

Oceanic and continental plates as seen by seismic anisotropy

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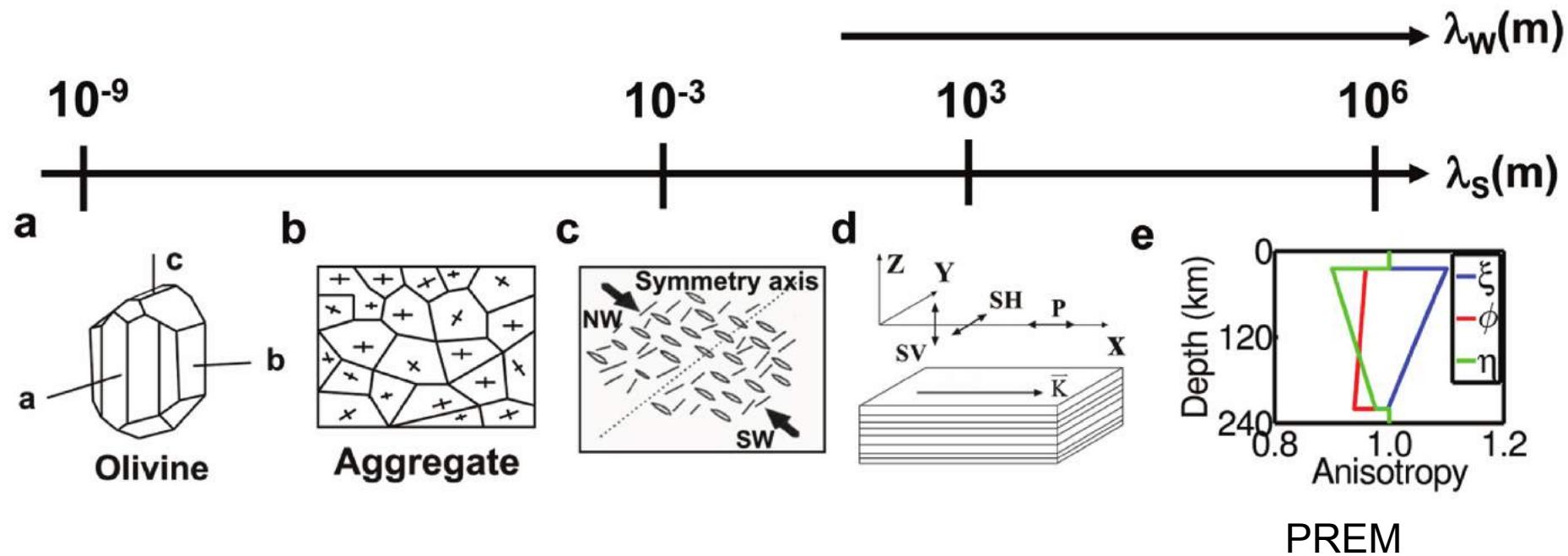
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2- Institut Universitaire de France

OUTLINE

- What is a plate? Role of continents?
Surface manifestation of mantle convection,
role of mantle plumes, slabs?
- Seismic Anisotropy: many processes, different
interpretations
- 3D- anisotropic structure of the Earth
Lithosphere- Asthenosphere Boundary (LAB)
Oceans (Burgos et al., 2014)
Continents: Indian Continent (Maurya et al., 2016)

Seismic Anisotropy at all scales



PREM: radial anisotropy: up to 10%

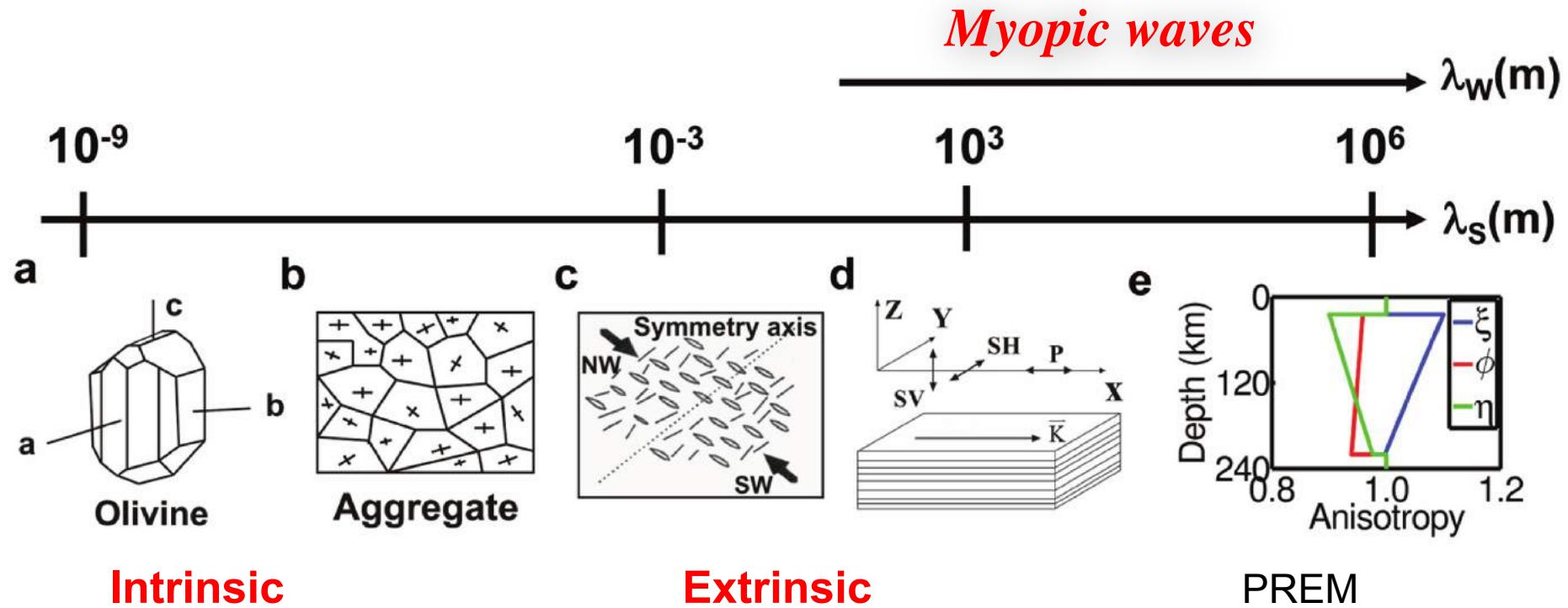
λ_w seismic wavelength

λ_s spatial scale

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(Wang et al., 2013)

Seismic Anisotropy at all scales

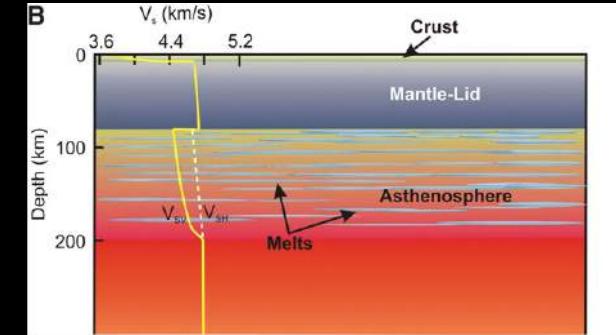
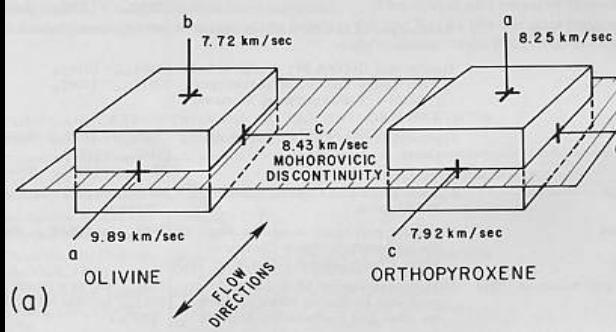
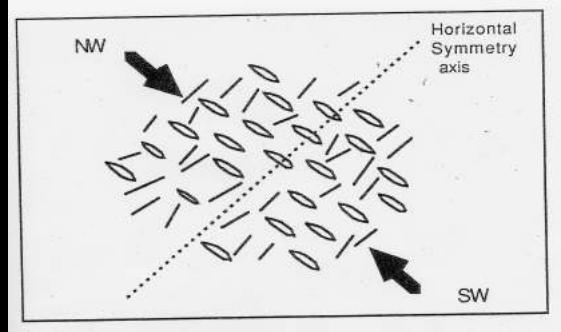


Observed (apparent) anisotropy
Intrinsic versus Extrinsic anisotropy

$$\alpha = p\alpha^{\text{int}} + (1-p)\alpha^{\text{ext}}$$

Different processes in different layers

-S.P.O. (stress) -L.P.O.(strain) Fine Layering



- **Mineralogy, Water and fluid content, layering**
- **Present day tectonics, geodynamic processes**
- **Mantle Convection**
- **Past processes (frozen anisotropy)**

**Separation of the different kinds of anisotropy in different layers => Different interpretations
(Stratification of anisotropy in the crust & mantle)**

Effect of anisotropy on the phase of surface waves

Effect on eigenfrequency ω_k for multiplet $k=\{n,l,m\}$ (Rayleigh's principle)

$$\frac{\delta\omega_k}{\omega_k} = \frac{\int_{\Omega} \varepsilon_{ij}^* \delta C_{ijkl} \varepsilon_{kl} d\Omega}{\int_{\Omega} \rho_0 u_r^* u_r d\Omega} = \frac{\delta V}{V} \Big|_k$$

ε strain tensor, u displacement, δC_{ijkl} elastic tensor perturbation,

V phase velocity (V_R Rayleigh; V_L Love)

Phase velocity perturbation $\delta V(T, \theta, \phi, \Psi)$ at point r (θ, ϕ)

(Smith & Dahlen, 1973; Montagner & Nataf, 1986)

Ψ Azimuth (angle between North and wave vector)

Azimuthal Terms: **0- ψ , 2- ψ , 4- ψ**

$$\delta V(T, \theta, \phi, \Psi) / V = \alpha_0(T, \theta, \phi) + \alpha_1(T, \theta, \phi) \cos 2\Psi + \alpha_2(T, \theta, \phi) \sin 2\Psi + \alpha_3(T, \theta, \phi) \cos 4\Psi + \alpha_4(T, \theta, \phi) \sin 4\Psi$$

- *Cijkl* 21 elastic moduli
- $\alpha_0 = 0$ - ψ term: 5 parameters A, C, F, L, N (PREM)

VTI Model (*transverse isotropy with vertical symmetry axis*)

- Best resolved parameters from surface waves (among 13 parameters when including azimuthal anisotropy 2ψ -, 4ψ - terms)

$$\textcolor{red}{L} = \rho V_{SV}^2 \quad \text{Isotropic part of } V_{SV} \text{ or } \textcolor{red}{\mu_{iso}} \approx (2L+N)/3$$

$$\xi = N/L = (V_{SH}/V_{SV})^2 \quad \text{Radial Anisotropy}$$

$$\textcolor{red}{G, \Psi_G} \quad \text{Azimuthal Anisotropy of } V_{SV}, \text{ also related to SKS splitting}$$

Geodynamic Interpretation: LPO

Convective cell: anisotropic parameters

Tomographies of:

-S- Velocity

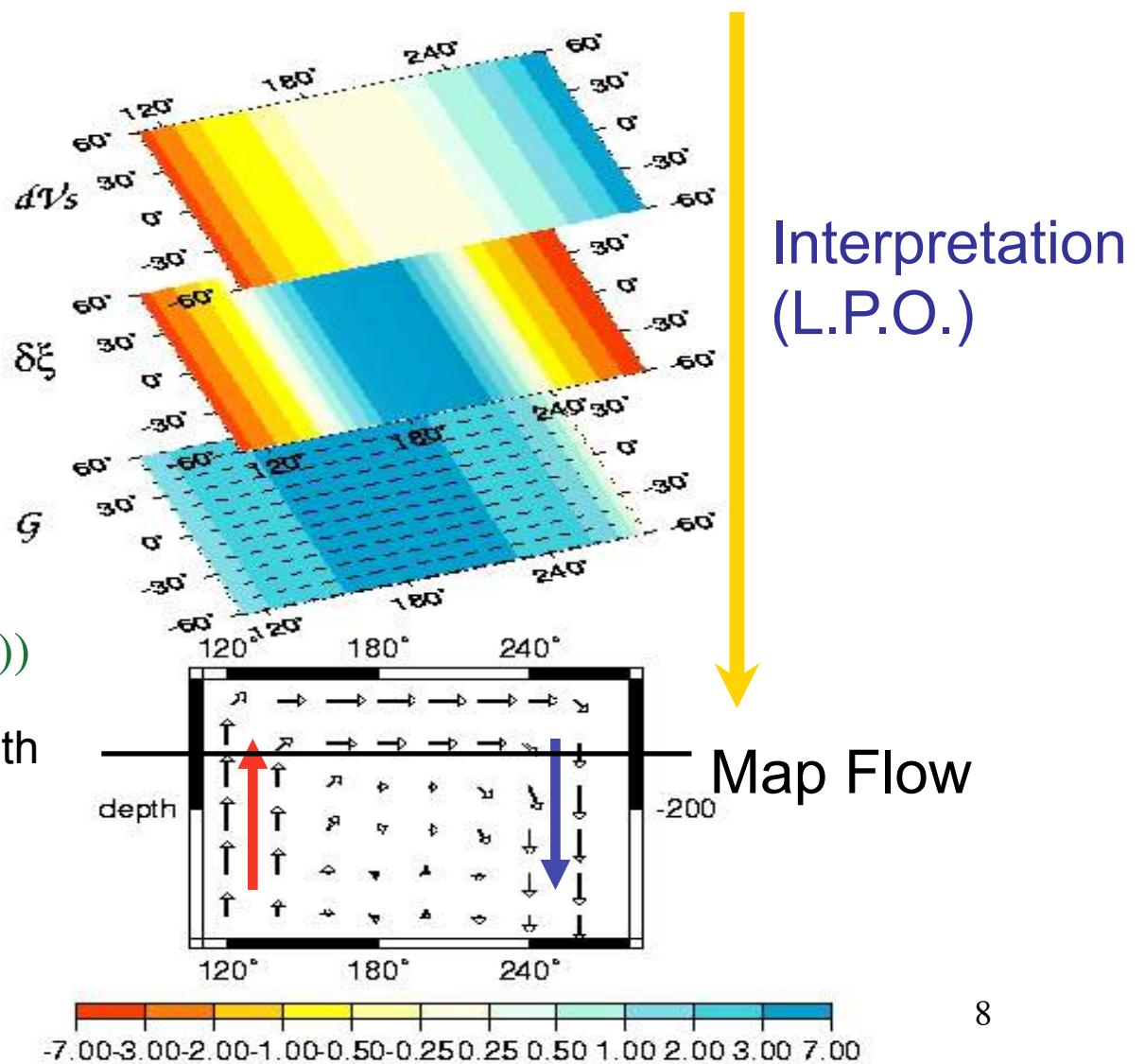
-Radial Anisotropy

$$\delta\xi = (V_{SH}^2 - V_{SV}^2) / V_{SV}^2$$

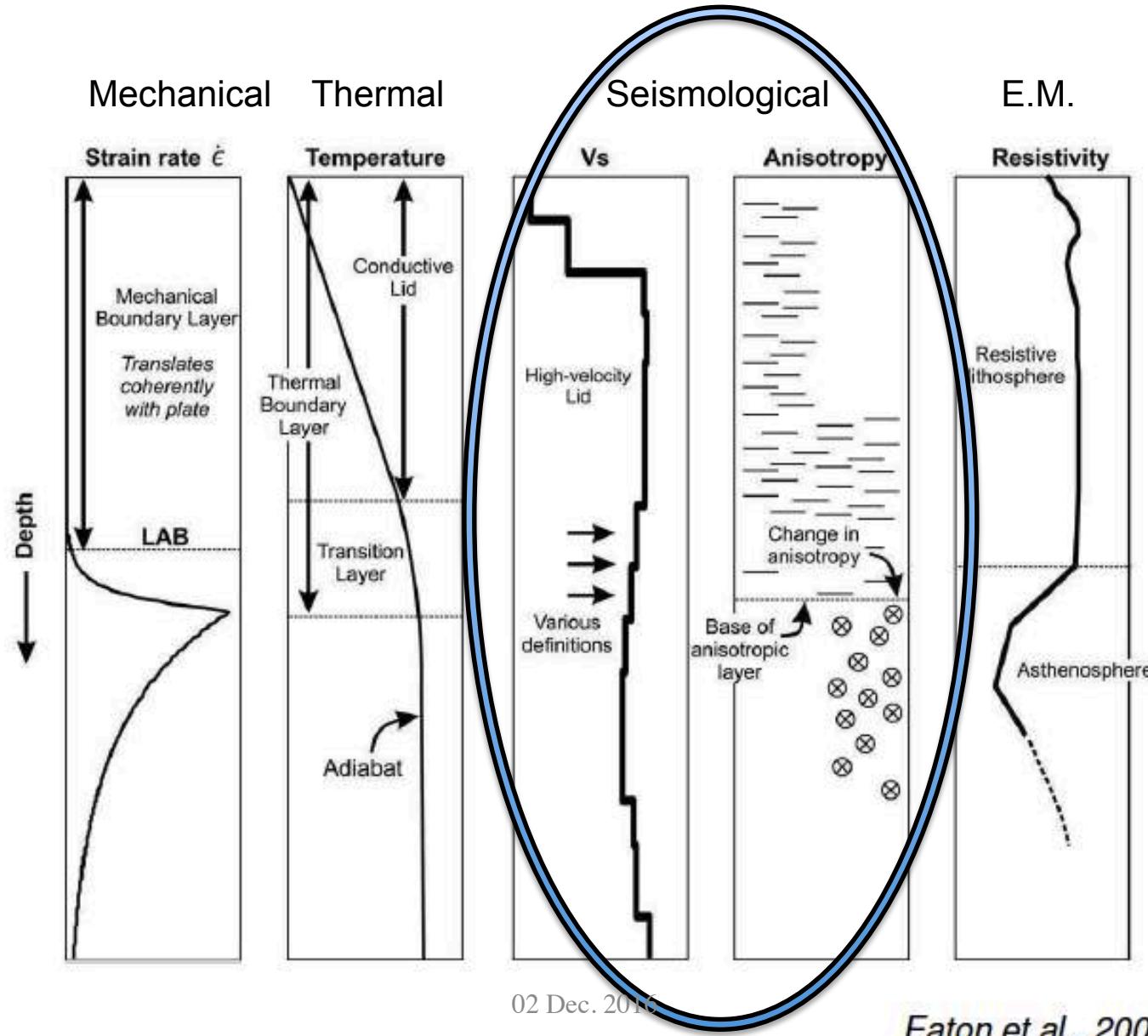
-Azimuthal Anisotropy

$$V_{SV} \approx V_{SV0} + \frac{1}{2} G \cos(2(\Psi - \Psi_G))$$

At a given depth

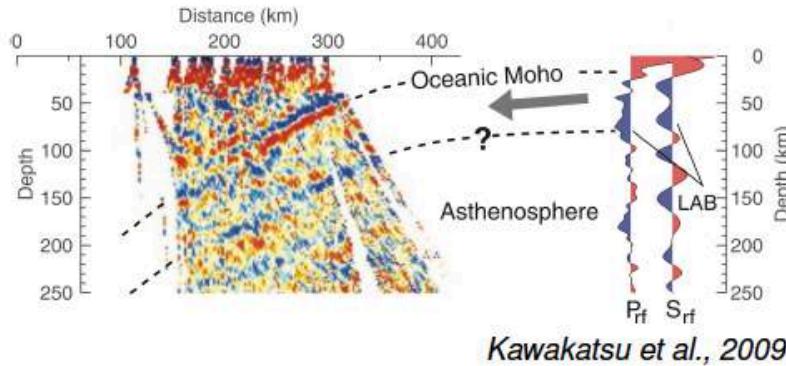


L.A.B.: Lithosphere-Asthenosphere Boundary (many different approaches and definitions)

