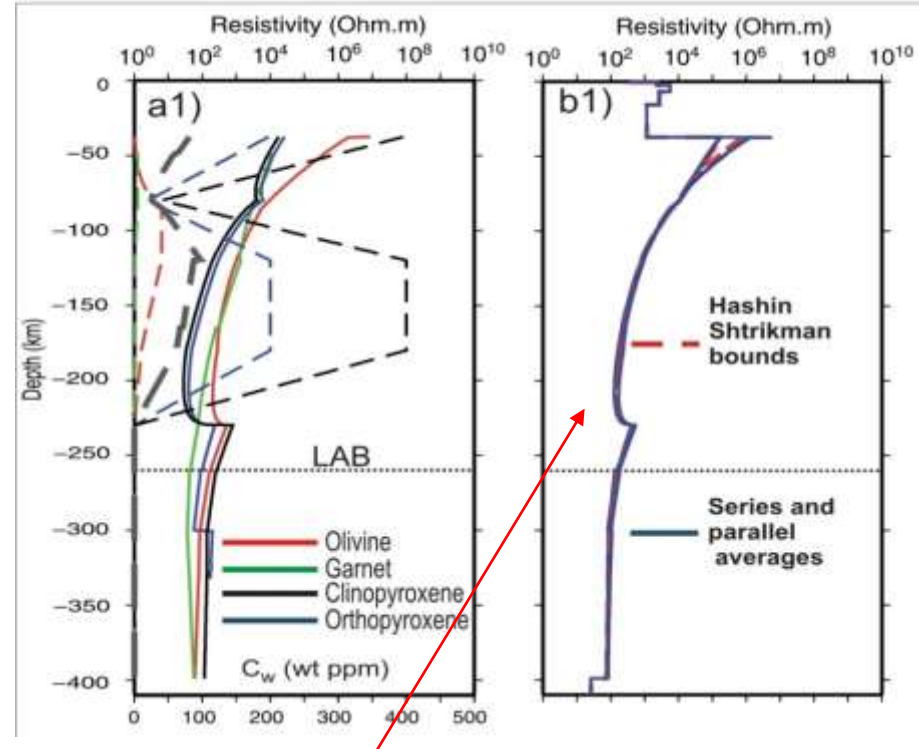


Low Velocity Zone

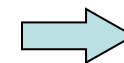


Temperature+Pressure+Attenuation

Magnetotelluric



Low Resistivity Zone



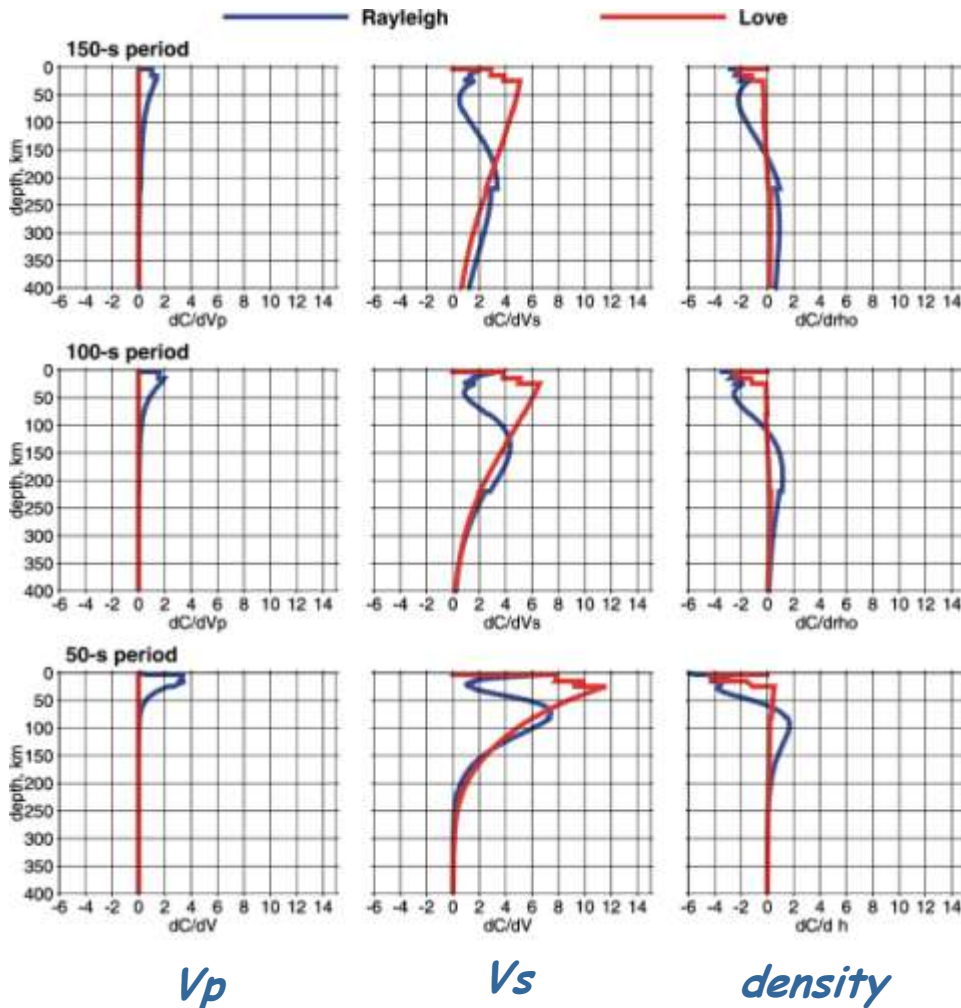
Temperature+Water

→ Unify LAB's criteria

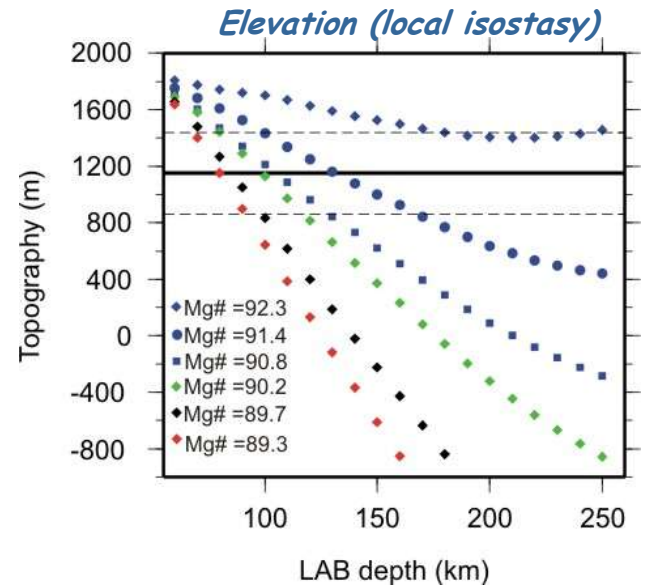
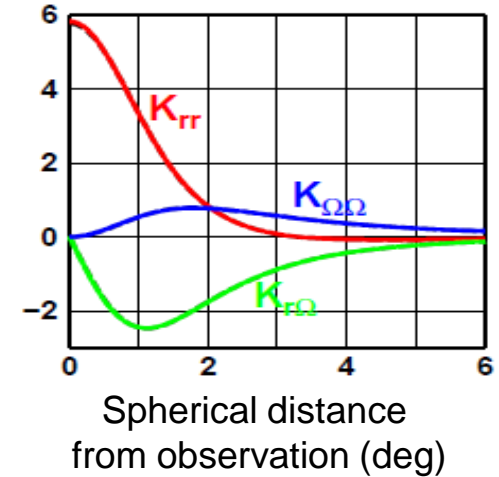
Why integrated modelling...?

Geophysical observables have different sensitivities to T, C variations (via seismic velocities, densities, conductivity...)

Surface waves sensitivity kernels



Gravity gradients kernels @ 255 km



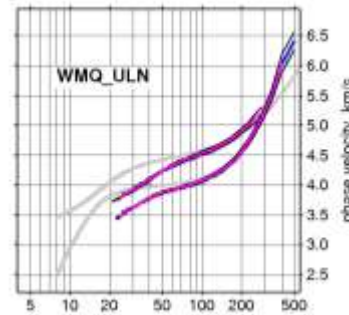
Why integrated modelling...?

Modelling all the observables together **avoids (some) inconsistencies** given the non-uniqueness of the physical problem at hand

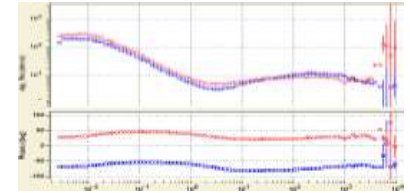
Xenoliths



Surface waves



Magnetotellurics



ρ_1, C_1

V_{s6}, V_{p7}, ρ_8

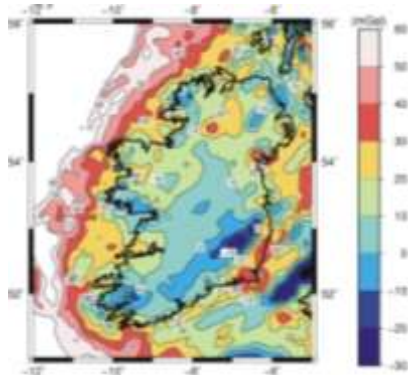
σ_5, C_5

Same T, C???