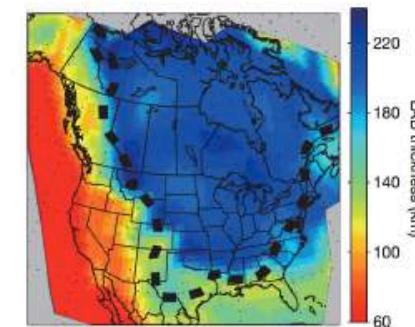
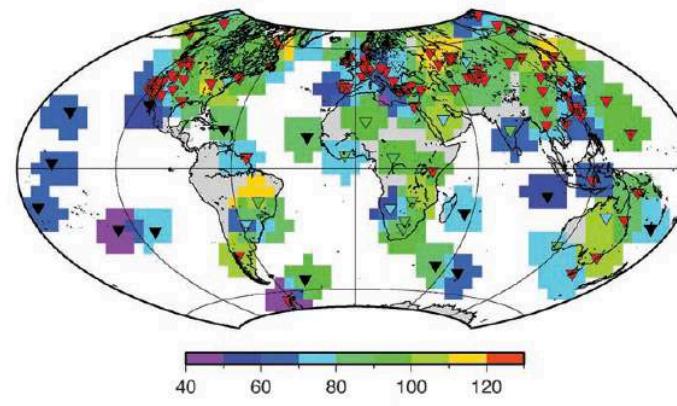
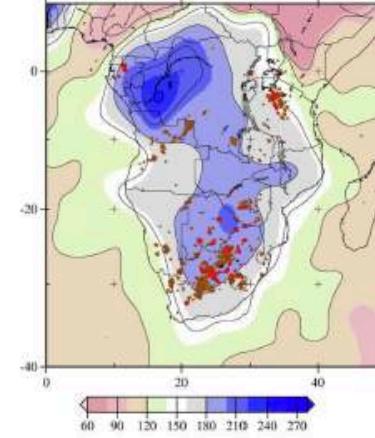


LAB : from seismic data

Receiver functions



Surface waves

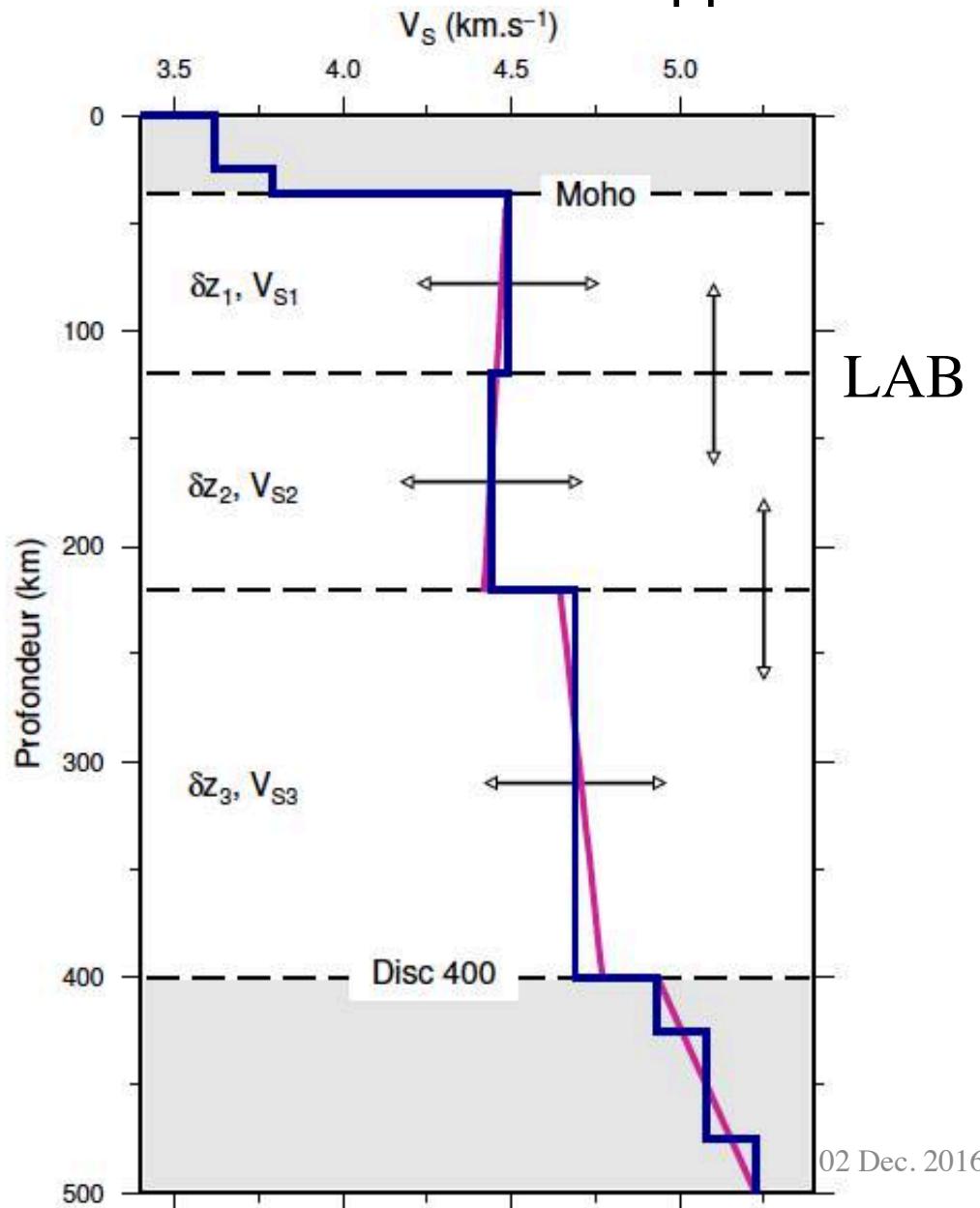


Much discrepancy between different estimates:
Global surface wave tomographies give 200-250km depth for continental roots

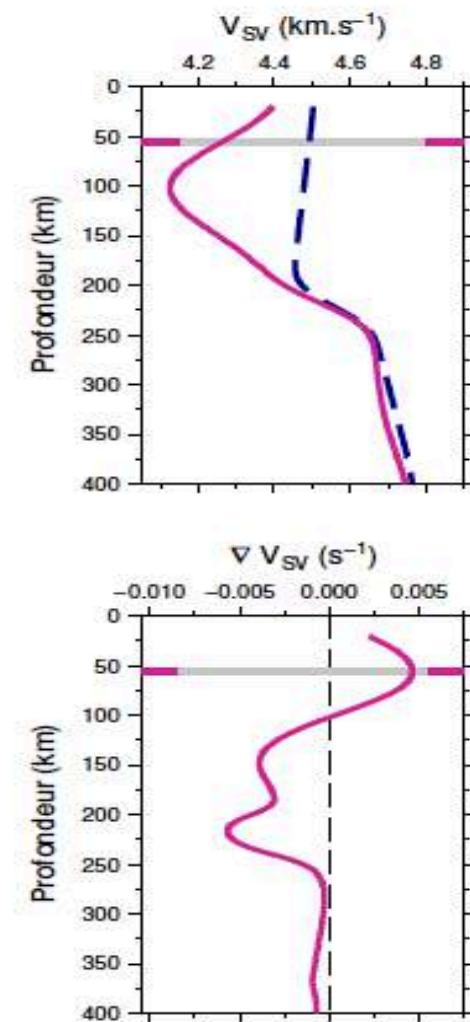
Mid-lithospheric Discontinuity (Yuan & Romanowicz, 2010)

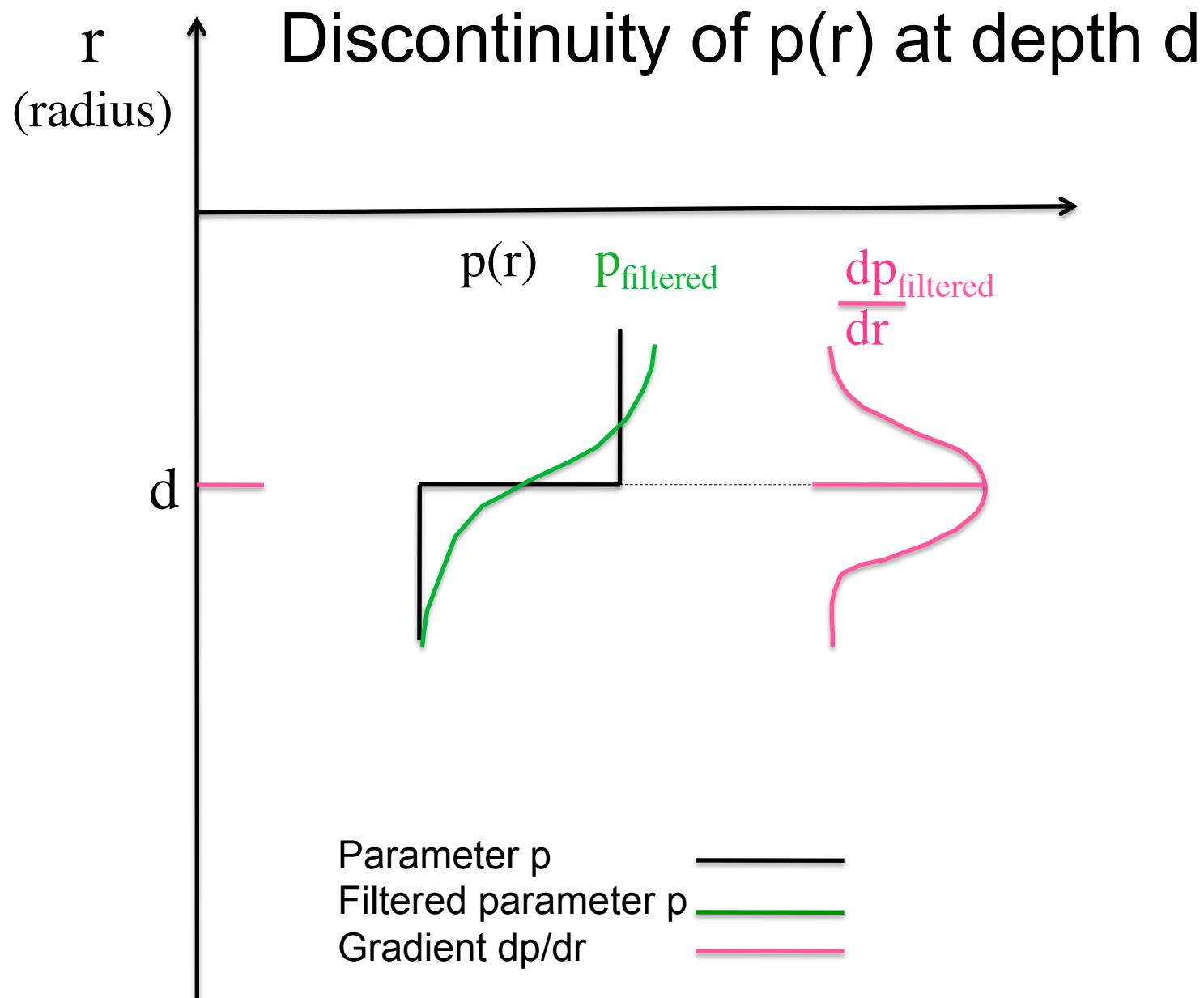
From Surface wave dispersion

Statistical Monte-Carlo Approach First order Perturbation theory



Proxy from parameter V_{sv}





Other proxies from parameters obtained from anisotropy tomographic models

Well resolved parameters:

V_{SV} S-wave velocity

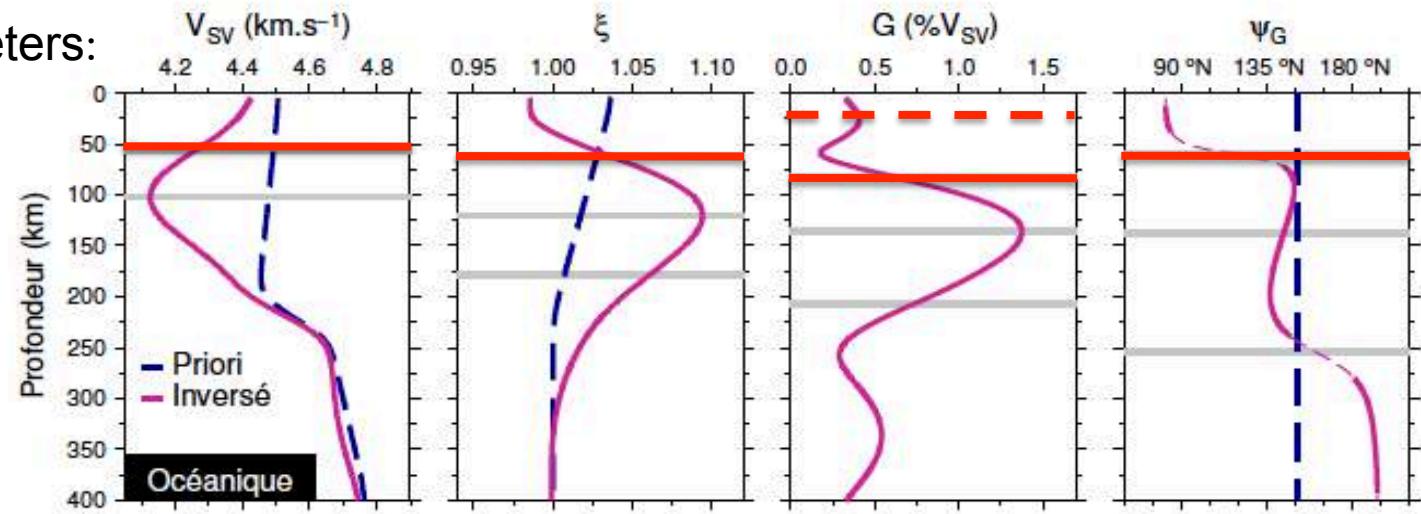
ξ , radial anisotropy

G, Ψ_G S-wave

azimuthal anisotropy

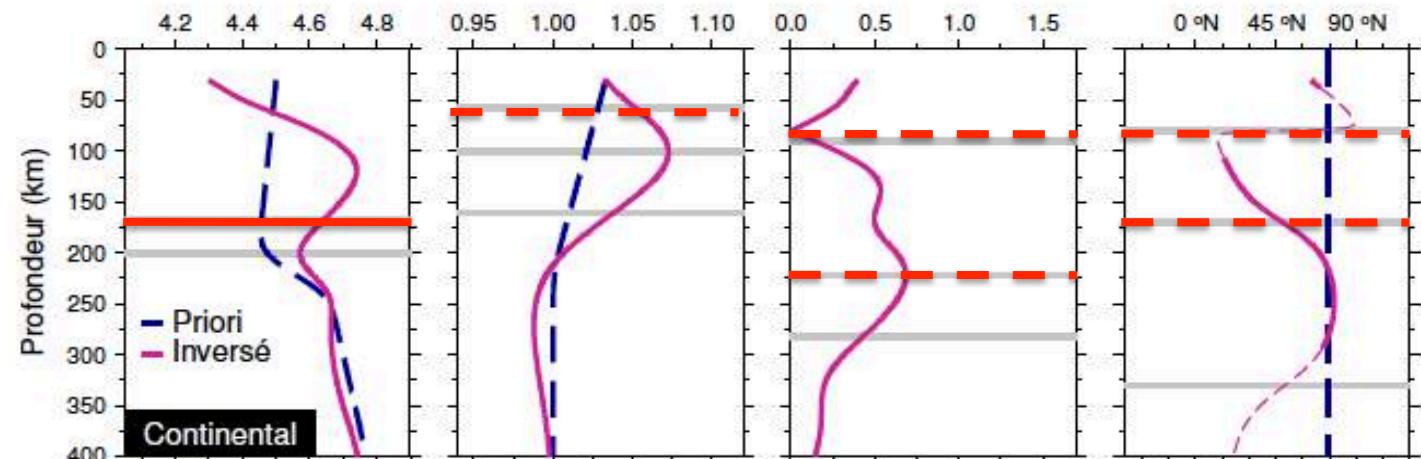
Oceanic profile

$\lambda=35^\circ, \phi=-35^\circ$



Continental profile

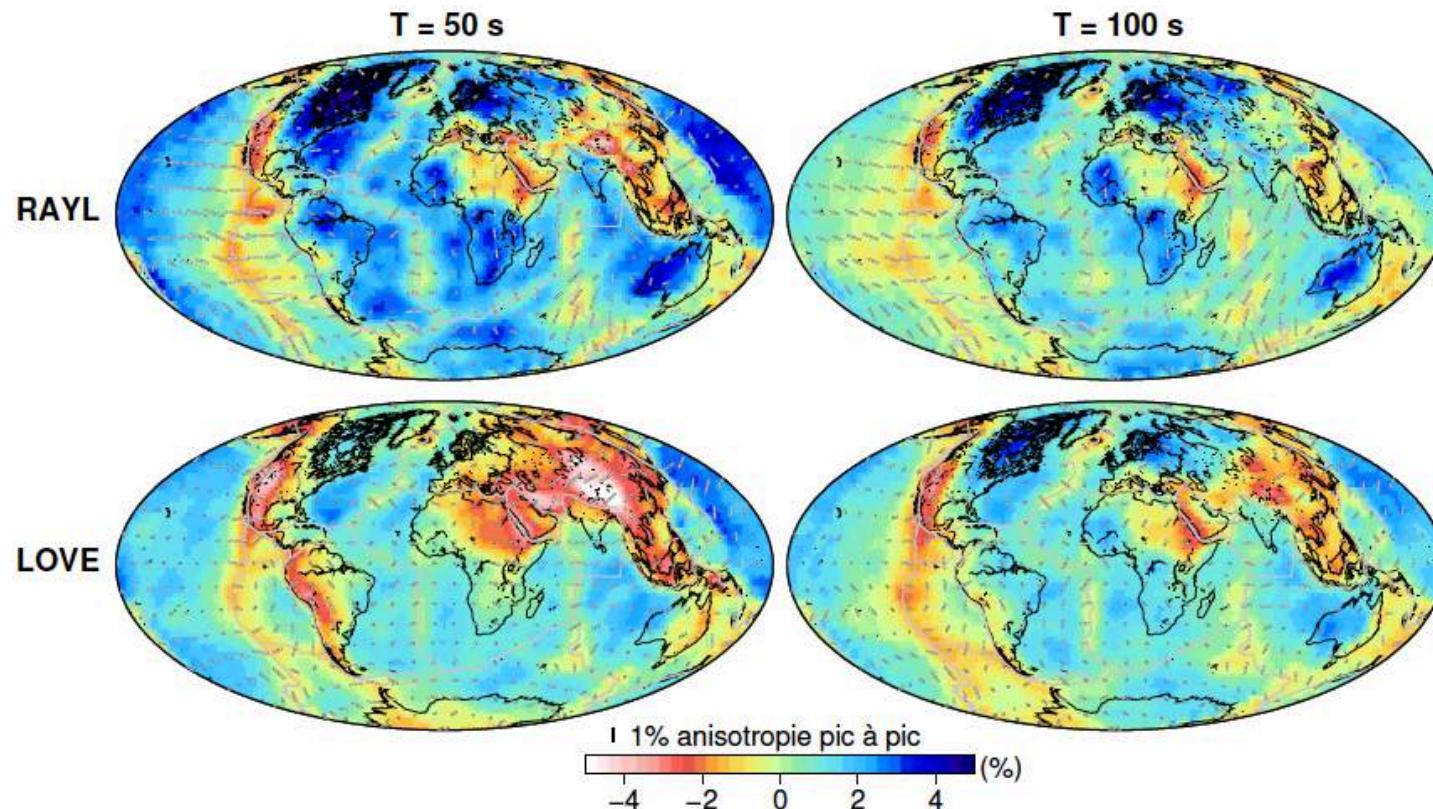
$\lambda=63^\circ, \phi=-96^\circ$



**Collection of surface wave dispersion data
(IPGP, Harvard, Utrecht, Boulder)**

First step: Regionalization => local dispersion velocity $V(T, \theta, \phi, \psi)$

Rayleigh phase velocity and azimuthal anisotropy



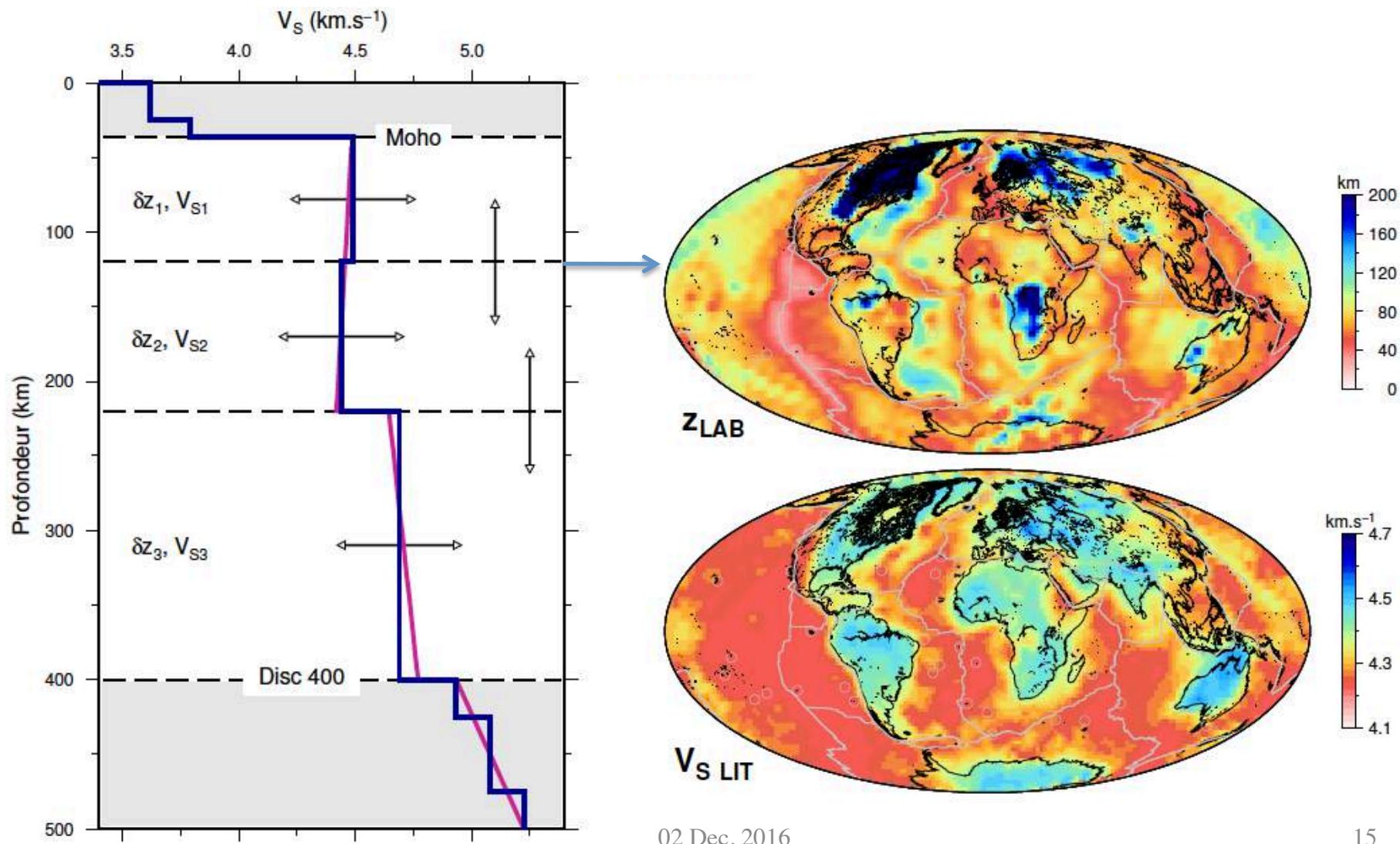
Second step: Inversion at depth

Statistical Monte-Carlo Inversion

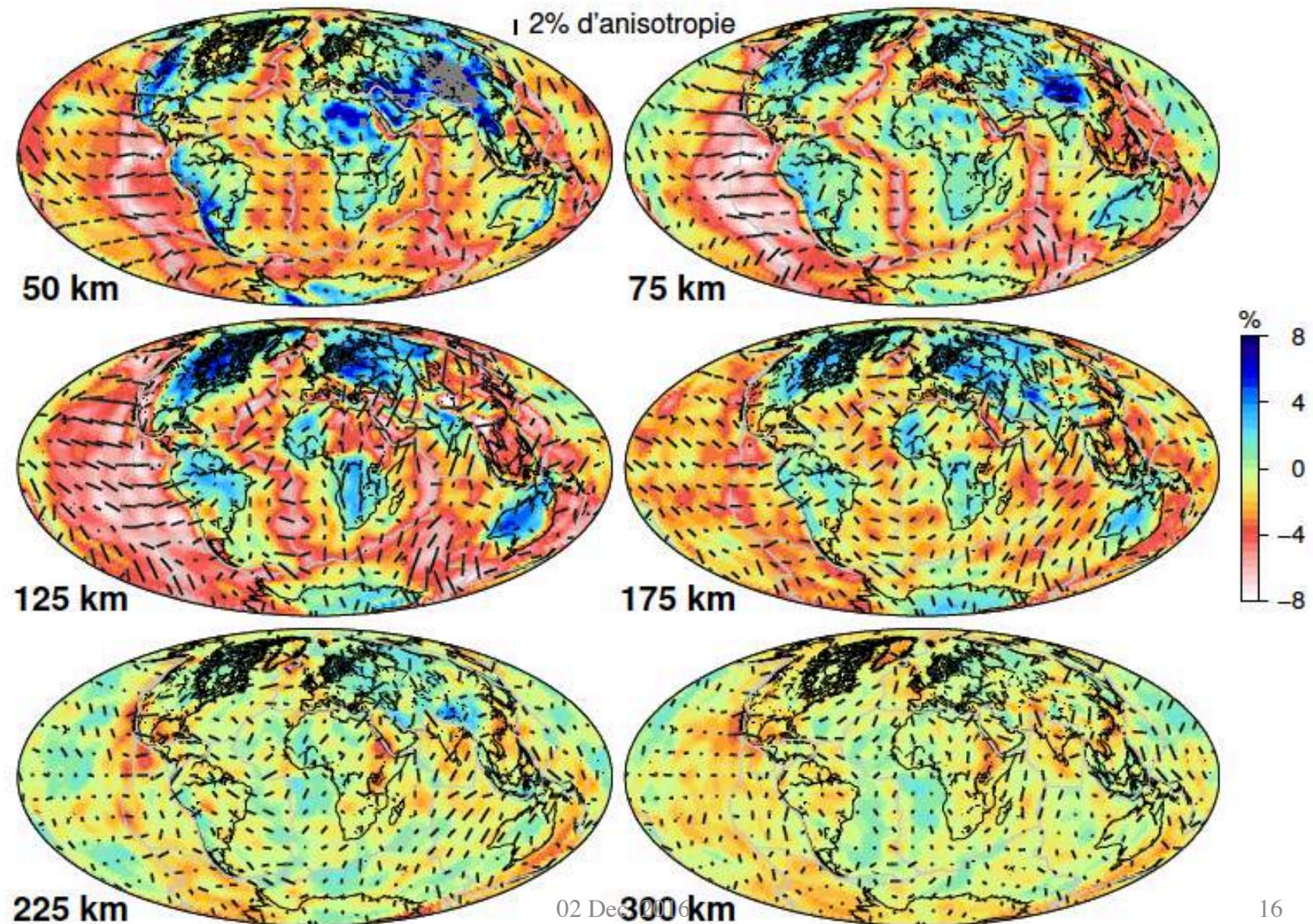
First order Perturbation

LAB: Statistical M.C. Isotropic Inversion

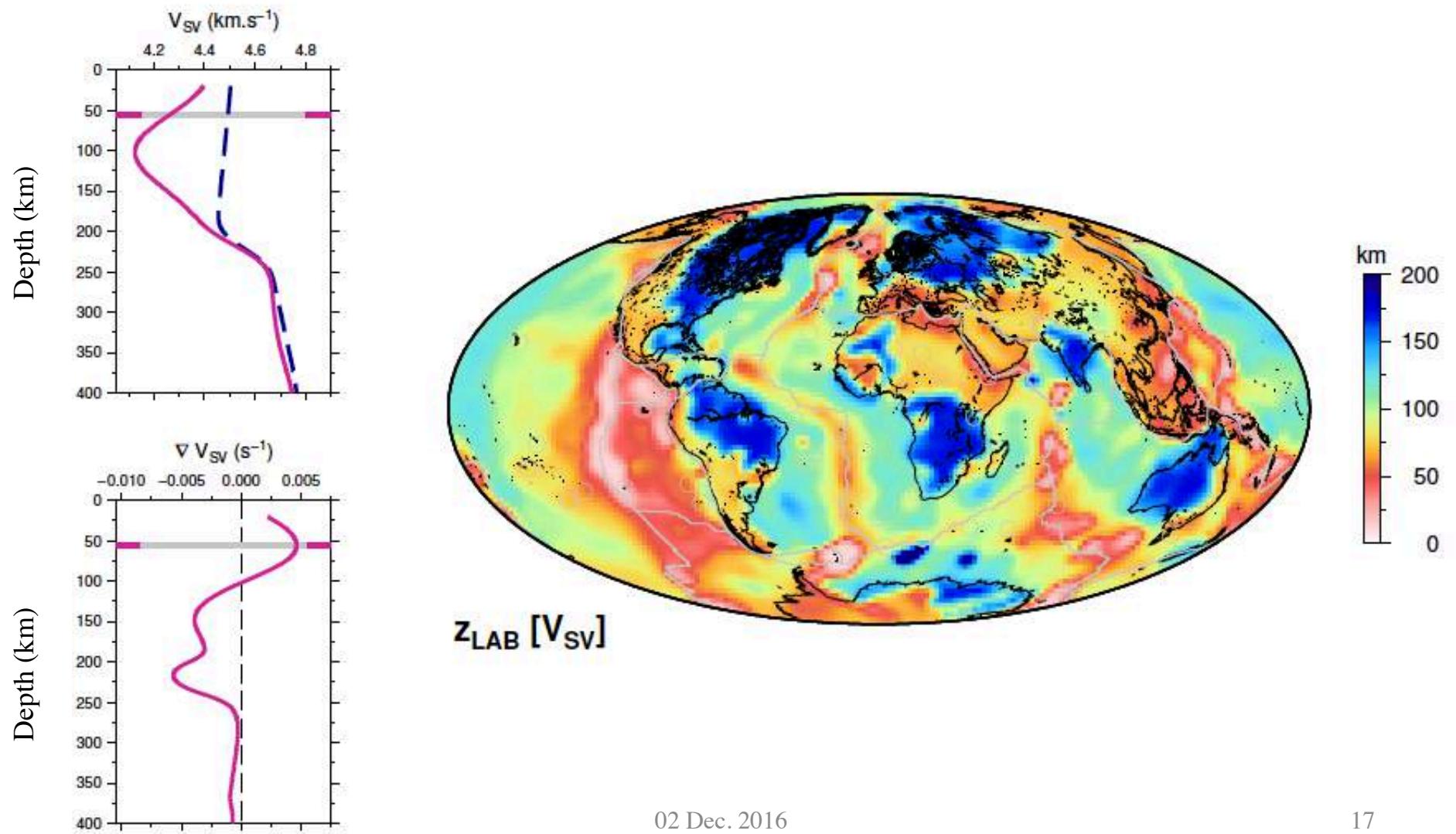
Data: C_R, C_L, U_R, U_L [30-300s], Parameters: 3Vs, 2 δz



First order perturbation Theory => depth distribution of V_{sv} , G (and ξ)

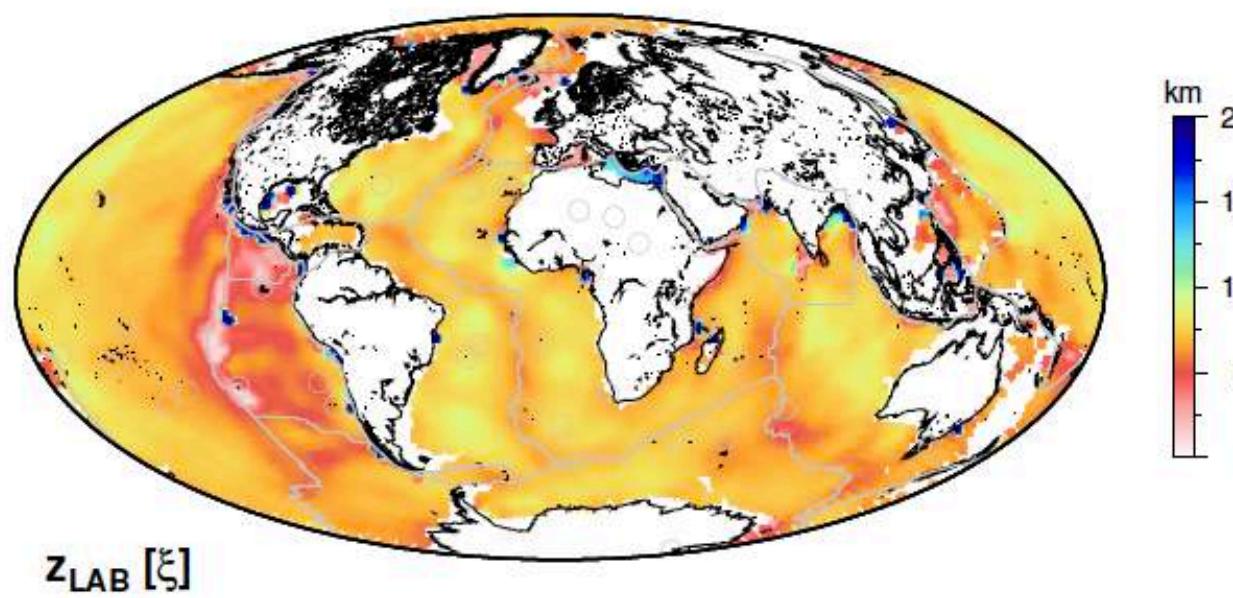
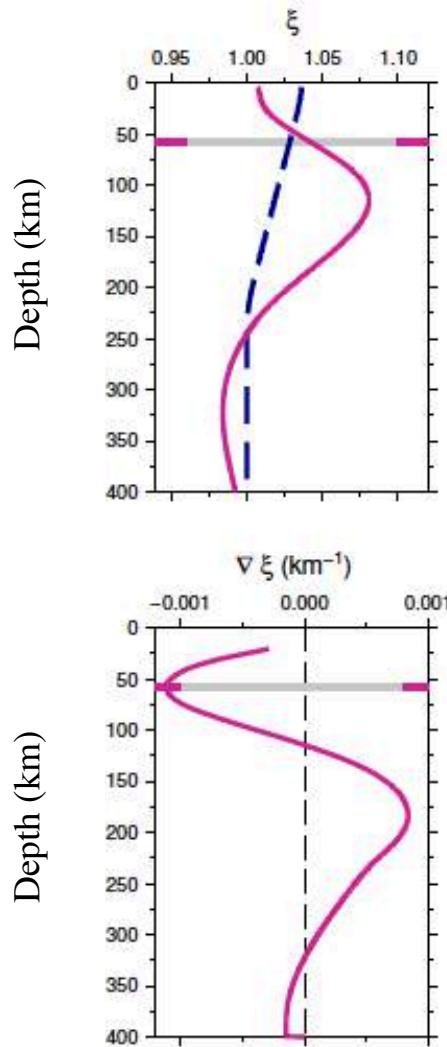