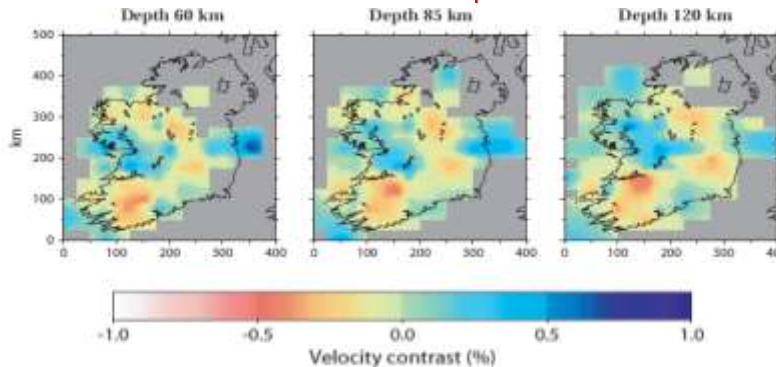
ρ_2

V_{s3}, V_{p3}

T_4



Potential fields



Seismic tomography



Surface heat flow

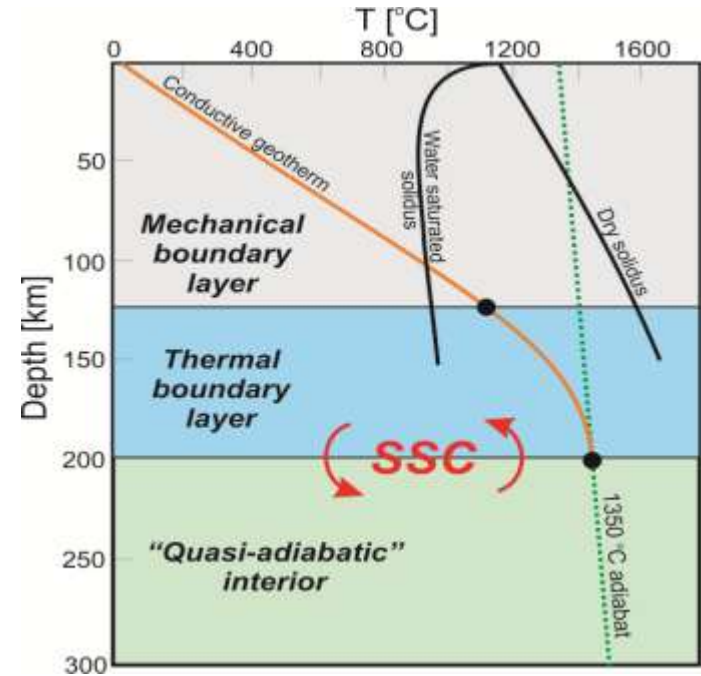
Heat transport equation:

$$\rho c_p \frac{\partial T}{\partial t} = \nabla \cdot (k \nabla T) + H - \rho c_p \vec{u} \cdot \nabla T$$

Lithosphere: conductive mantle

Steady-state conduction equation: $dT/dt=0$ and $U=0 \rightarrow$ Diffusion PDE

$$\nabla \cdot [k(\vec{x}, T, P) \nabla T(\vec{x})] = -H(\vec{x})$$

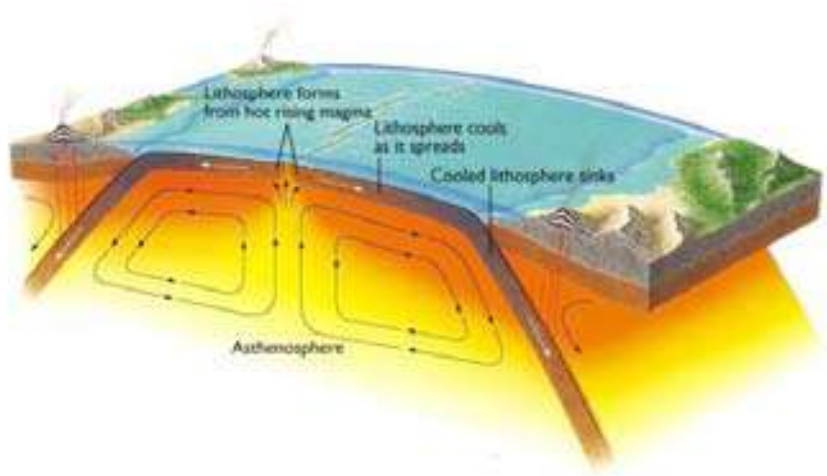


Sub-Lithosphere: mantle convection

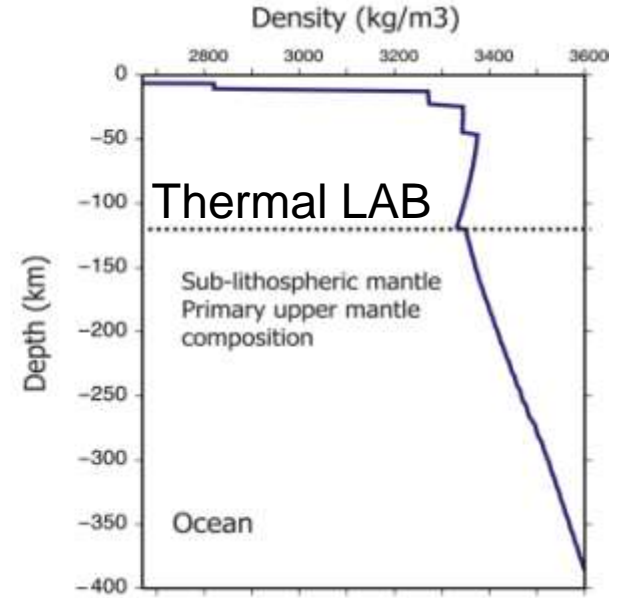
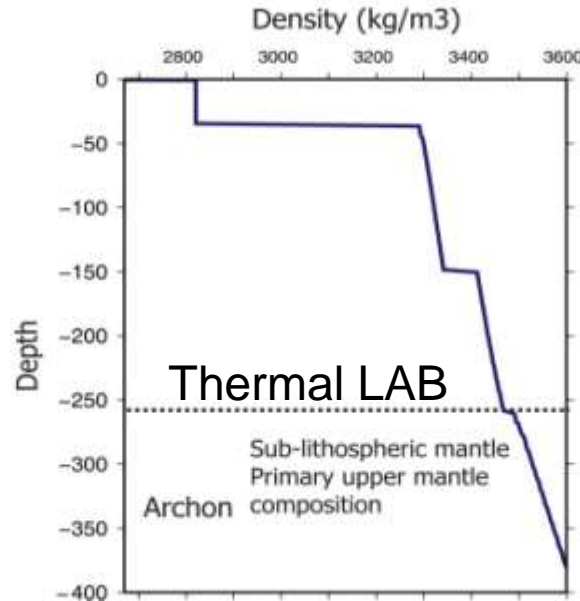
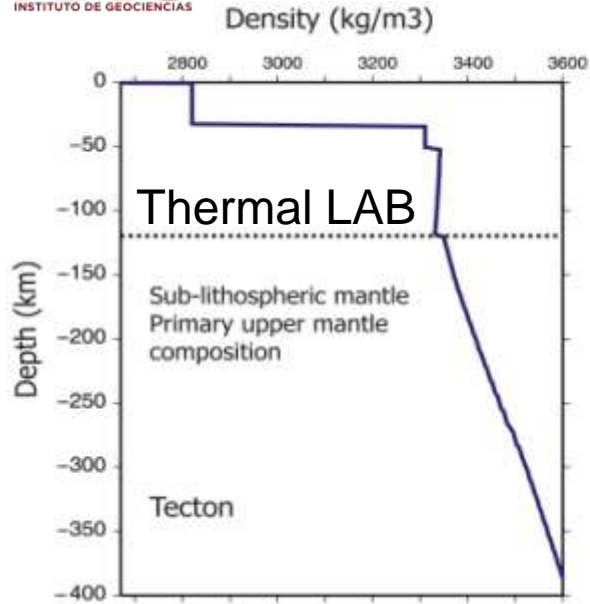
Convection in the mantle (i.e. no heat interexchange with the surroundings). Fast heat transport mechanism compared to conduction

$$\left(\frac{\partial T}{\partial r} \right)_S = \frac{T \alpha g}{c_p}$$

Adiabatic gradient: typically 0.45-0.6 K/km in the uppermost mantle



Ultimately most of the **LAB** definitions depend upon temperature/pressure and **composition**



	Aver. Archon Gnt. SCLM*	Aver. Kaapvaal Harzburg.	Aver. Tecton Gnt. SCLM *	PUM J79 [#]	PUM MS [#] *
SiO ₂	45.7	45.9	44.5	45.2	45
TiO ₂	0.04	0.05	0.14	0.22	0.2
Al ₂ O ₃	0.99	1.3	3.5	4	4.5
Cr ₂ O ₃	0.28	0.34	0.4	0.46	0.38
FeO	6.4	6.0	8.0	7.8	8.1
MnO	0.11	0.1	0.13	0.13	0.14
MgO	45.5	45.5	39.8	38.3	37.8
CaO	0.59	0.5	3.1	3.5	3.6
Na ₂ O	0.07	0.07	0.24	0.33	0.36
NiO	0.3	0.28	0.26	0.27	0.25
Mg#	92.7	93.1	89.9	89.7	89.3
Cr/(Cr+Al)	0.16	0.27	0.07	0.07	0.05

In the mantle, stable mineral assemblages can be computed by Gibbs free energy minimization either within the system CaO-FeO-MgO-Al₂O₃-SiO₂ (CFMAS) or Na₂O-CaO-FeO-MgO-Al₂O₃-SiO₂ (NCFMAS) [Connolly, 2005]. Each mantle body is therefore characterized by a specific **major-element composition** (in wt.%), which translates into specific bulk-rock properties.

Why integrated modelling...?

All thermophysical properties (e.g. **density, seismic velocities, electrical conductivity**) depend on **T, P, and Composition**

$$dG = V dP - S dT + \sum_i \mu_i dX_i + Dd\vec{E}$$

Thermodynamic equilibrium ($T > 500$ C)

$$V = \left(\frac{\partial G}{\partial P} \right)_T$$

$$V\alpha = - \left(\frac{\partial S}{\partial P} \right)_T = \left(\frac{\partial^2 G}{\partial P \partial T} \right)$$

$$C_P = -T \left(\frac{\partial^2 G}{\partial T^2} \right)_P$$

$$c_{ijkl} = \frac{1}{V} \left(\frac{\partial^2 G}{\partial S_{ij} \partial S_{kl}} \right)_{P,T}$$

$$\sigma = \underbrace{\sigma_0 \exp\left(\frac{-\Delta H(X_{Fe}, P)}{k_B T}\right)}_{\text{small polaron}} + \underbrace{f(C_w) \exp\left(\frac{-\Delta H_{wet}(C_w)}{k_B T}\right)}_{\text{Proton conduction}} + \underbrace{\sigma_{0i} \exp\left(\frac{-\Delta H_i}{k_B T}\right)}_{\text{Ionic conduction}}$$

small polaron
(electron hopping
between ferric Fe³⁺
and ferrous Fe²⁺
ions) at $T < 1300$ °C

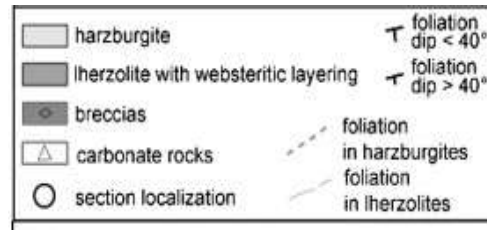
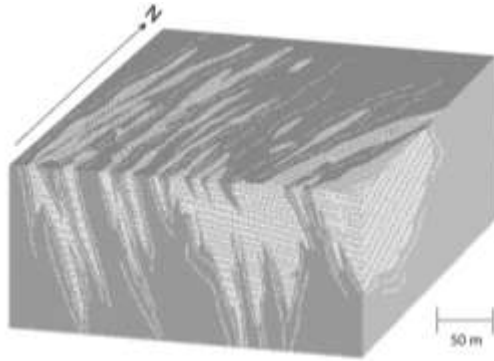
Proton
conduction
(H⁺ bound to
structural O
atoms)

Ionic
conduction
(Mg
vacancies)
 $T > 1300$ °C

→ connect laboratory studies & thermodynamics with geophysics

Why integrated modelling...?