LAB from the gradient of VSV parameter

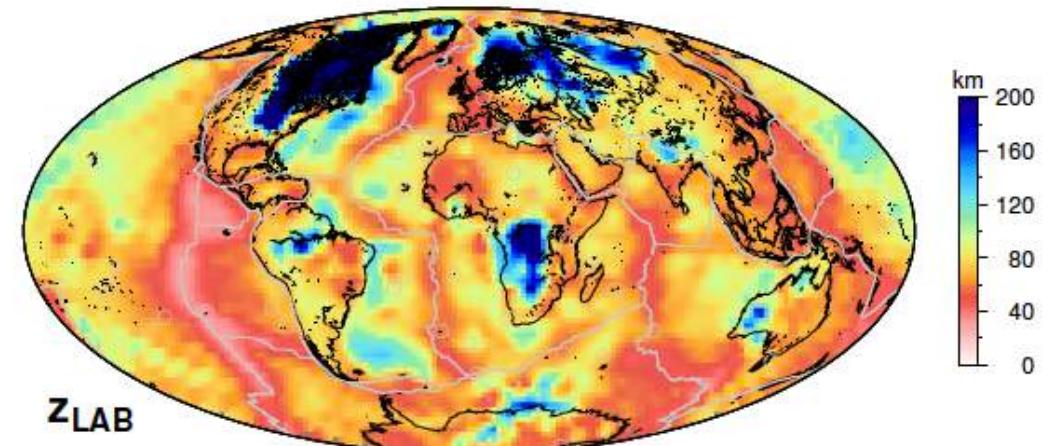


LAB from the gradient of ξ parameter (only oceans)

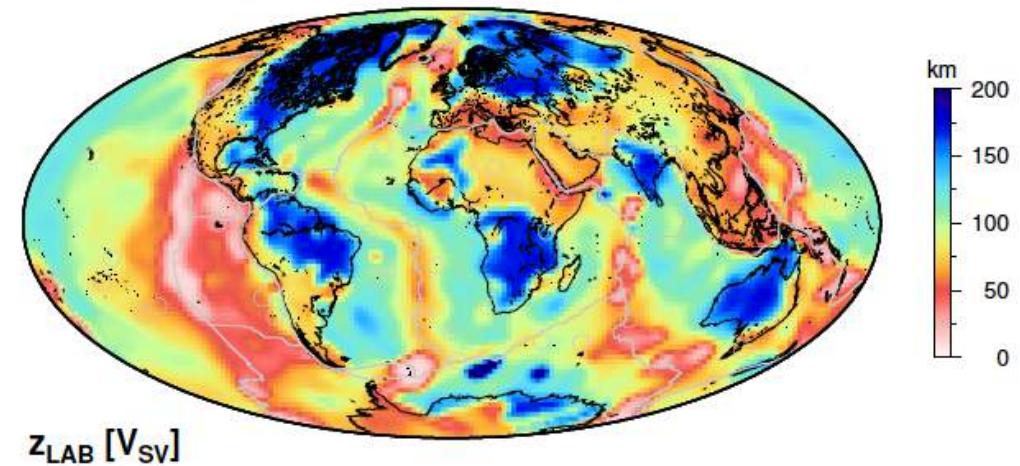
$$\text{Radial anisotropy } \xi = (V_{SH}/V_{SV})^2$$



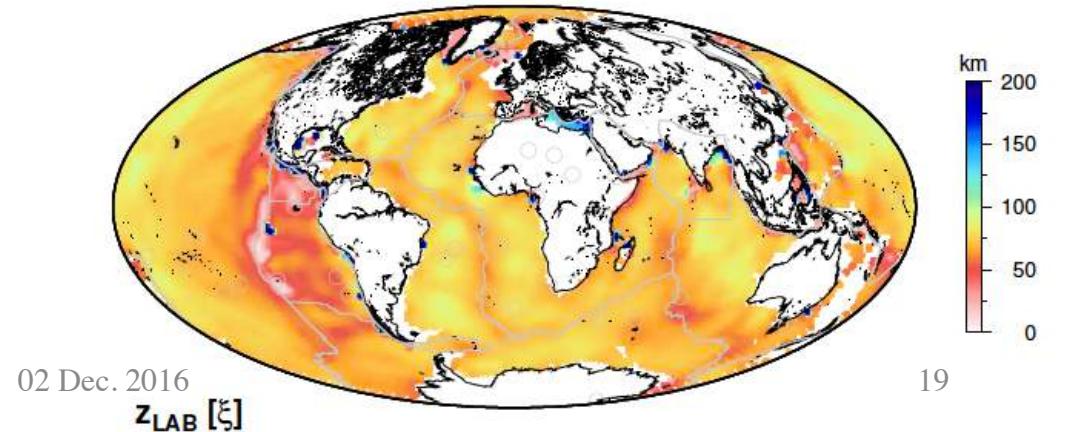
Statistical MC Isotropic Inversion



Vsv proxy (1st order Perturbation Theory)



ξ proxy (1st order Perturbation Theory)

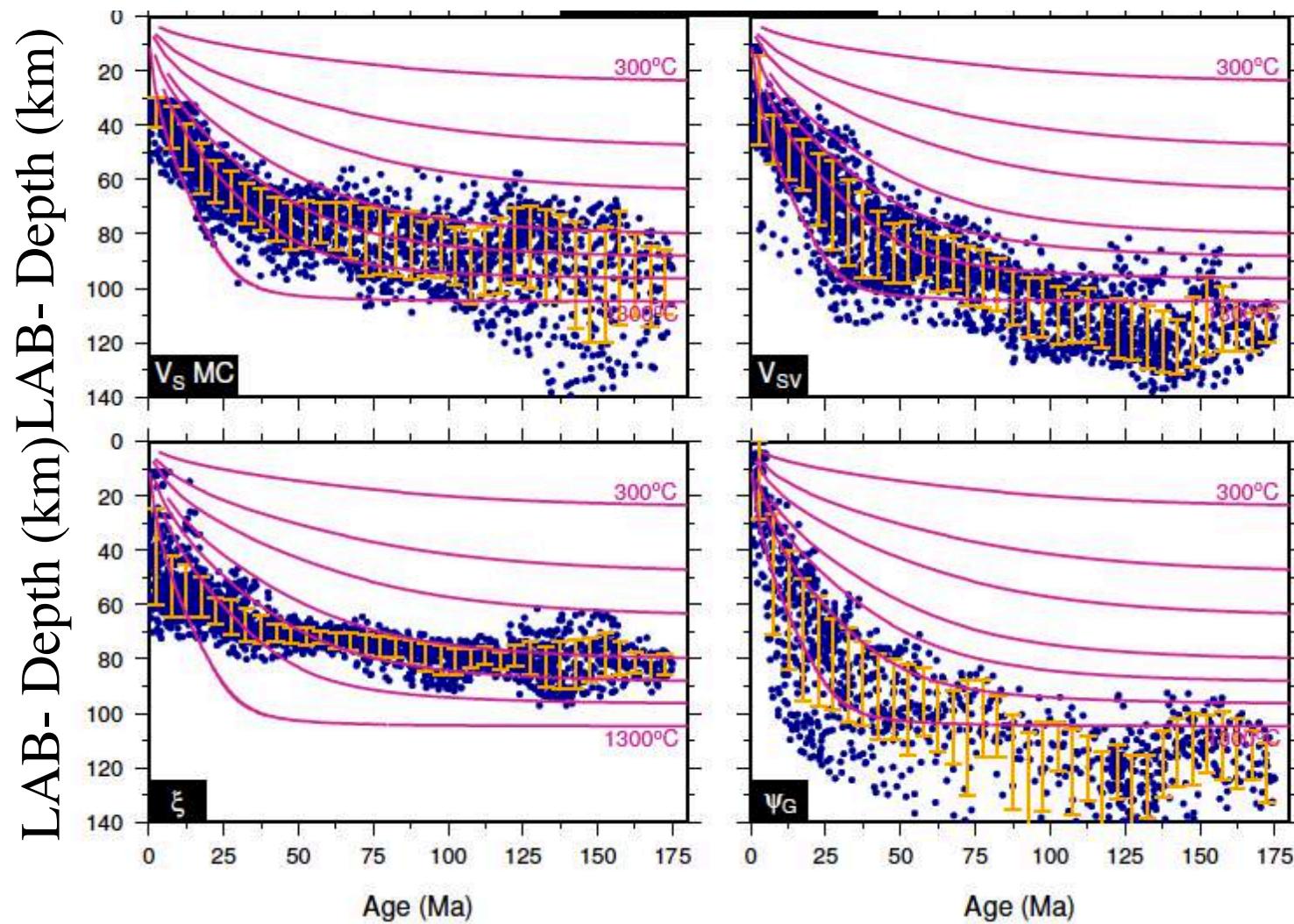


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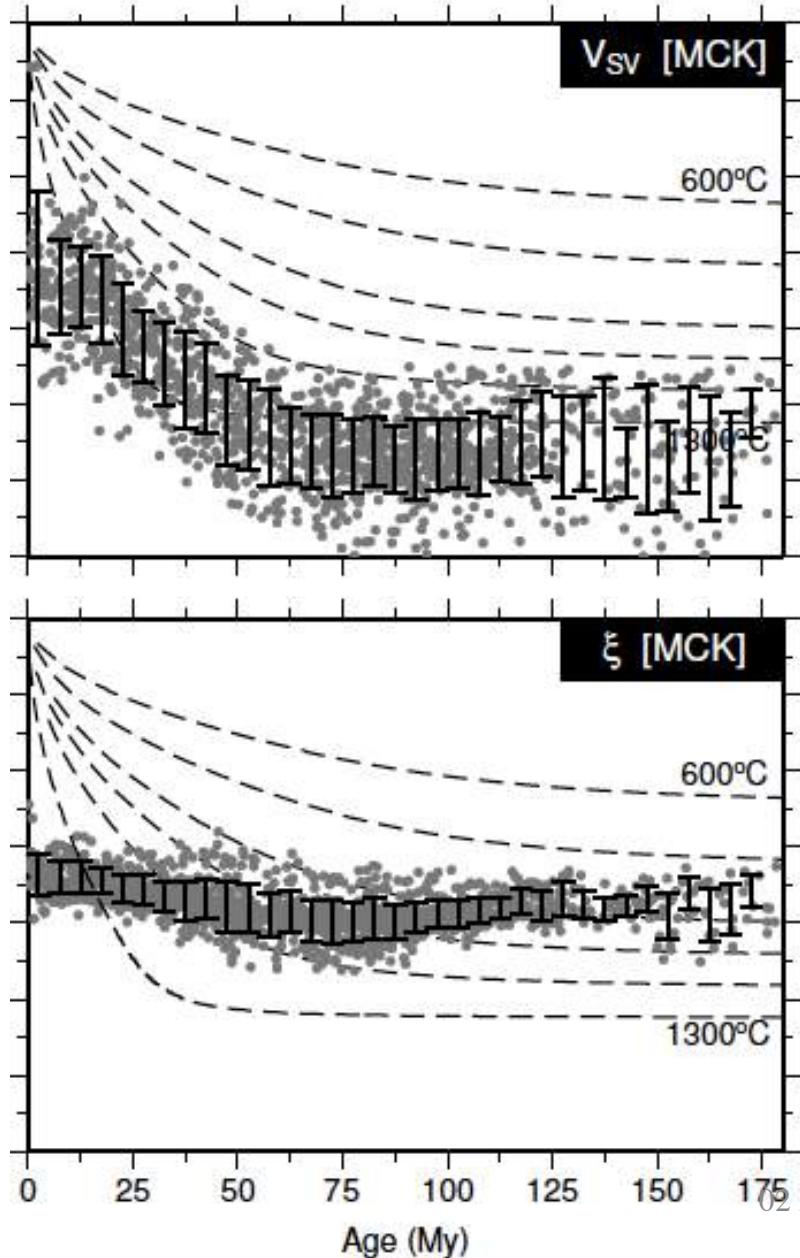
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Age Variation of LAB depth in oceanic regions

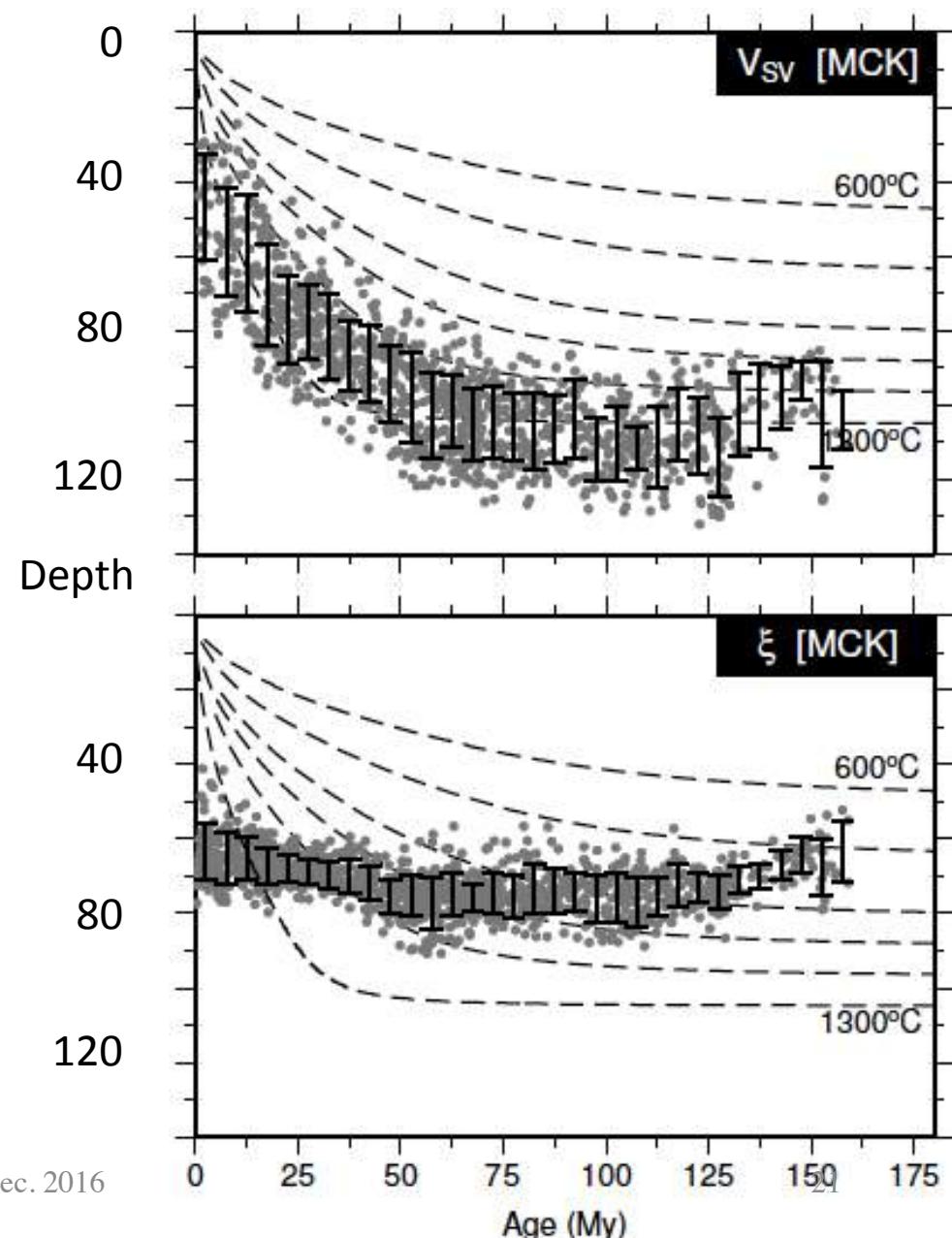
Compared with Plate model



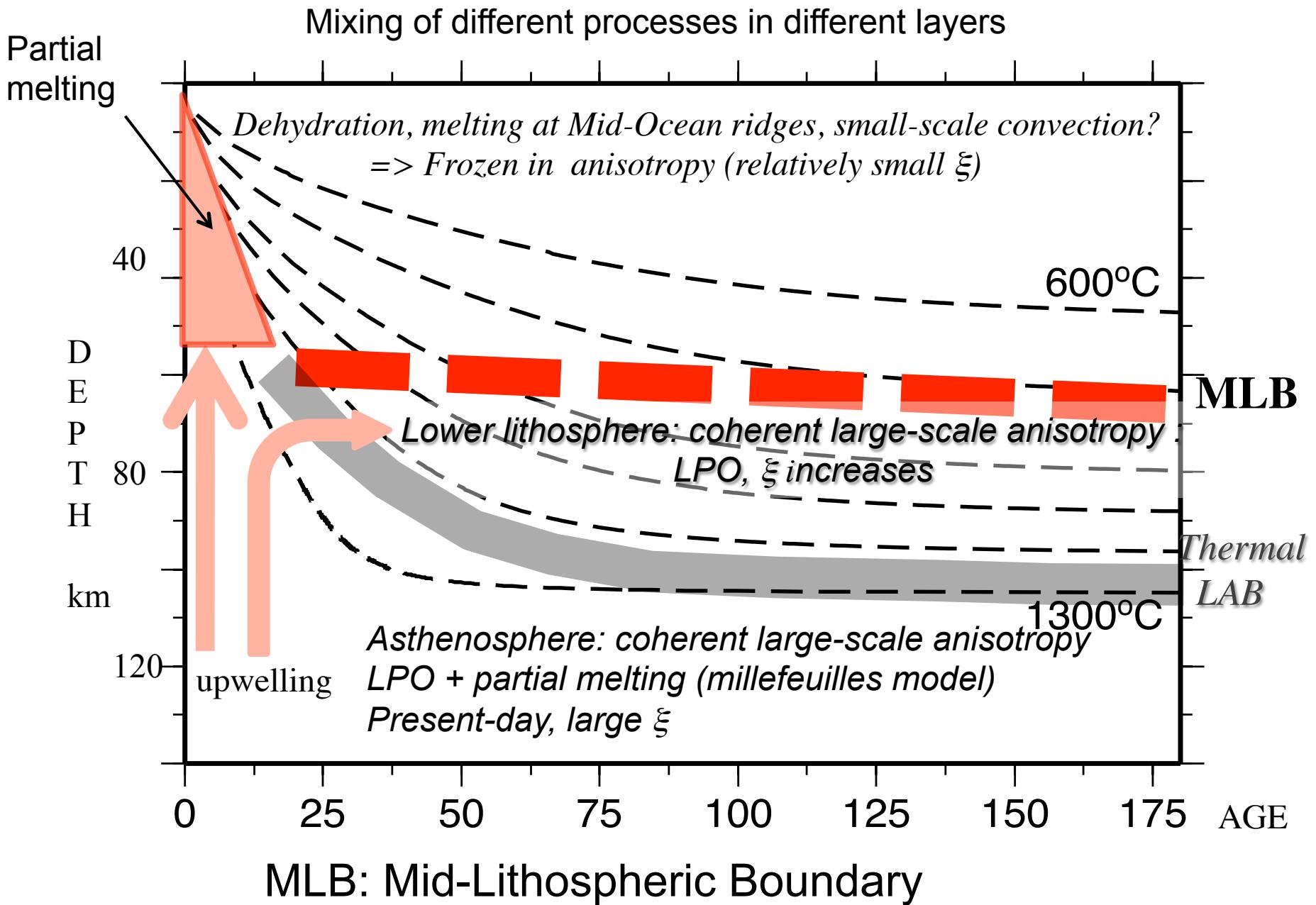
Atlantic Ocean



Indian Ocean



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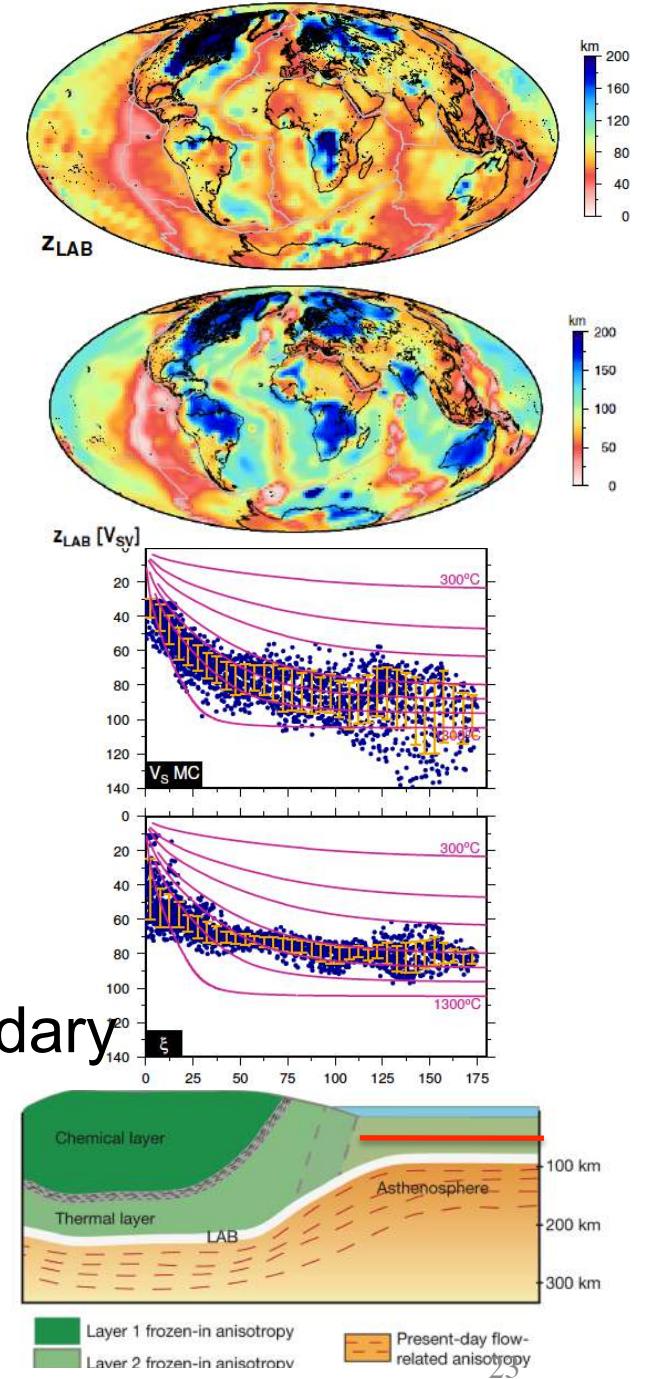
New Discontinuity within the lithosphere

- LAB topography derived from surface wave data on a global scale
- The ocean lithosphere not so simple!
 - For oceans, the model of formation of lithosphere must be revisited in view of results from radial and azimuthal anisotropies.
- Existence of a strong gradient of ξ between 60-80km (plate; related to dehydration boundary layer?)

Mid-Lithospheric Boundary

CONTINENTS?

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Yuan and Romanowicz, 2010

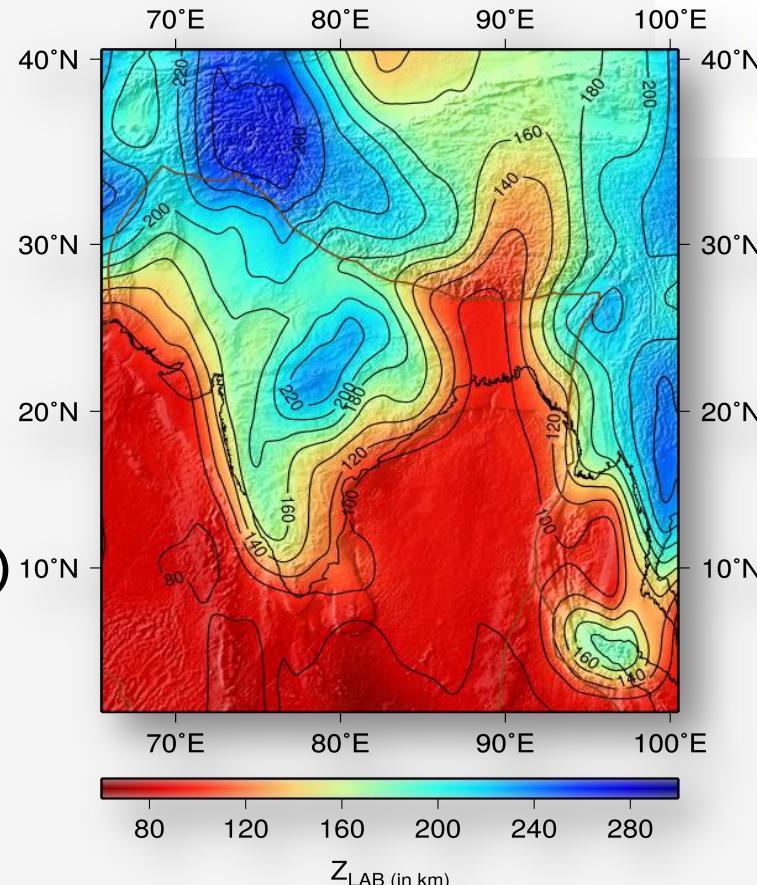
Deep Structure of the Indian Continent



1: IPG-Paris, France (J.-P. Montagner)



2: NGRI-Hyderabad, India (Ravi Kumar)

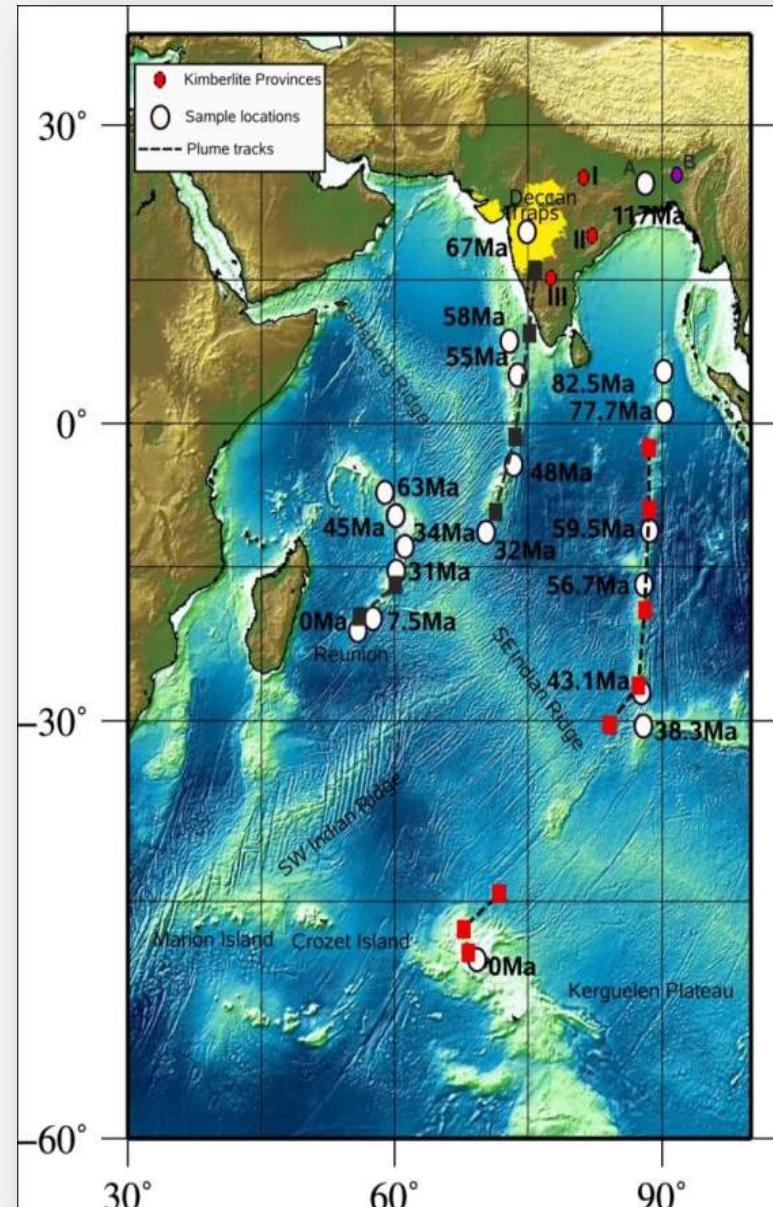


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Satish Maurya
IPG-Paris

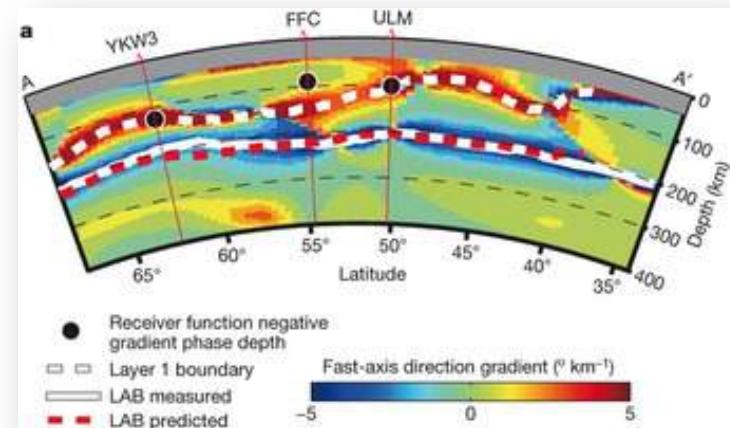
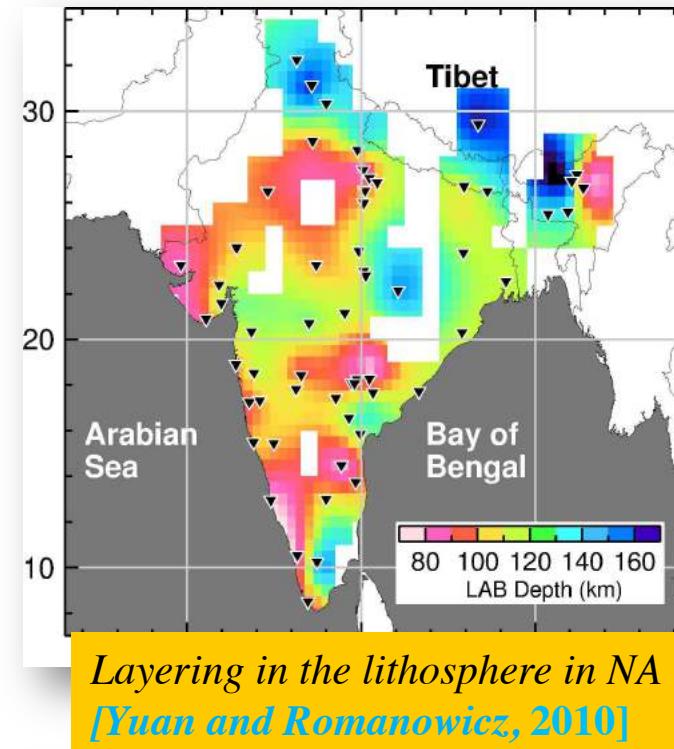
Motivation and Scientific Challenges

- Indian continent is **unique** in many respects.
- Indian plate moved at exceptionally high speeds of **18-20 cm/yr** after its breakup from Gondwanaland ~65 Myr. Ago (Paleomagnetic data).
- Ravaged by hotspots and experienced large scale magmatism.
- Five cratons. Indian continental lithosphere might be underlain by an anomalously hot sub-lithospheric mantle
- Interaction with plumes (*Réunion, Marion, Crozet and Kerguelen*)

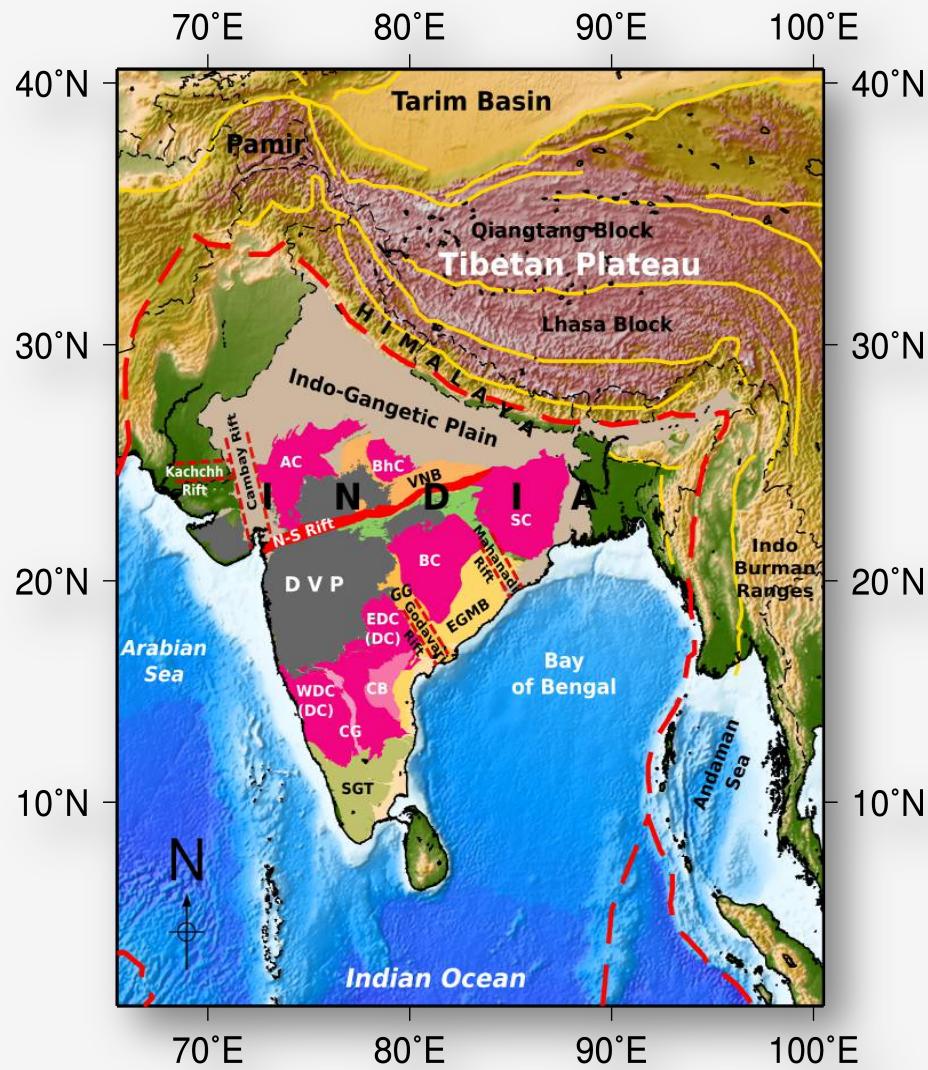


Scientific Challenges – Debate on Indian LAB

- **Super mobility** due to a thin seismic lithosphere (~100km) [Kumar et al., 2007 & Negi et al., 1986] - in total disagreement with common consensus on cratons.
- Pressure-Temperature studies on mantle xenoliths reveal a thick lithosphere that is **~230 km** at ~1100 Ma. underneath **Dharwar Craton** [Griffin et al., 2008].
- **~175 km** at ~ 65 Ma underneath **Bastar craton** [Babu et al. 2009].
- Evidence for postcollisional flexuring of the Indian plate with a wavelength of ~1000 km



Study Area



Precambrian

- South Granulite Terrane (SGT)
- Cratons (AC,BC,BhC,DC,SC)
- Closepet Granite(CG)
- Cuddapah Basin (CB)
- Eastern Ghat Mobile Belt (EGMB)
- Godavari Graben (GG)
- Vindhyan Basin (VNB)

Phanerozoic

- Deccan Volcanic Province (DVP)
- Gondwana Rocks (GR)
- Indo-Gangetic Plains (IGP)
- Alluvium