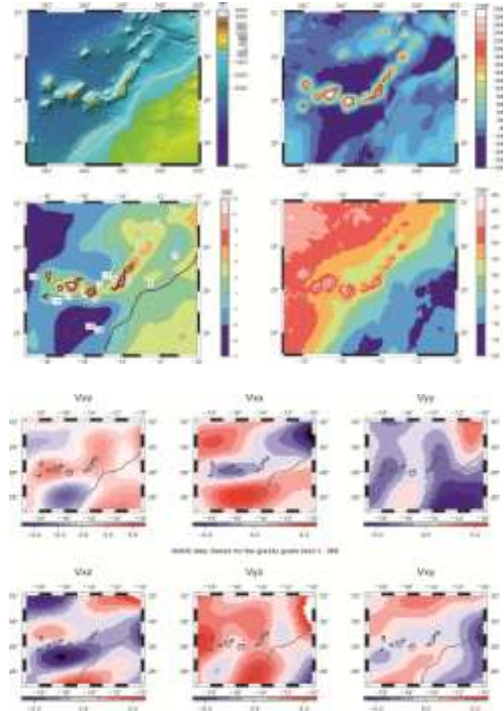
Representativeness of observed mantle samples (xenoliths, peridotite massifs etc..) on the lithospheric scale



Forward modeling



Geophysical data



Depleted

Depleted +metasomatised

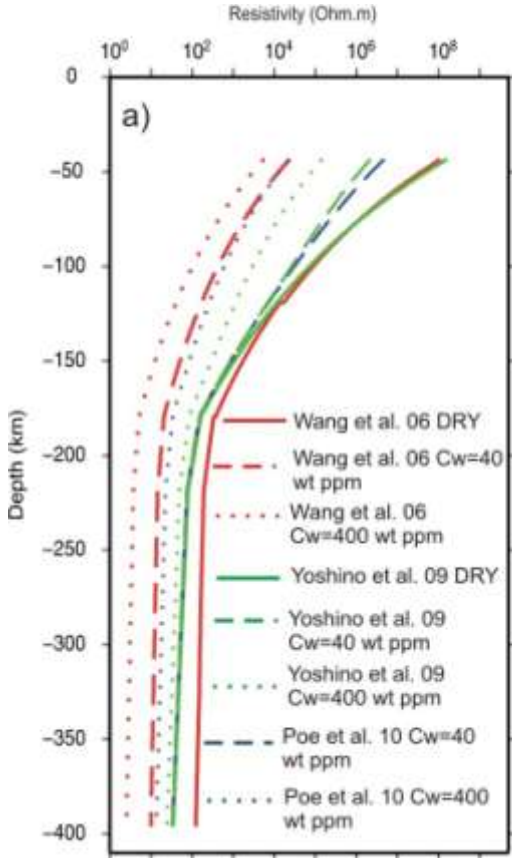
	1) Av. Harz. Lanzarote (wt%) ^a	2) Av. HEXO Tenerife (wt%) ^b	3) Av. Harz. La Palma av. (wt%) ^c	4) Av. HLCO Tenerife (wt%) ^b	5) Av. Ocean floor peridot. (wt%) ^d	6) Av. Middle Atlas (wt%) ^e
SiO ₂	43.78	43.32	43.07	42.14	45.09	43.48
Al ₂ O ₃	0.7	0.61	0.53	0.73	2.33	2.38
FeO	7.79	8.04	8.43	8.8	8.4	8
MgO	46.1	45.31	45.19	44.14	41.23	42.6
CaO	0.6	0.81	0.68	1.68	1.32	2.83
Na ₂ O	0.1	0.14	0.17	0.18	0.23	0.24
Mg#	91.34	90.96	90.53	89.94	89.7	90.47

Mantle Depletion (partial melting)

Mantle metasomatism (refertilization)

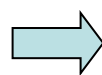
Why integrated modelling...?

olivine



For olivine there are two models for proton conduction based on lab studies:

$$\sigma_p = \sigma_{0p} C_w \exp\left(\frac{-\left(\Delta H_0 - \alpha C_w^{1/3}\right)}{k_B T}\right)$$



e.g. Yoshino et al., 09;
Manthilake et al., 09;
Poe et al., 10

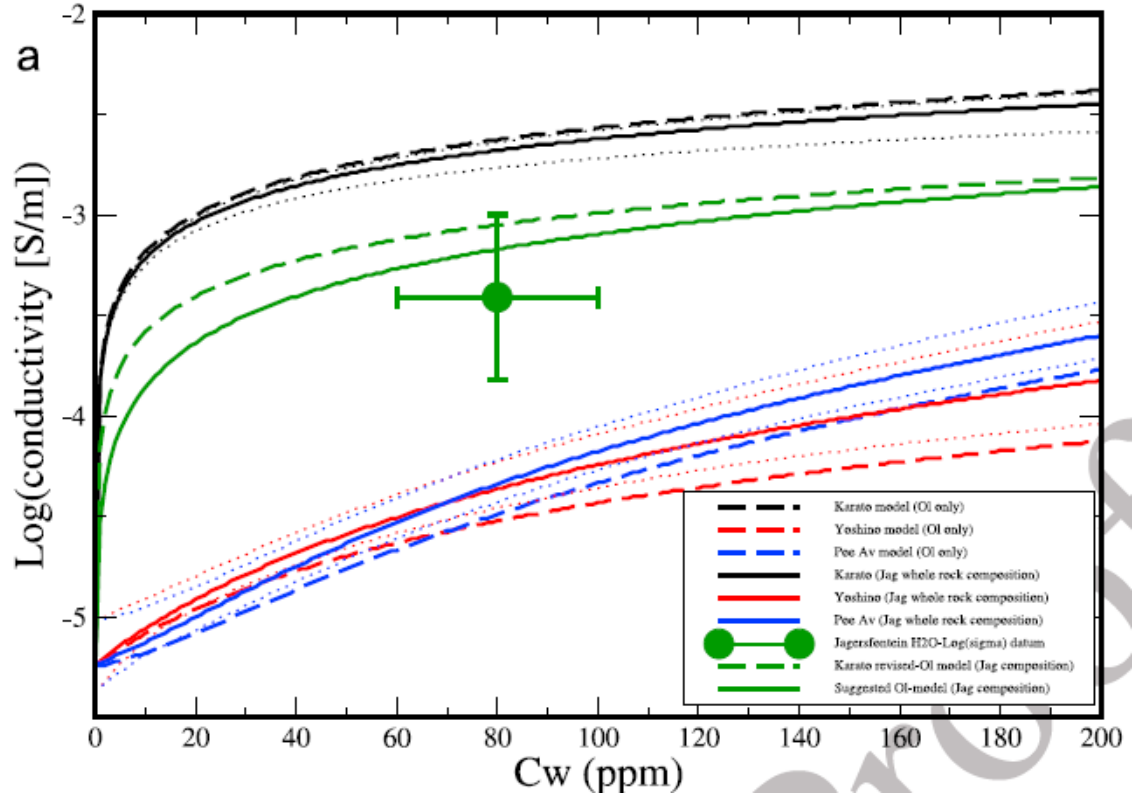
$$\sigma_p = A C_w^r \exp\left(\frac{-\Delta H_{wet}}{k_B T}\right)$$