AC: Aravalli craton

BC: Bastar craton

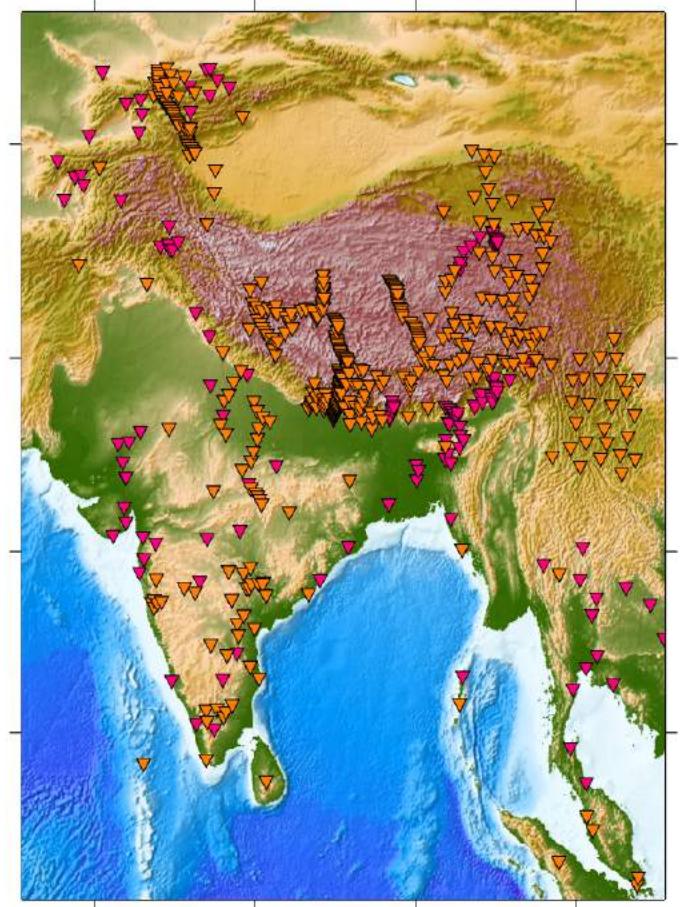
BhC: Bundhelkhand craton

DC: Dharwar craton

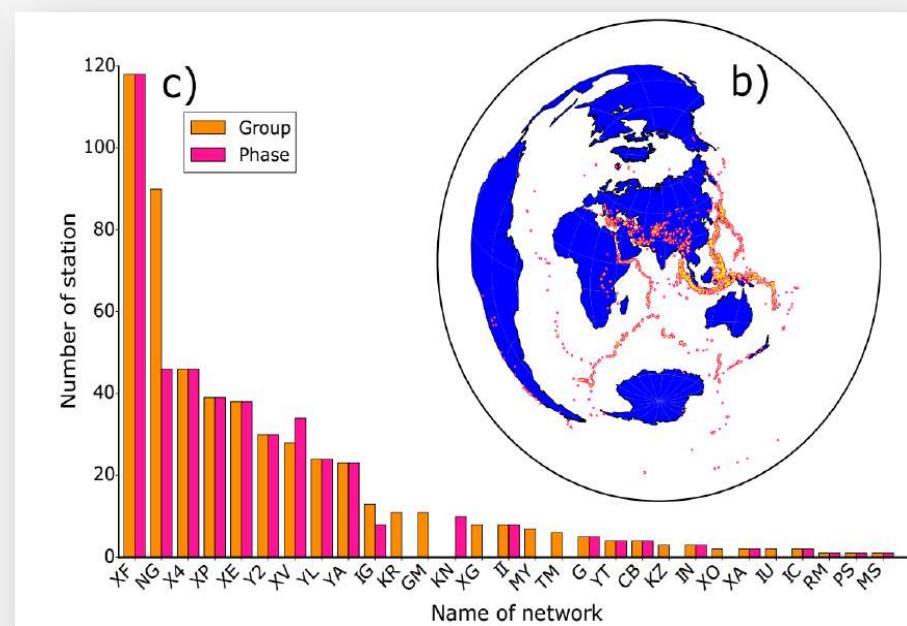
SC: Singhbhum craton

Data

Stations



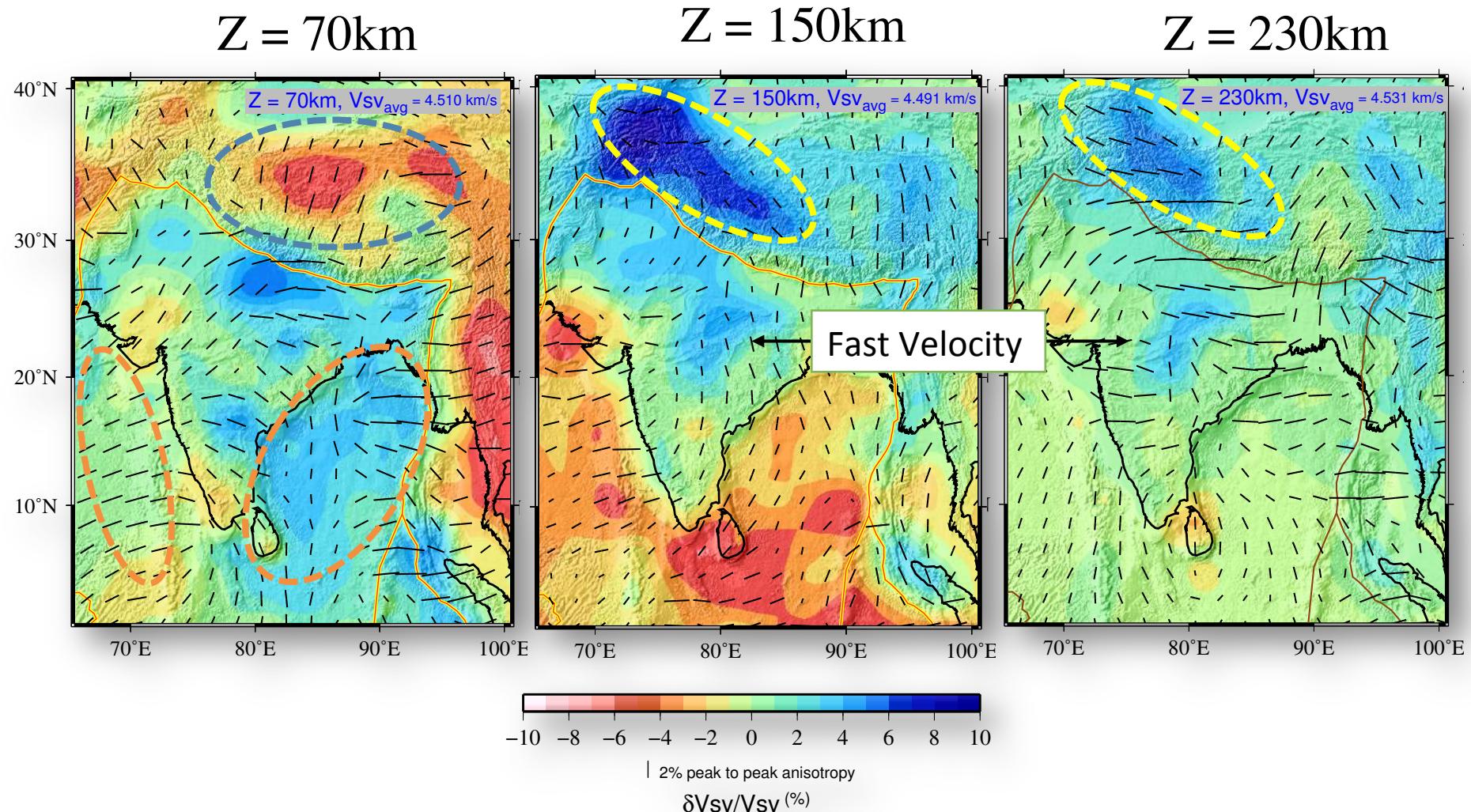
Earthquakes



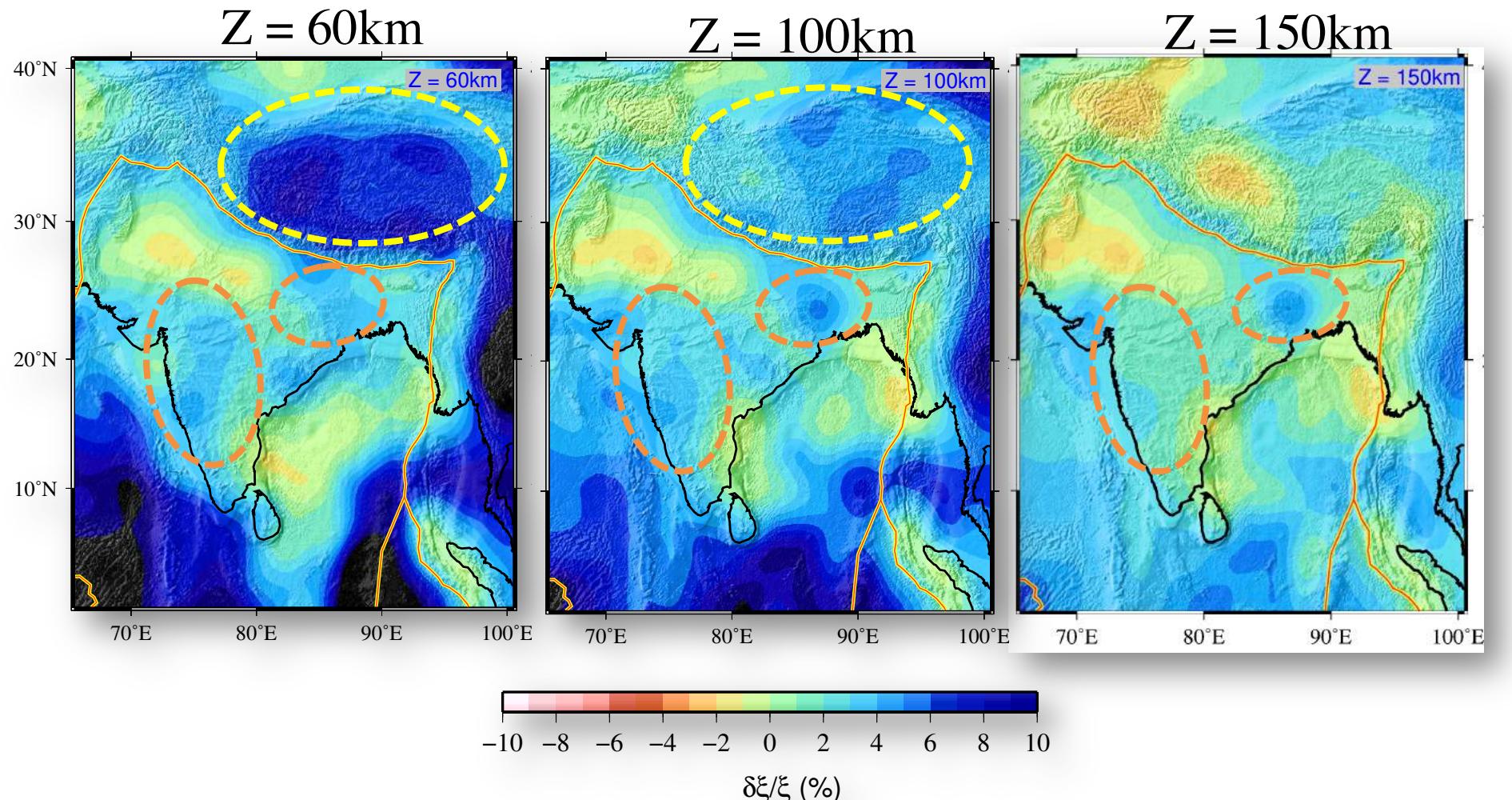
- 29 Seismic broadband Networks (**global and regional**)
- Over 550 seismic stations
- Earthquakes of magnitude >5.5
- Surface wave data in the period range of 10-400s.

3-D tomography model of the Indian continent

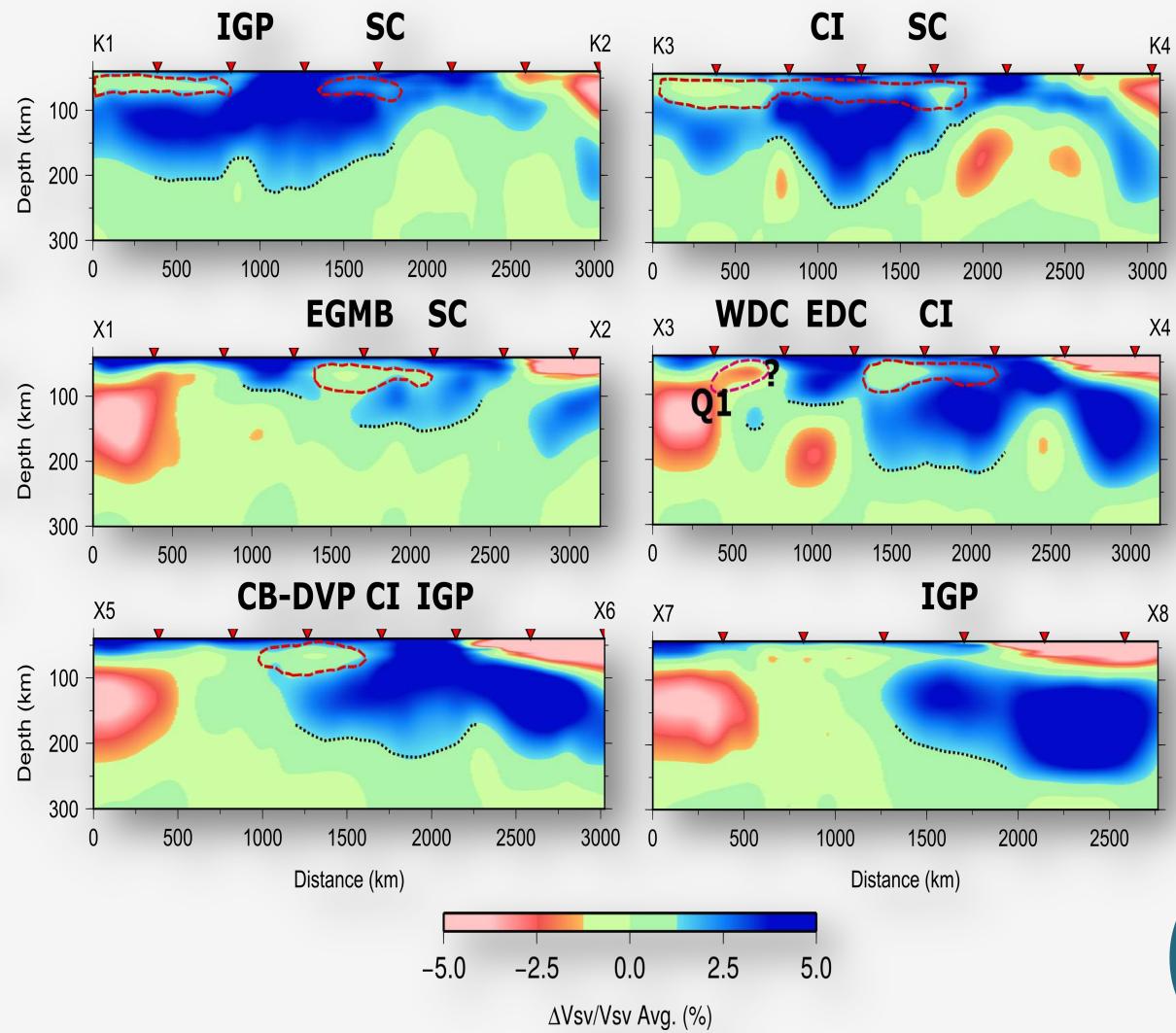
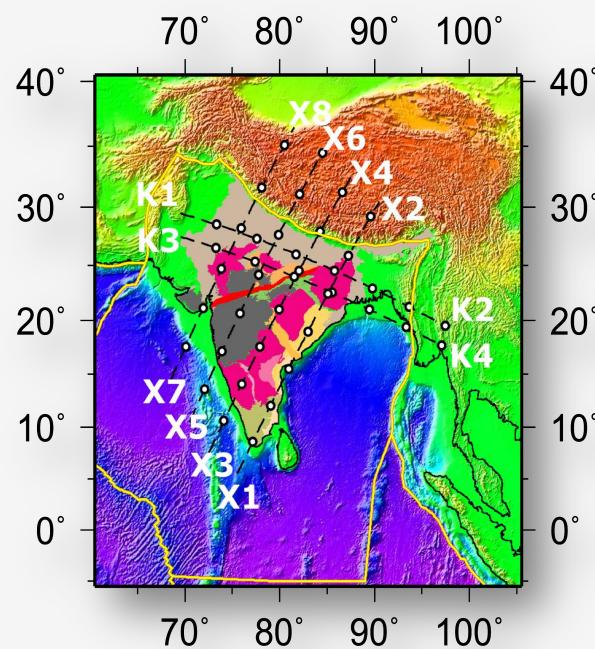
Velocity and Azimuthal Anisotropy



3-D tomography model of the Indian continent Radial Anisotropy ξ

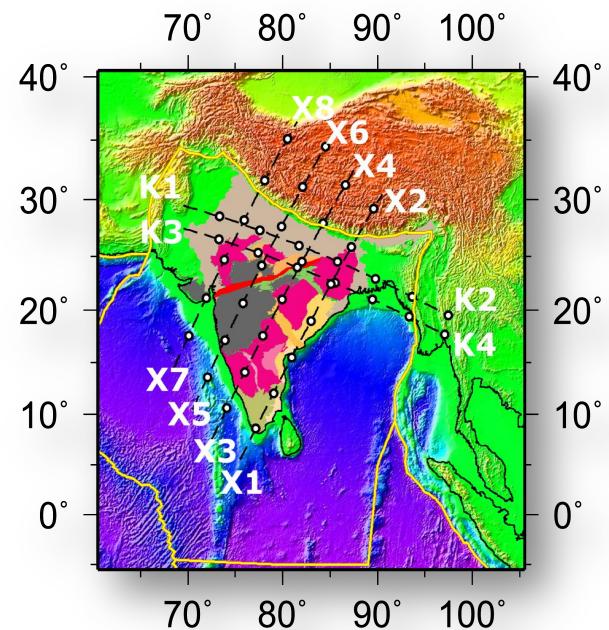


3D-Perturbation model

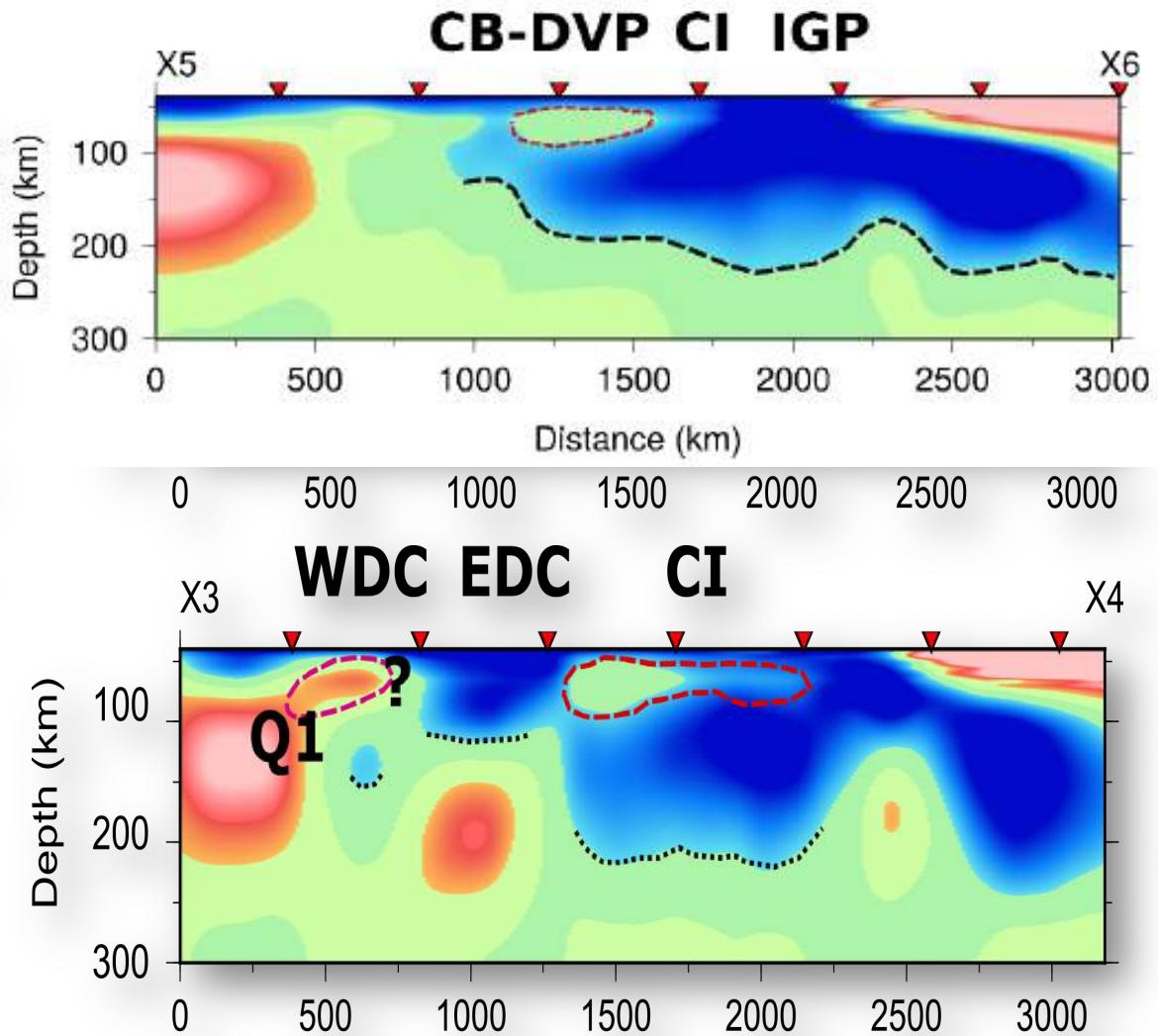


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3D-Perturbation model



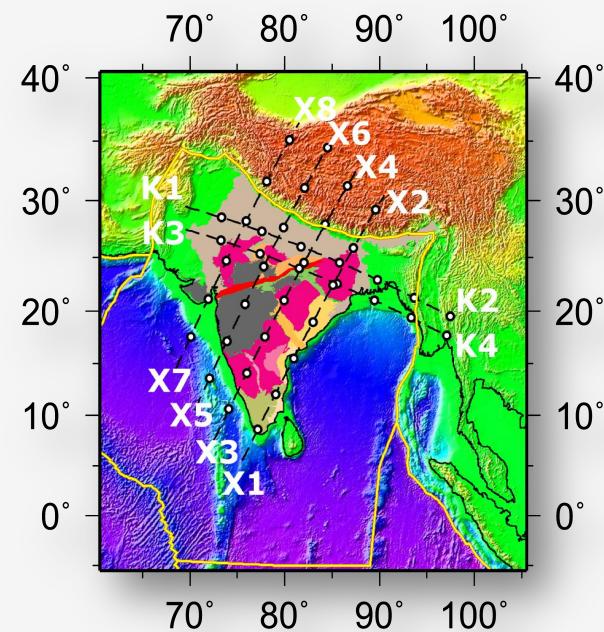
DVP: Low velocity zone
Remnant of hotspot birth