e.g. Wang et al., 06; Dai and Karato 09

Jagersfontein: Comparison of H2O models

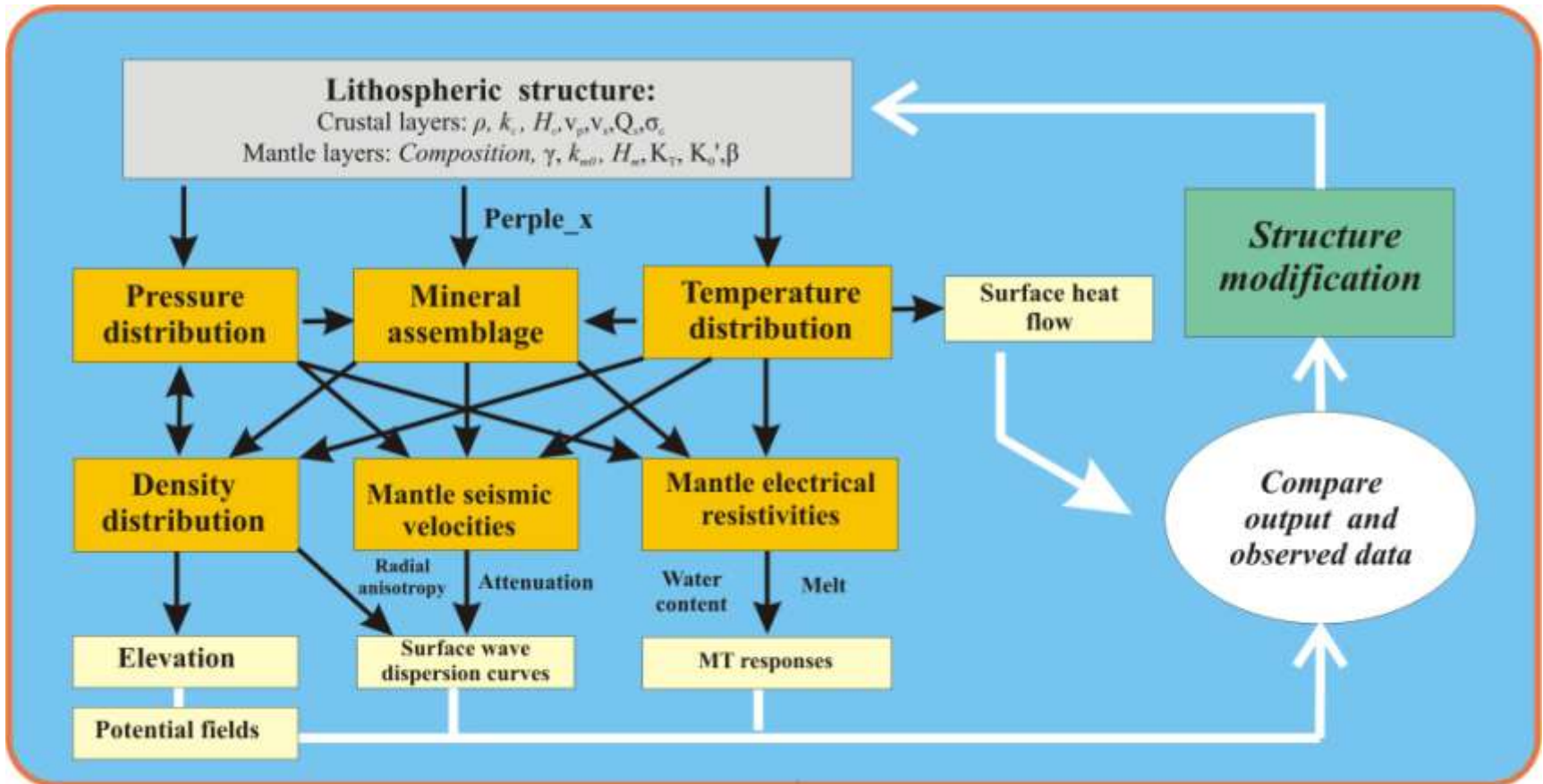
T=740 C, P=3.2 Gpa, Mg#=93.2, Ol=69.5%, Opx=24.17%, Cpx=3.4%, Gt=3.4%, Sp=0.32%



→ Calibrate lab studies with geophysics

Forward modelling: trial and error

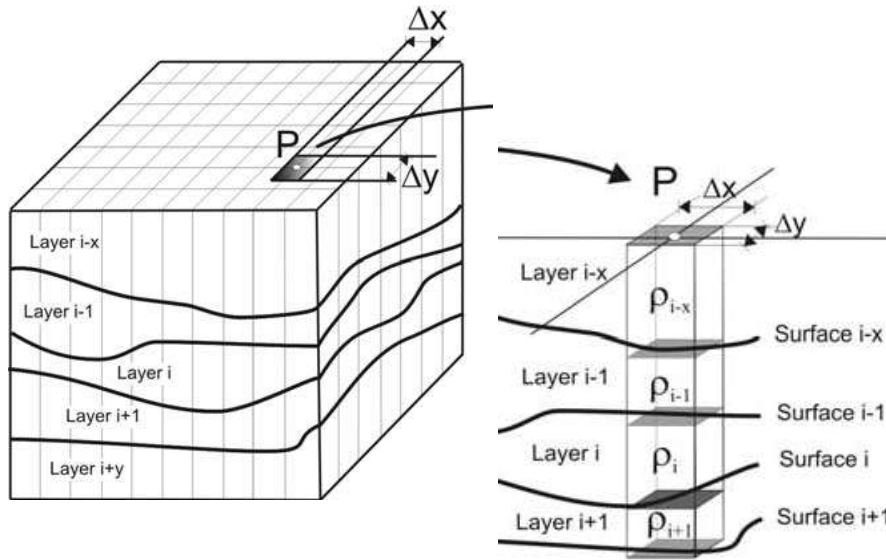
Lithospheric model → Observables (e.g., topography, potential fields, seismic data, MT...)



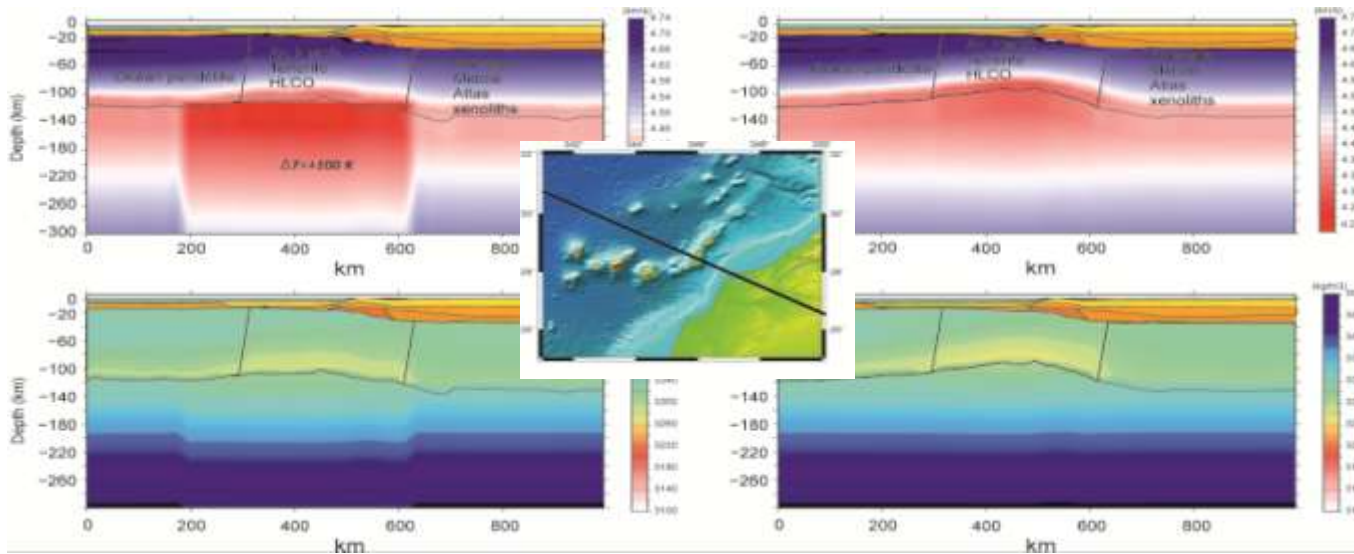
All necessary files containing thermodynamic information can be generated with the freely available software **Perple_X** [www.perplex.ethz.ch, Connolly, 2005].

Forward codes **LitMod** (1D,2D,3D) for integrated modelling available at [<http://eps.mq.edu.au/~jafonso/Software1.htm> , Afonso et al., 2008; Fullea et al., 2009].

Forward modelling: advantages



*High spatial resolution (fine grids) depending on computational power available



*Exploratory nature, hypothesis testing

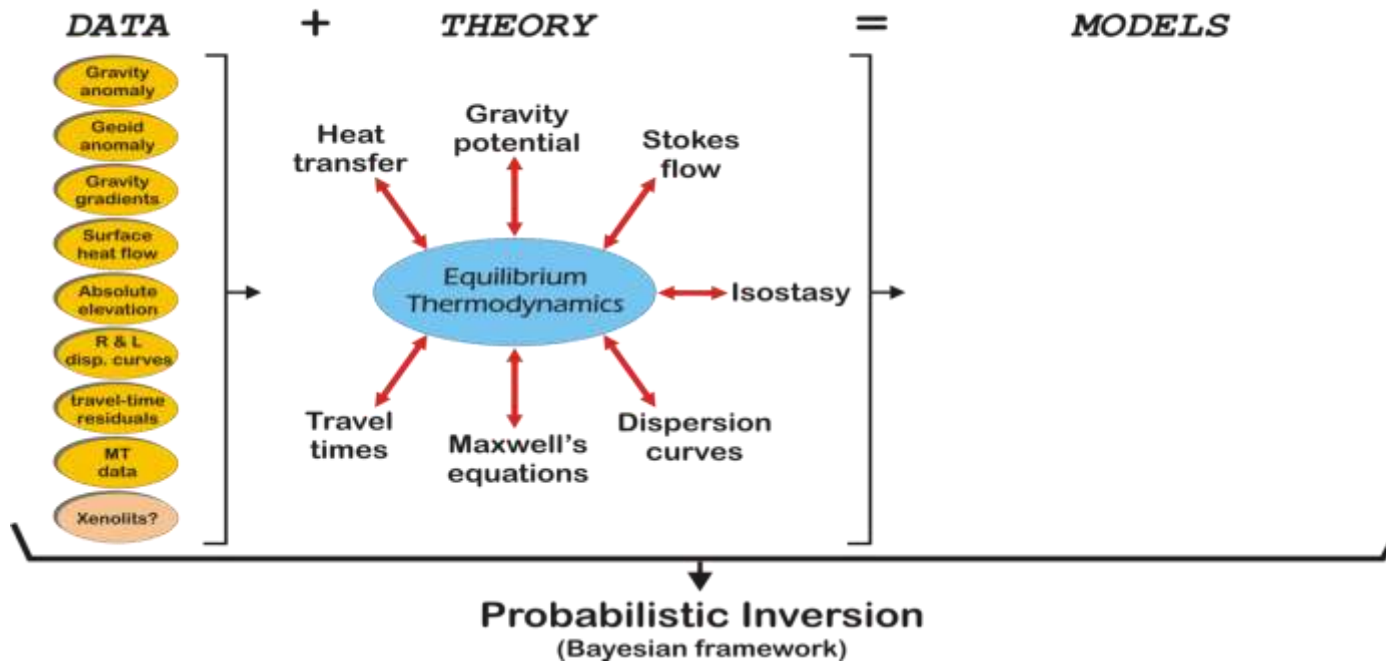
Forward modelling: disadvantages

- *Requires a priori/complementary info
- *Huge (potential) parameter space, time demanding task
- *Possible bias from the modeller's prejudices/background



Inversion "version" of LitMod??

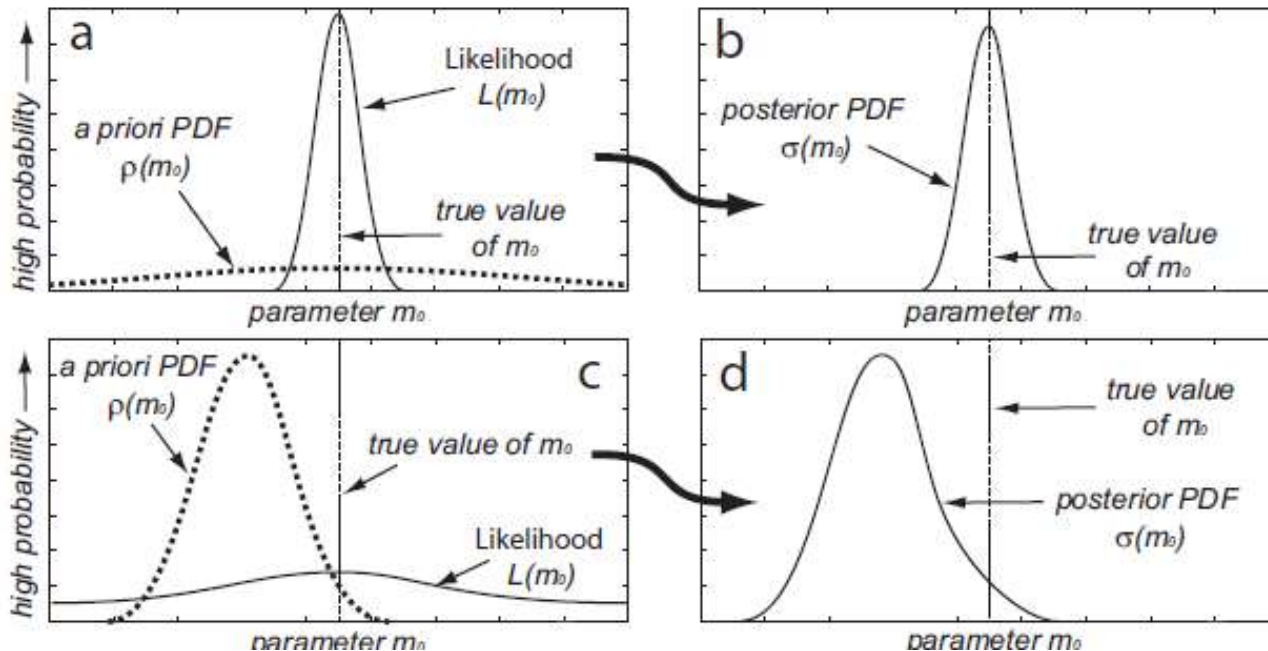
LitMod_4INV



Probabilistic non-linear inversion

- * Looking for T and C (5 oxides+water) at every model node
- * Typical forward model: 50x50x200 (5E5) nodes \rightarrow 5E5*7 param/node=3.5 millions of parameters!!
- * Systematic brute-force inversion scheme not affordable
- * The physical problem is highly non-linear
- * A good control on prior PDF $\rho(\mathbf{m})$ (a priori info) and likelihood $L(\mathbf{m})$ (observational and theoretical uncertainties, covariance matrix) is essential.

$$\sigma(\mathbf{m}) = k\rho(\mathbf{m})L(\mathbf{m}) \quad L(\mathbf{m}) \propto \exp \left\{ -\frac{1}{2} [\mathbf{g}(\mathbf{m}) - \mathbf{d}_{\text{obs}}]^T C_D^{-1} [\mathbf{g}(\mathbf{m}) - \mathbf{d}_{\text{obs}}] \right\}$$

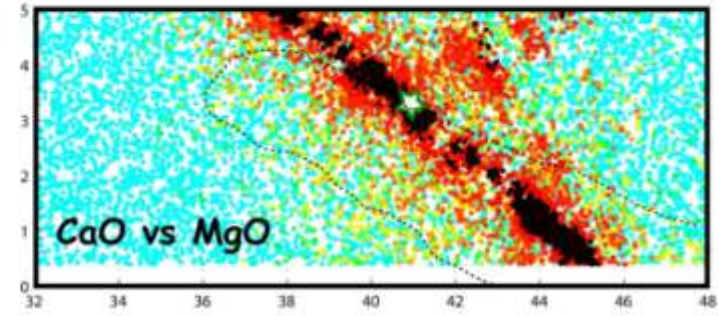
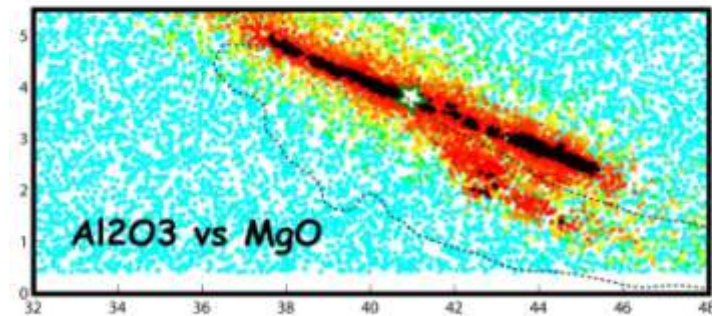
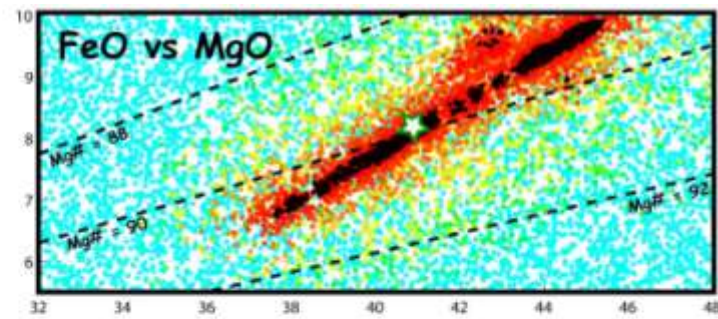
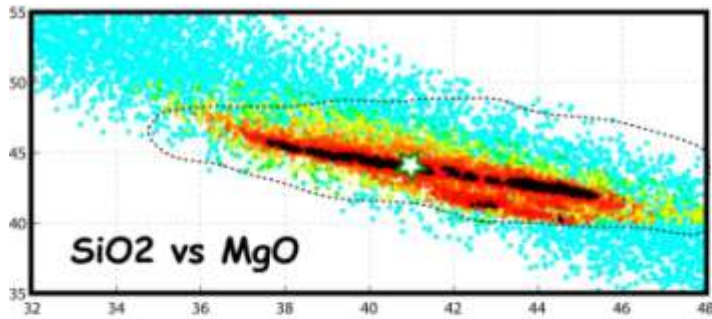


Probabilistic non-linear inversion

- * Trade-offs between T and C
- * T has a greater effect than C in most of the observables
- * Non uniqueness of compositional field (worse in the lithosphere than in the sublithosphere)

$$\sigma(\mathbf{m}) = k\rho(\mathbf{m})L(\mathbf{m})$$

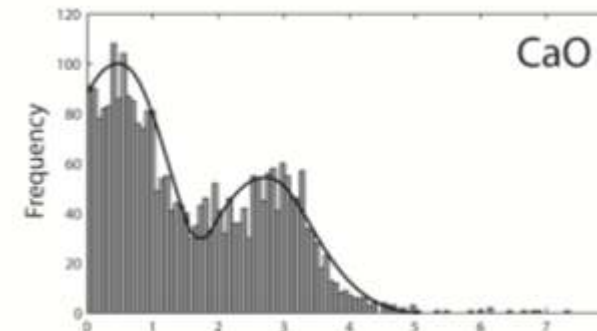
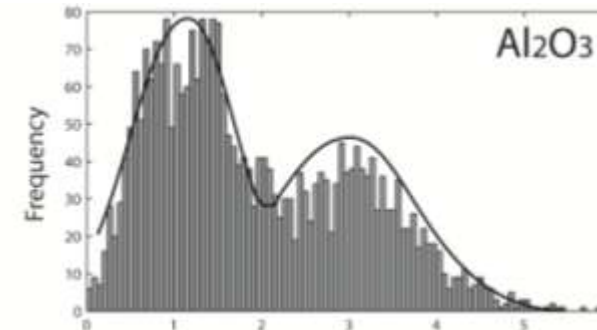
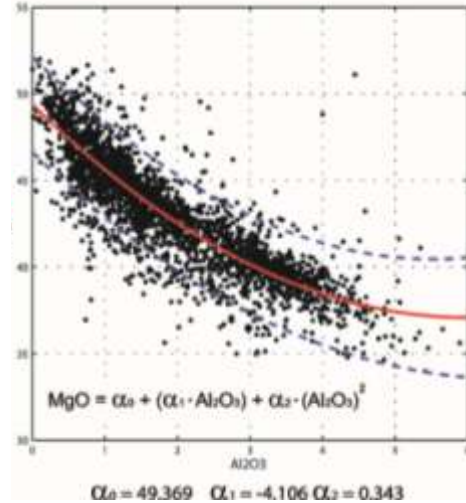
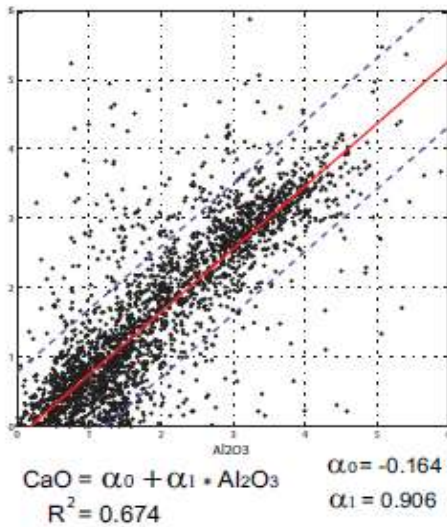
$$L(\mathbf{m}) \propto \exp \left\{ -\frac{1}{2} [\mathbf{g}(\mathbf{m}) - \mathbf{d}_{\text{obs}}]^T C_D^{-1} [\mathbf{g}(\mathbf{m}) - \mathbf{d}_{\text{obs}}] \right\}$$