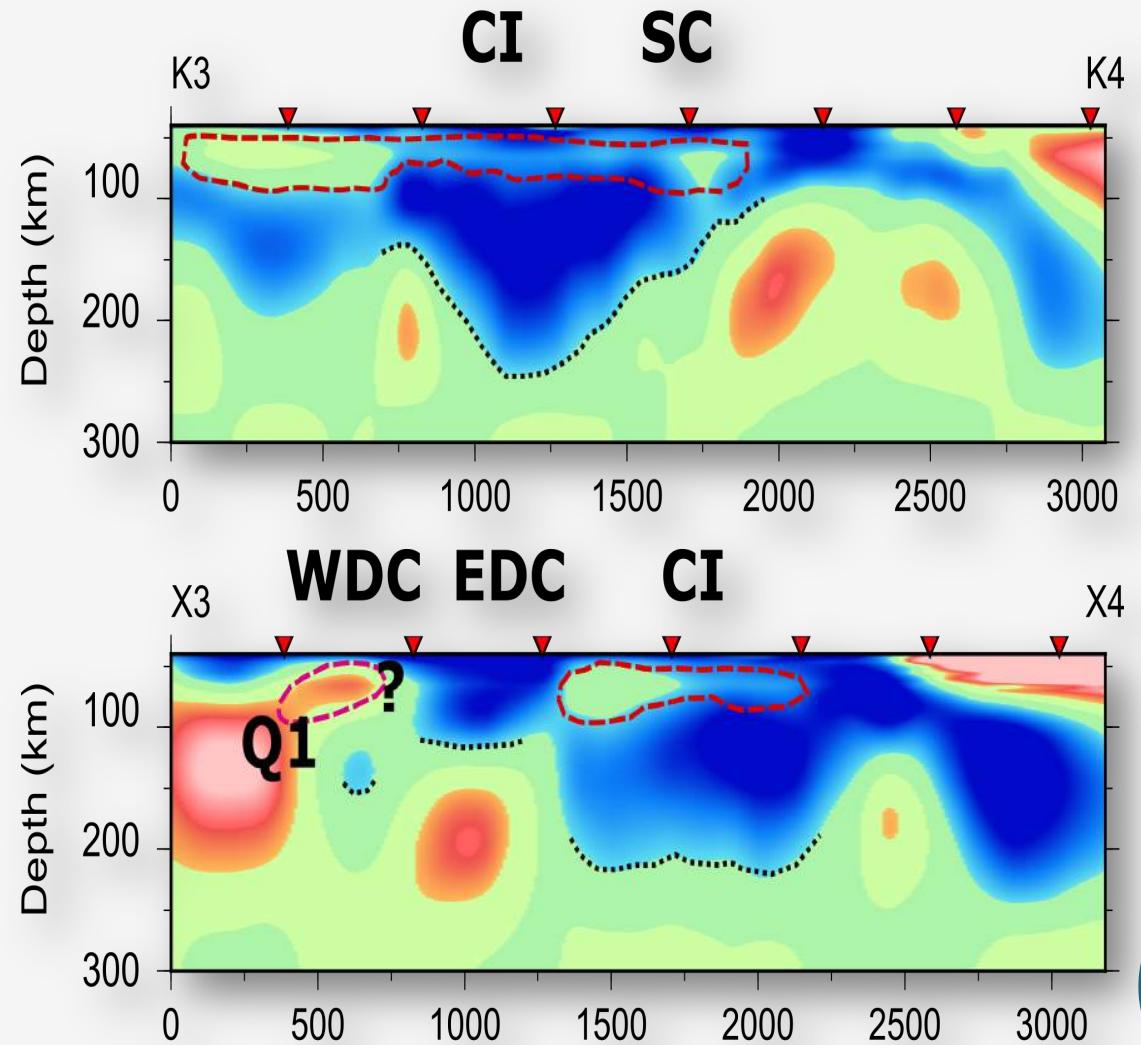
MLB: Mid-lithospheric low velocity zone

3D-Perturbation model

Indian Keel



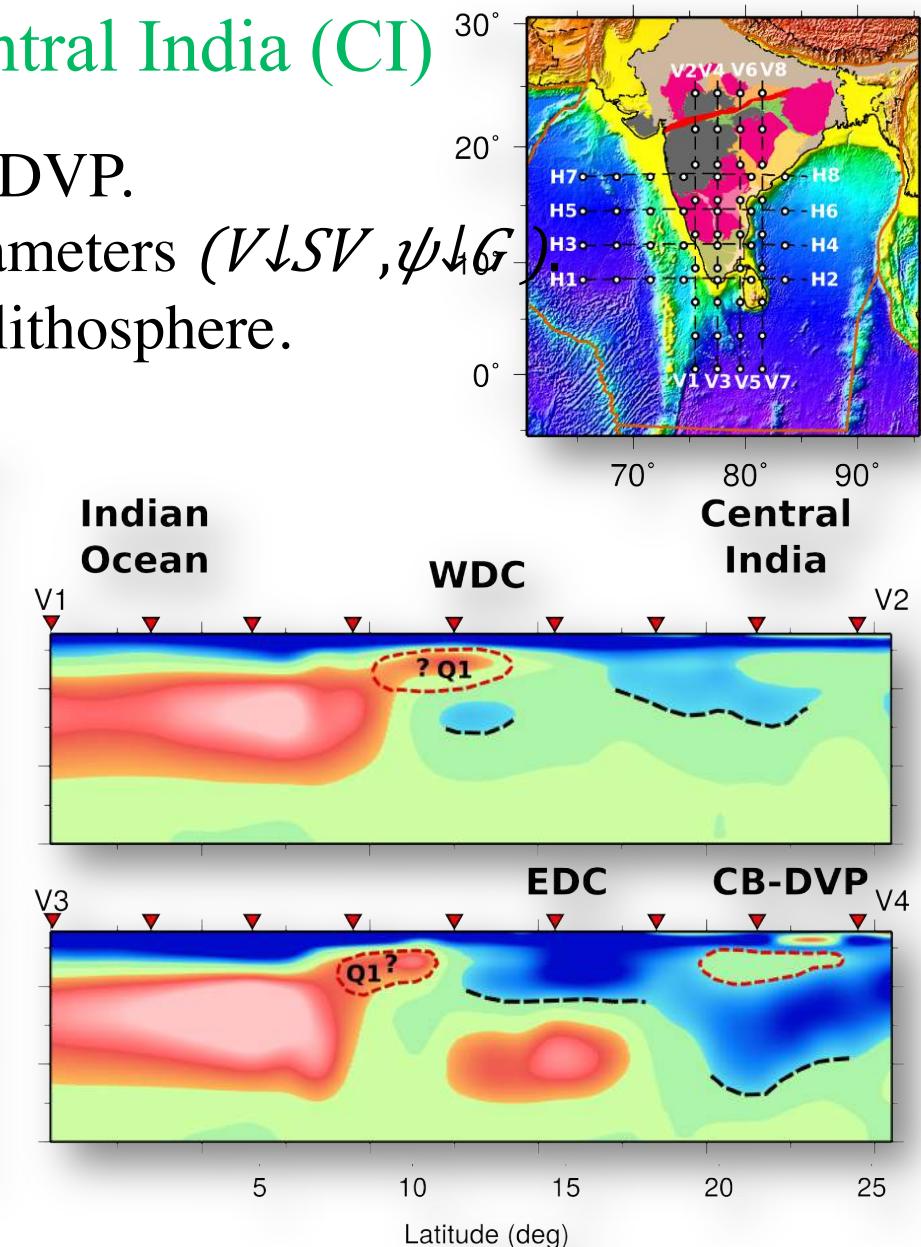
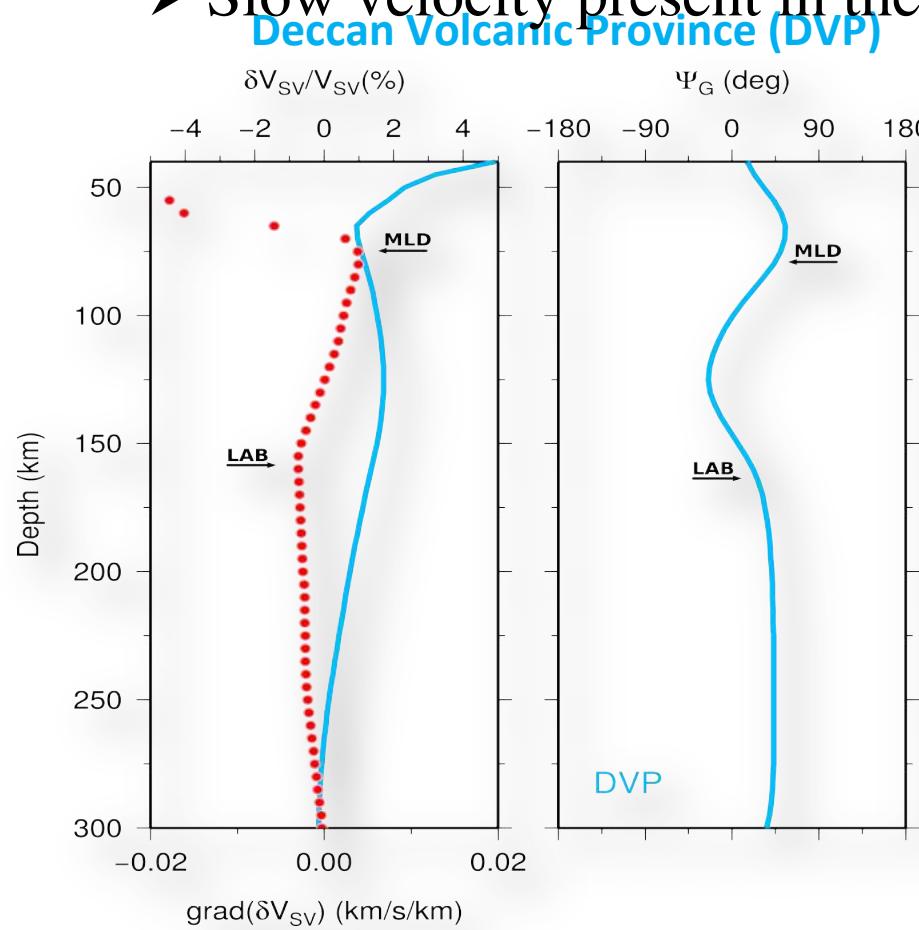
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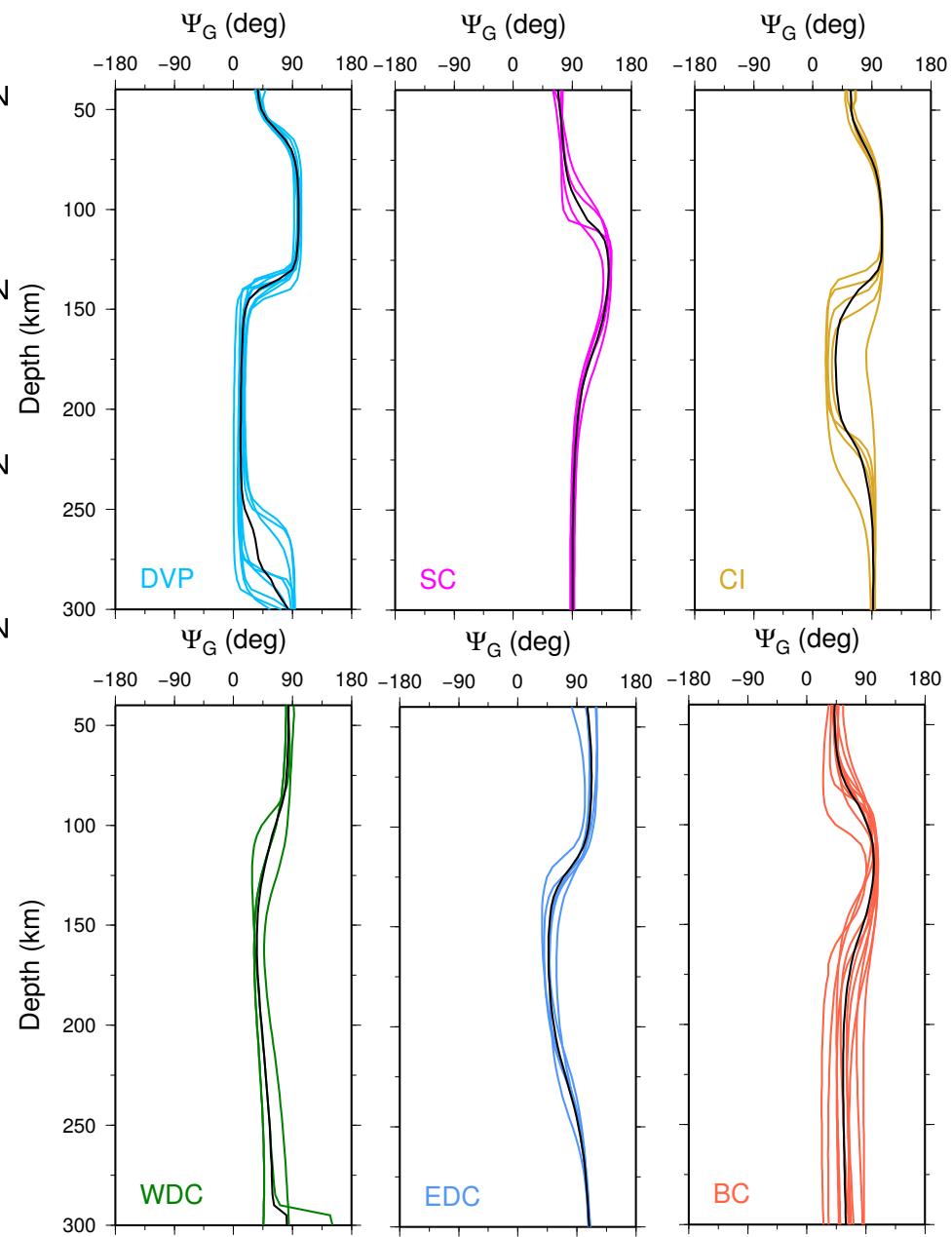
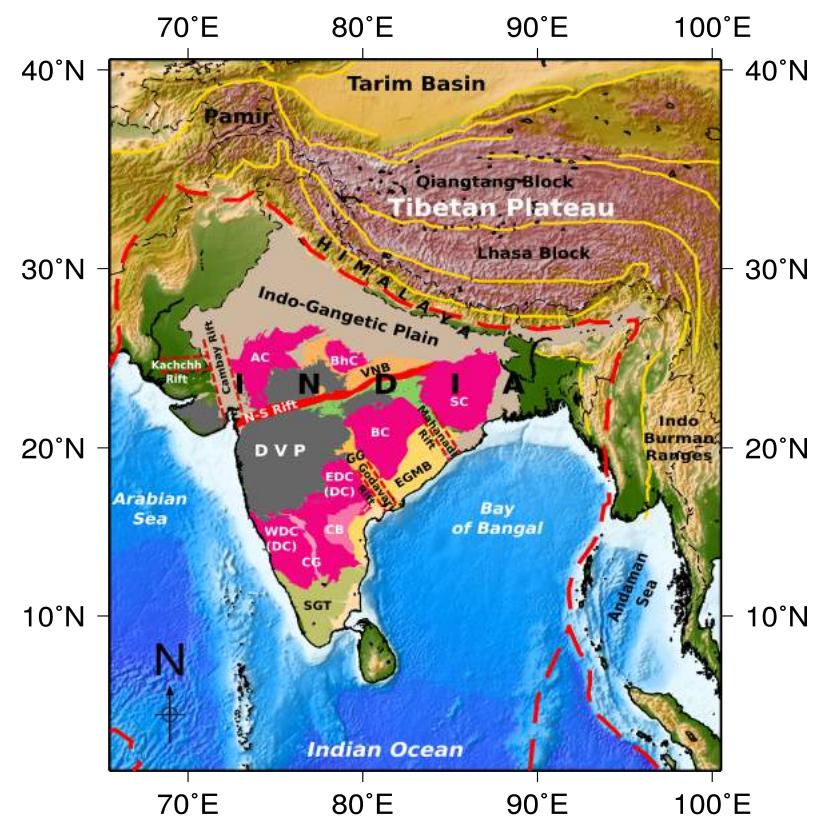