combined residual V_p, V_s, ρ

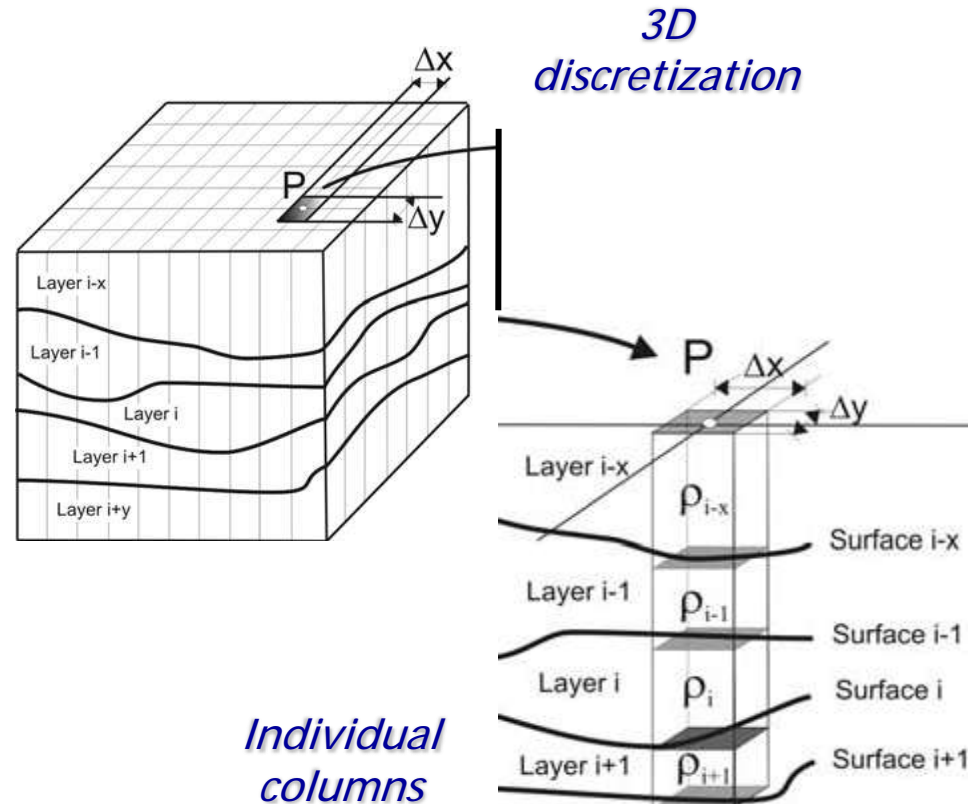
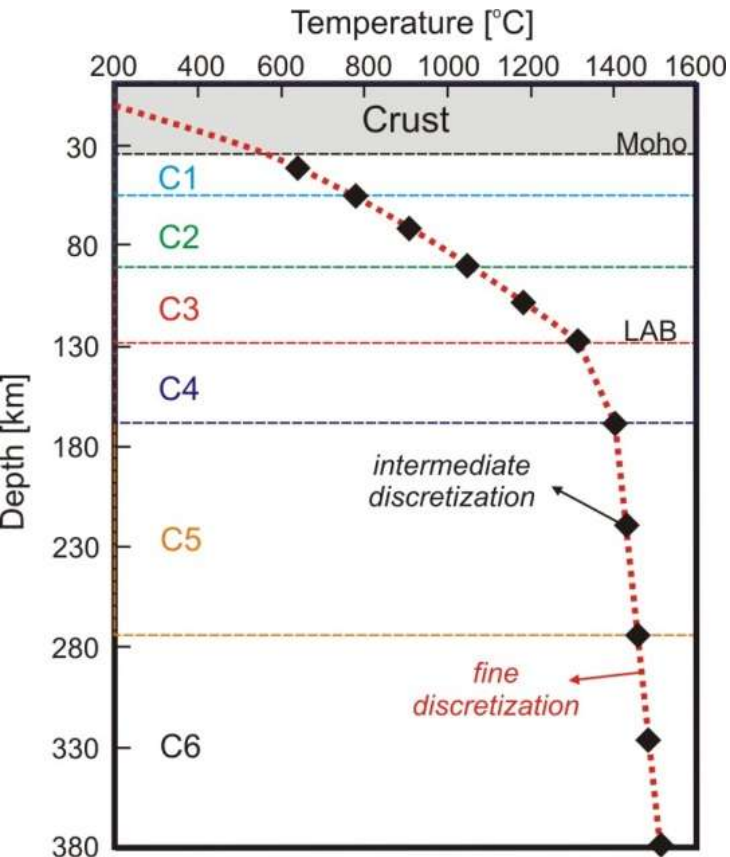
Defining prior PDF $\rho(m)$

- * A priori petrological data base (>2900 samples from xenoliths, perid. Massifs and ophiolites)
- * Correlation between oxides (Al₂O₃ and FeO as independent C param.), regardless of tectonic age or facies.
- * Possible bias in database (e.g. double peaks) due to
- * Wide Al₂O₃ and FeO ranges (>95% of natural variability) with uniform probability density



Reducing the parameter space

- * Lower resolution: 6 compositional layers and 12 “thermodynamic nodes” (vs 200 vertical nodes in fwd. models)
- * Split the 3D problem into 1D columns (for 1D data) → first order PDF’s used as priors in the full 3D inversion
- * PDF’s sampled via MCMC simulation using Metropolis-Hastings algorithm



Individual columns

Method I

First Part
(search & appraisal in 1D)

Input

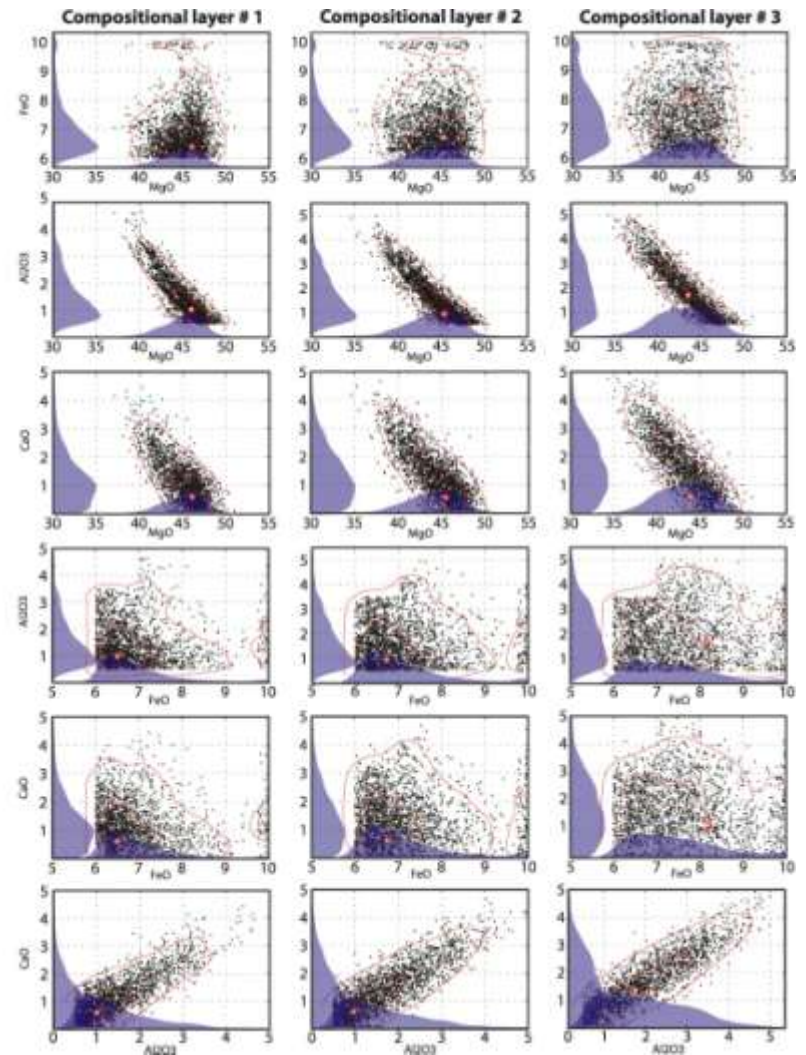
- * Dispersion curves
- * 1D geoid anomaly
- * xenolith data
- * Vp structure
- * elevation
- * MT data
- * surface heat flow

Nonlinear inversion

(Search for acceptable models
"column by column" and their
PDFs using NA)

Output

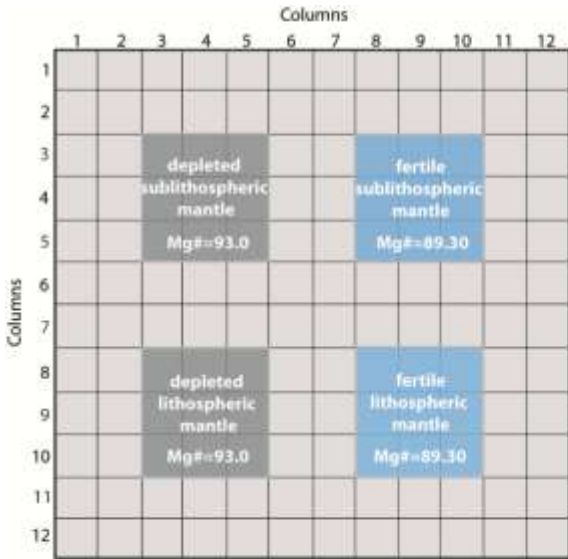
(PDFs for all parameters
in all individual columns)



1D Search focused in T: average lithospheric/sub-lithospheric C, $dT > 200$ K
 1D search focused on C: 6 compositional layers using LAB posterior PDF

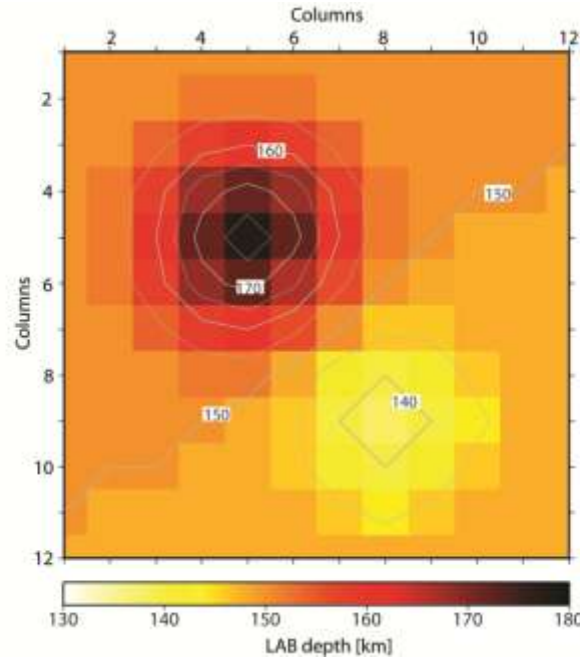
Synthetic test

*4 compositional domains + laterally varying LAB and Moho → its forward responses+noise serve as input for the inversion

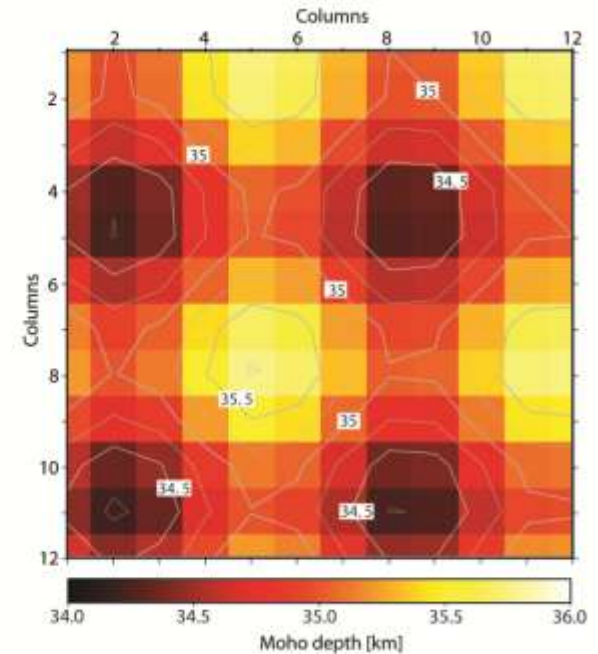


Background Mg#=89.9

Domains



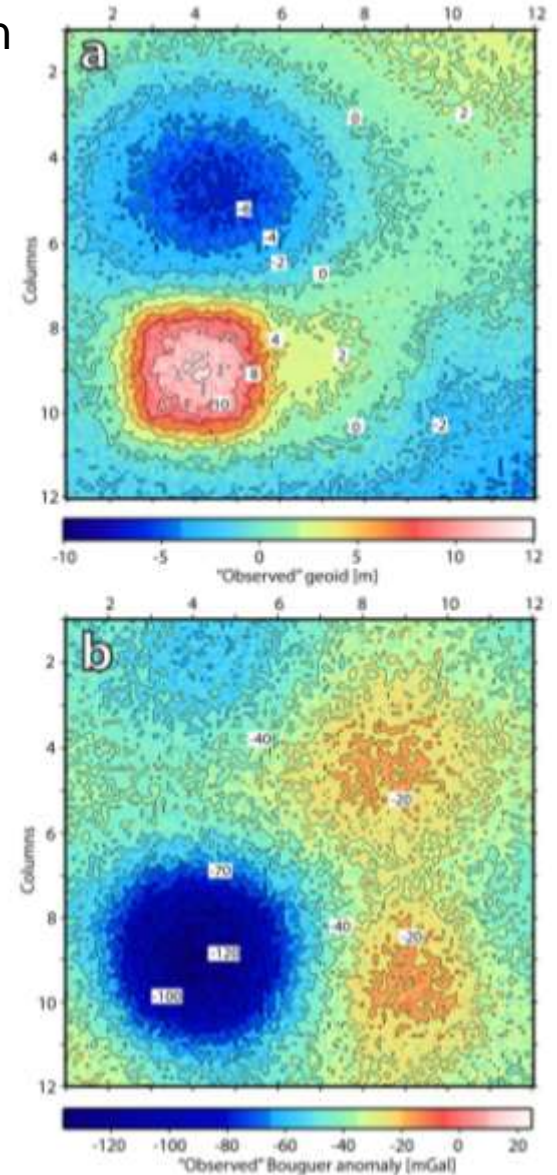
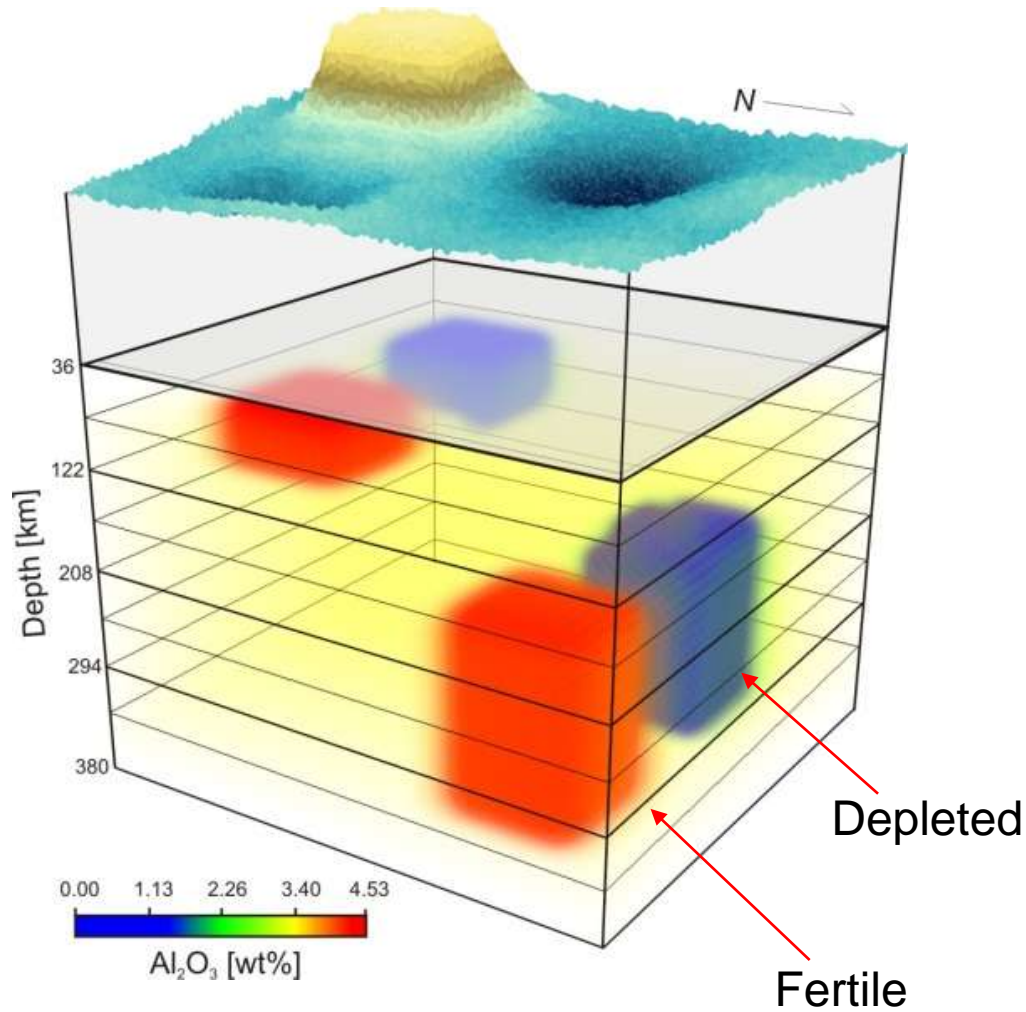
LAB



Moho

Synthetic test

Inputs (synthetic forward+noise) used in the inversion

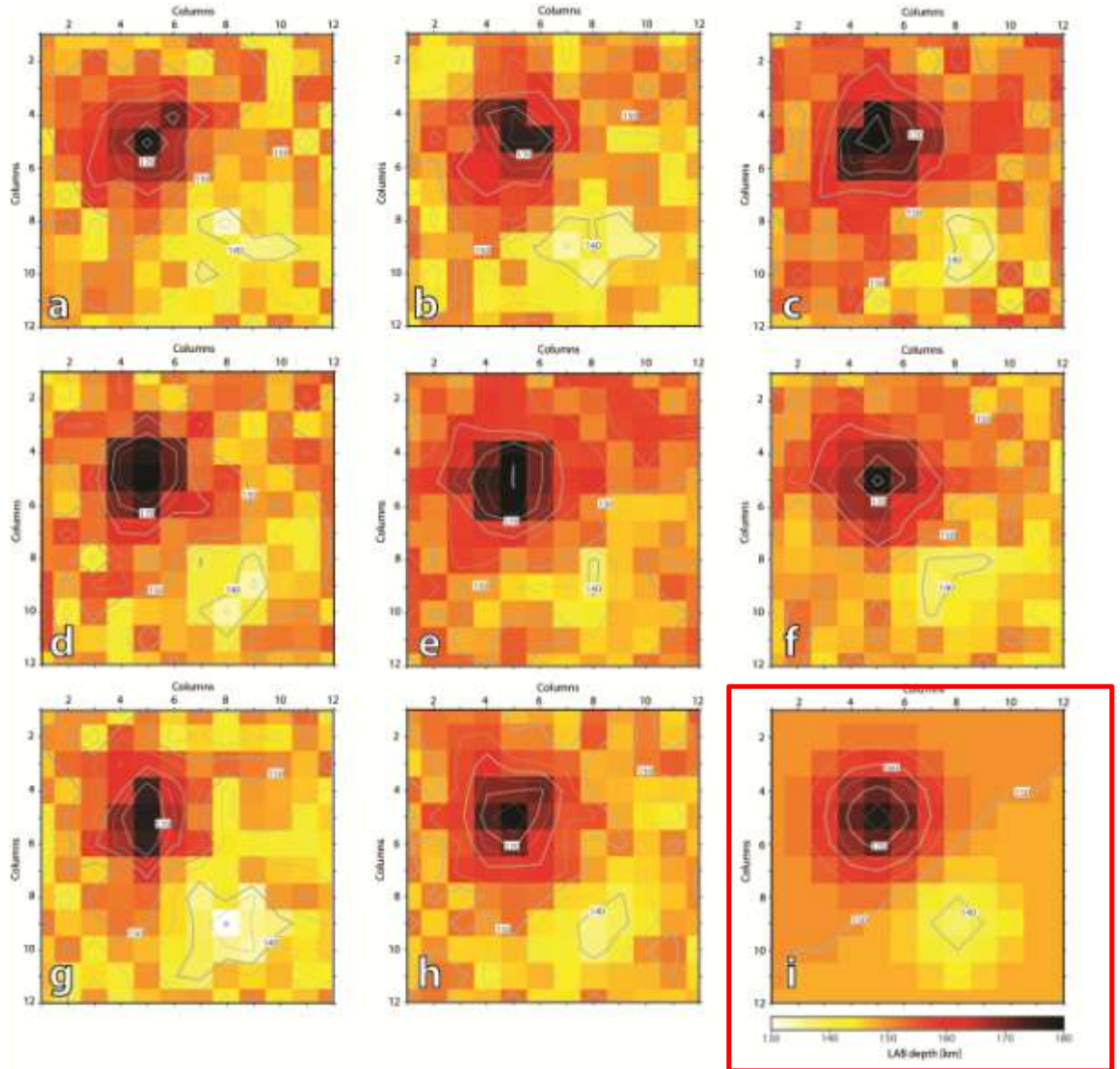


Synthetic test

Results for LAB geometry

Mean models of 8 random ensembles with 500 samples each taken from the total posterior.

Regular and smooth posterior PDF → averaging is meaningful

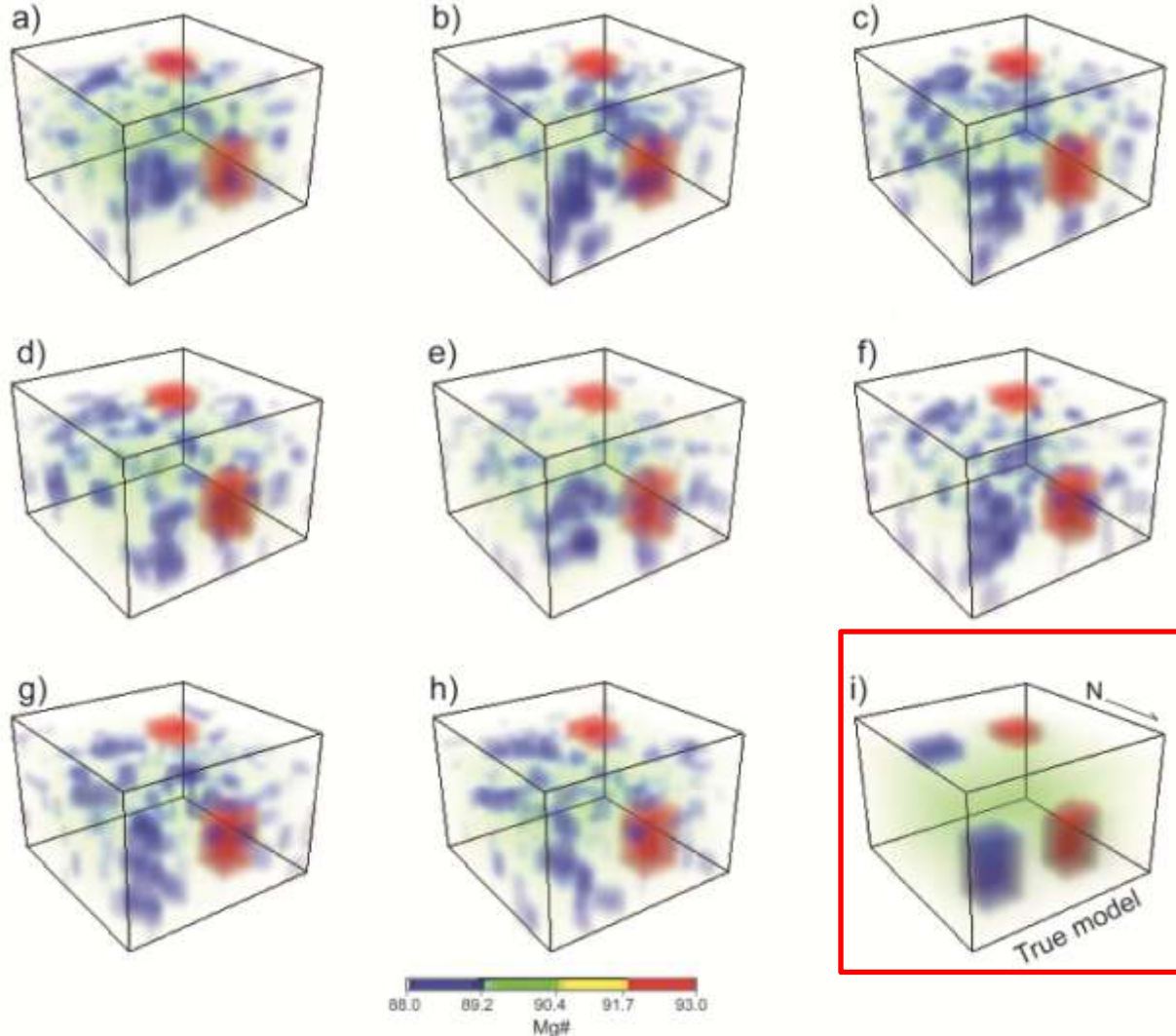


Synthetic test

Results for Mg#

Mean models of 8 random ensembles with 500 samples each taken from the total posterior

Depleted domains (high Mg#) are recovered, fertile ones (low Mg#) are blurred

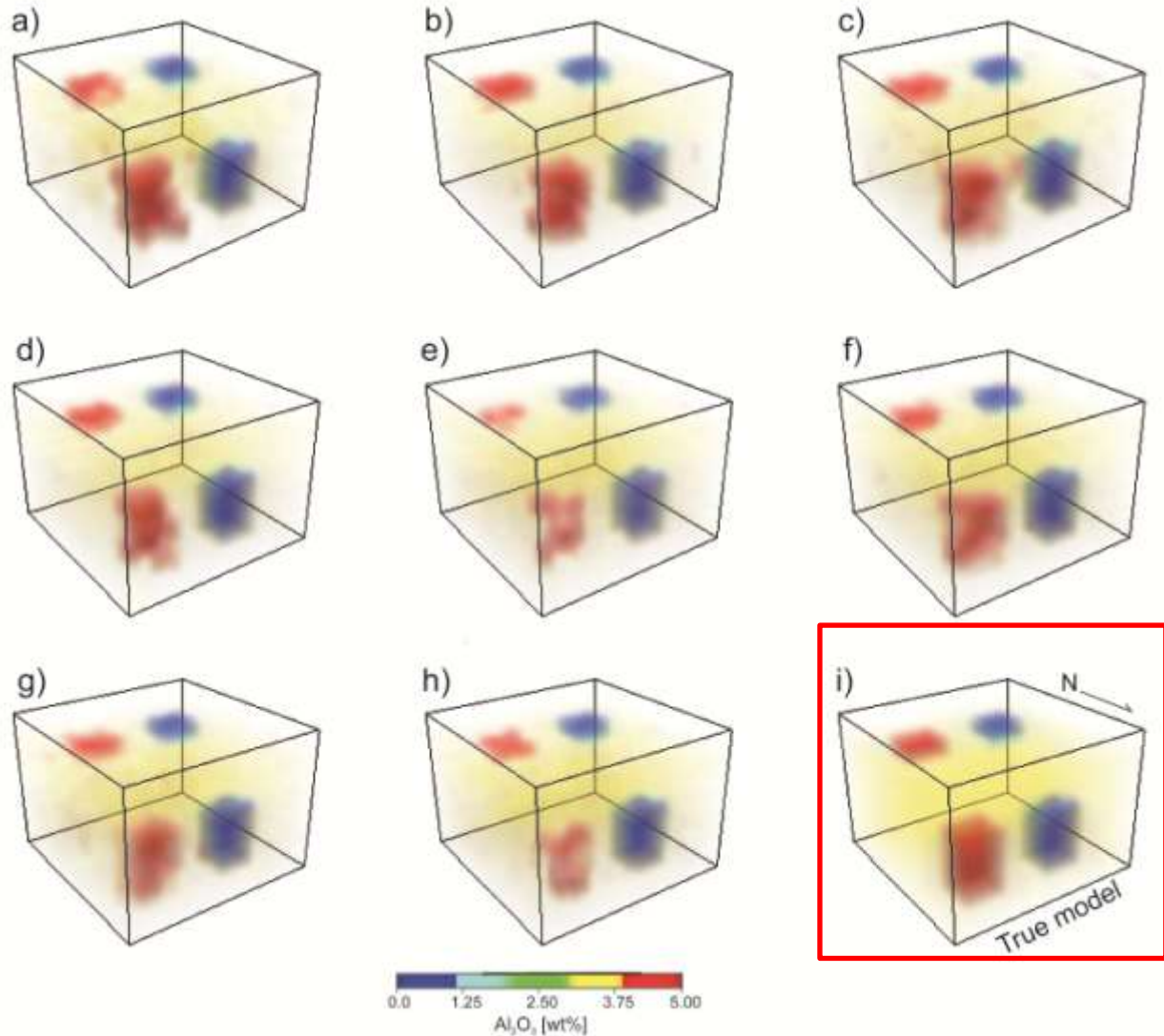


Synthetic test

Results for bulk Al_2O_3

Mean models of 8 random ensembles with 500 samples each taken from the total posterior

Fertile and depleted domains are recovered $\rightarrow Al_2O_3$ is a sensible indicator for compositional end-members



Internally consistent combination of geophysical observables with different sensitivities to T and C

Integrated modelling: forward vs (and) inversion approaches

**Forward model:*

- Requires a priori/complementary info
- Huge (potential) parameter space, time demanding task
- Possible bias from the modeller's prejudices/background
- + Relatively fast (seconds-minutes)
- + High resolution affordable

**Probabilistic Inversion:*

- + It does not require a priori/complementary info
- + Parameter space (T, C) effectively explored in hours-days
- + No bias from the modeller's prejudices/background
- Relatively low resolution

The best of both worlds...



(1) Inversion and (2) detailed forward modelling based on (1) for hypothesis testing

Thank you for your attention!