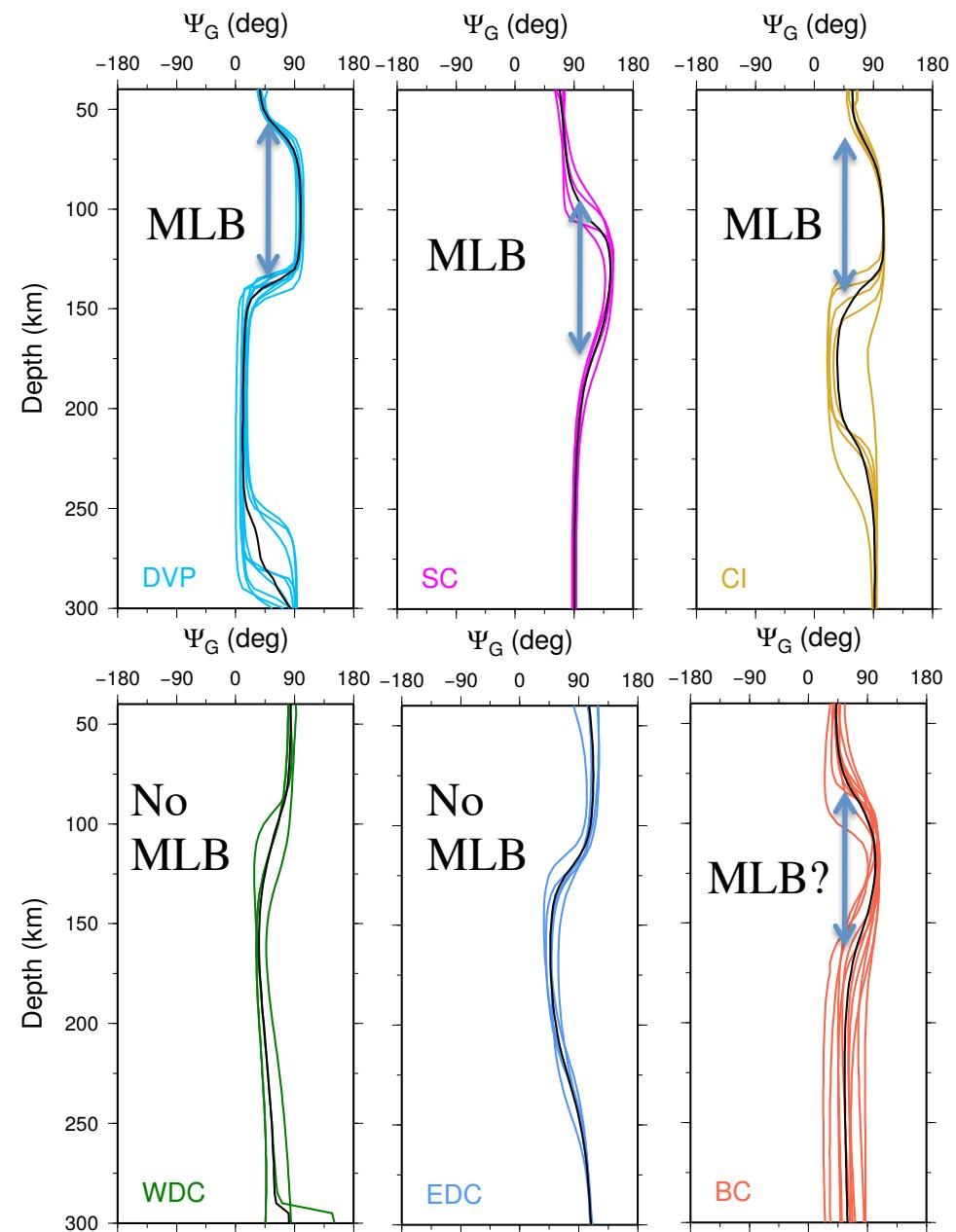
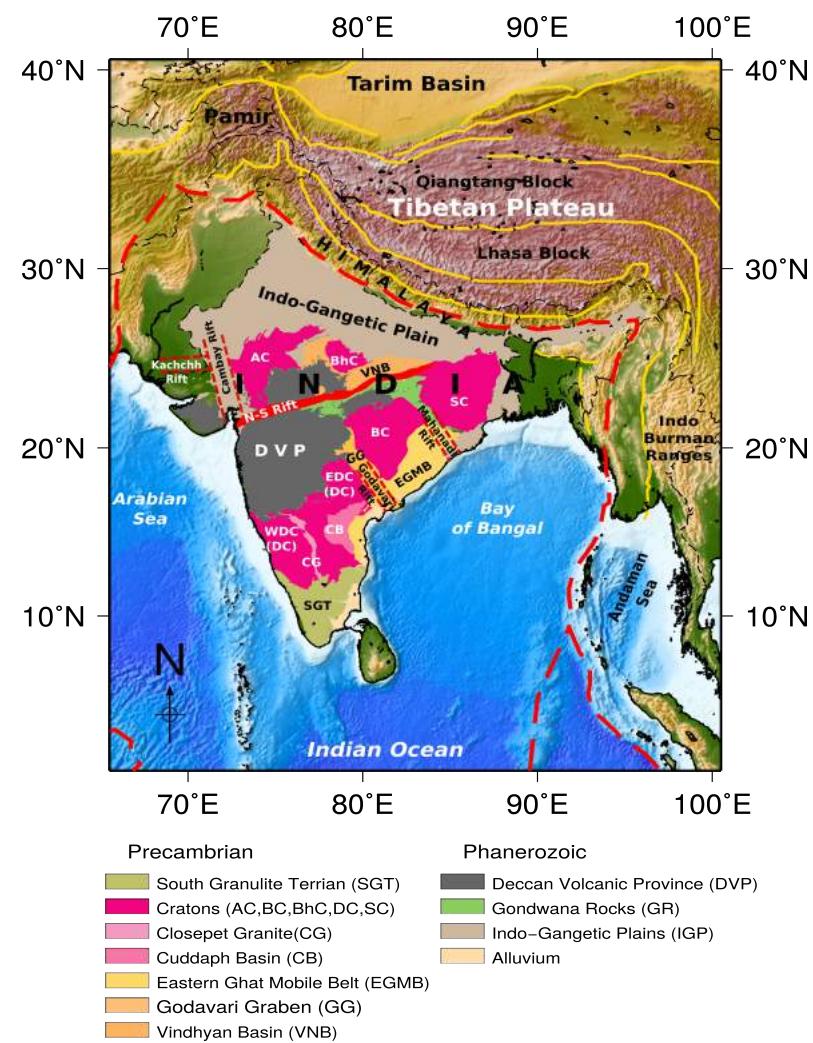
MLB: Mid-lithospheric low velocity zone

Discussion: Central India (CI)

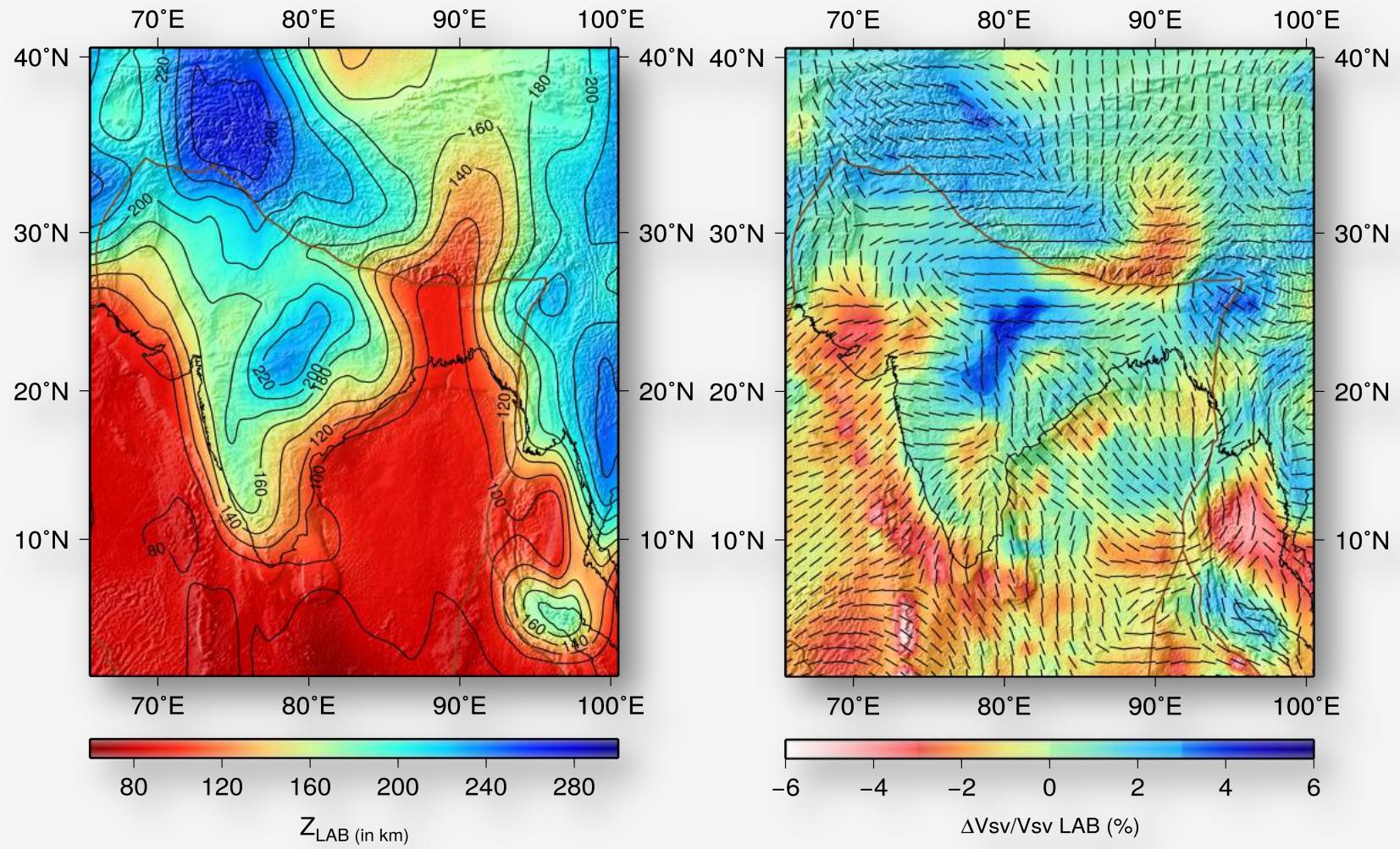
- Small velocity drop beneath DVP.
- Good agreement in both parameters ($V\downarrow SV, \psi\downarrow G$)
- Slow velocity present in the lithosphere.





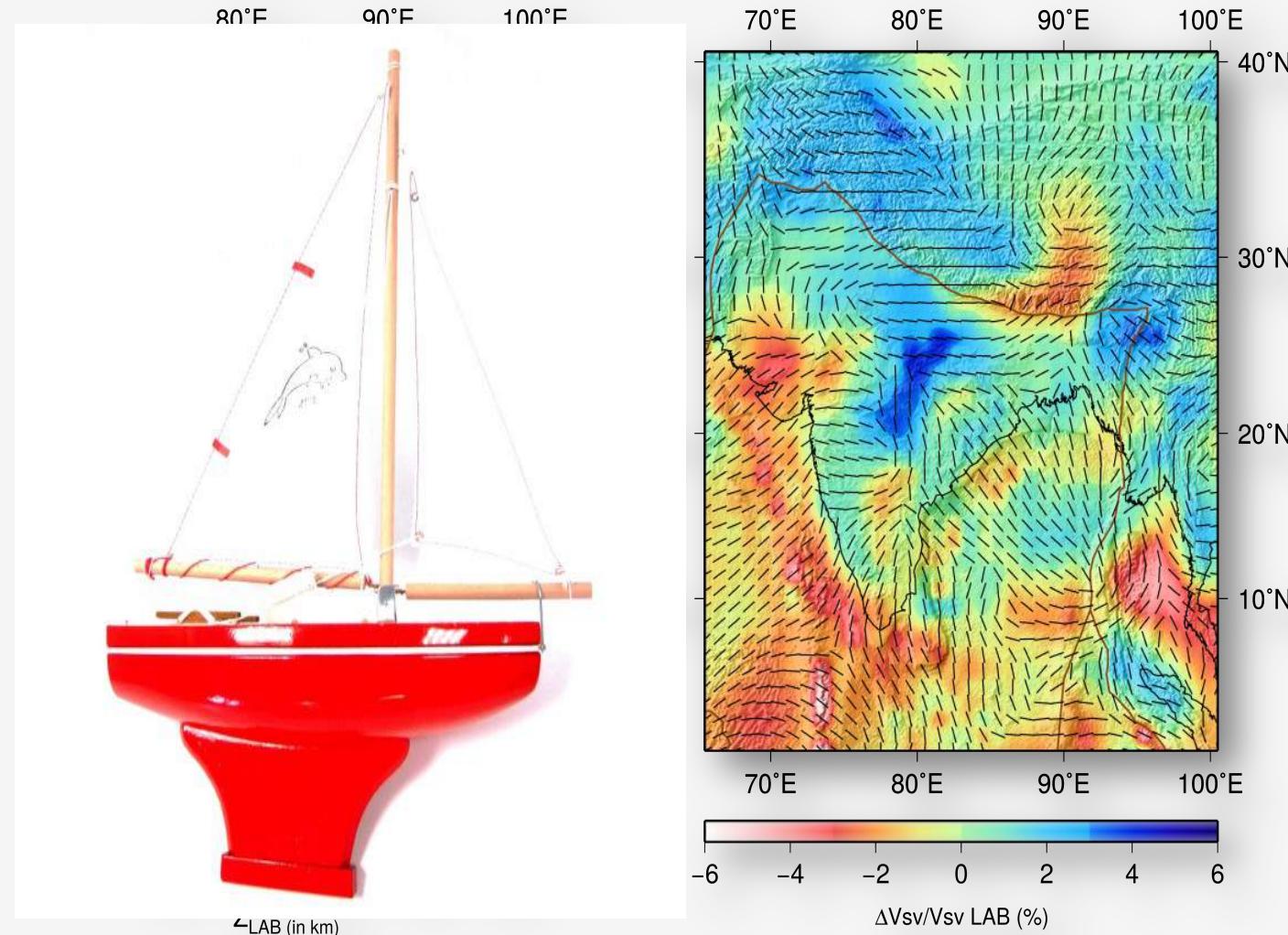


Indian Plate LAB (Lithosphere-Asthenosphere Boundary)



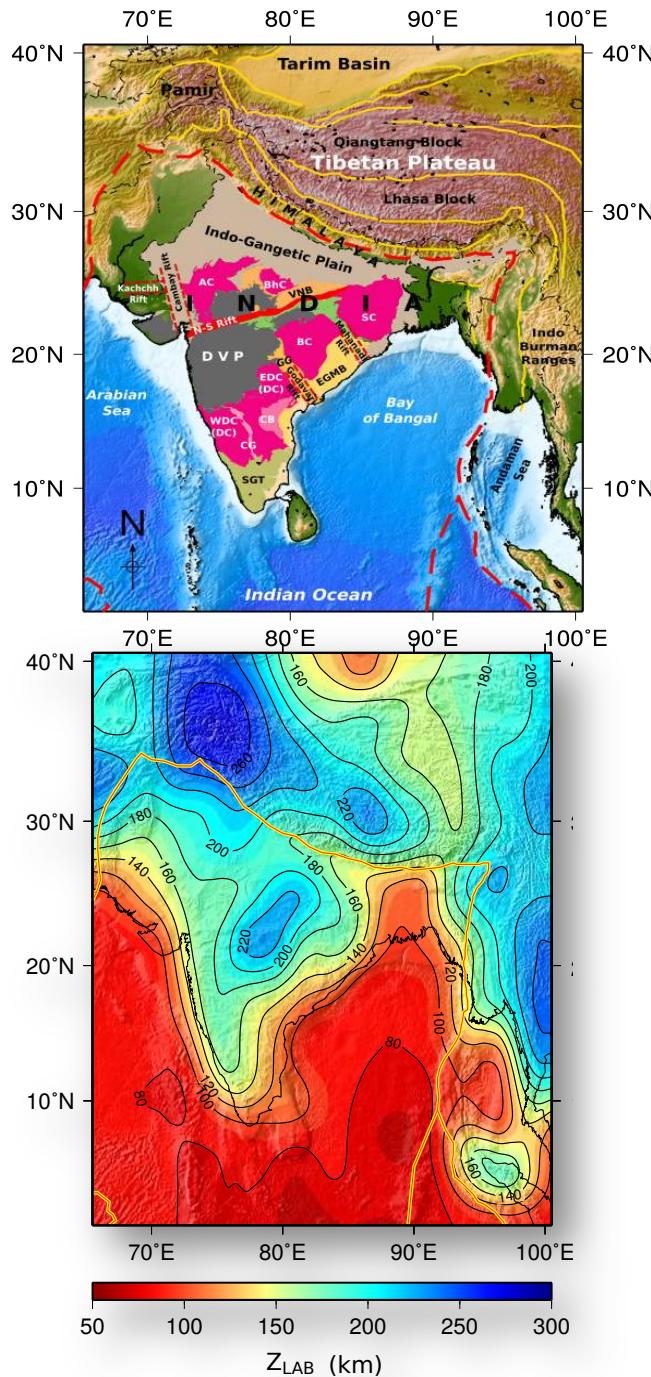
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Indian Plate LAB: Keel



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Geodynamic Role?
Plume influence?



- Large variability of craton thicknesses
- MLB (ML-LVZ): low velocity zone
- MLB: Change in azimuthal anisotropy
- MLB not present in all blocks

- DVP (Deccan Volcanic Province)
MLB: memory of La Réunion Hotspot birth

- Indian Keel: geodynamic role?

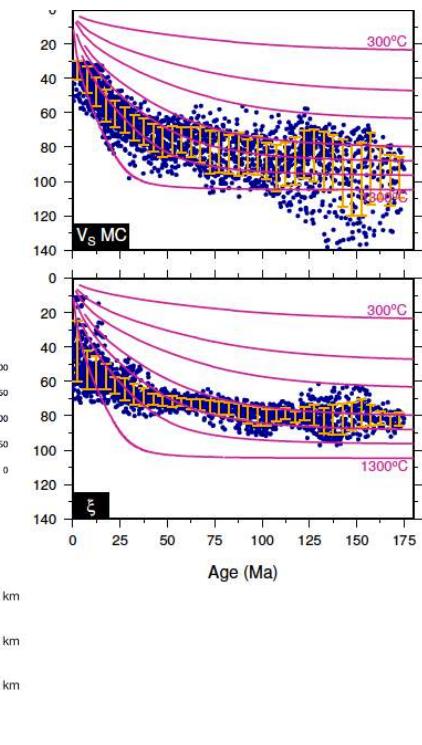
- Continents much more complicated than oceans

Conclusions

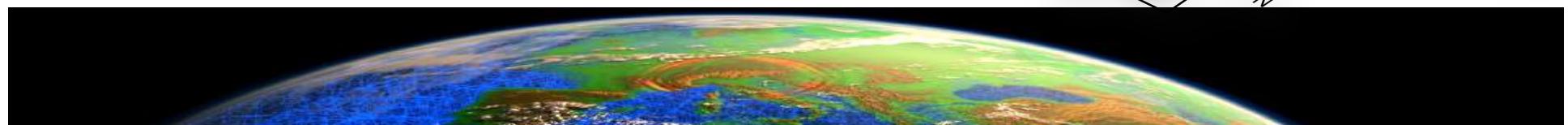
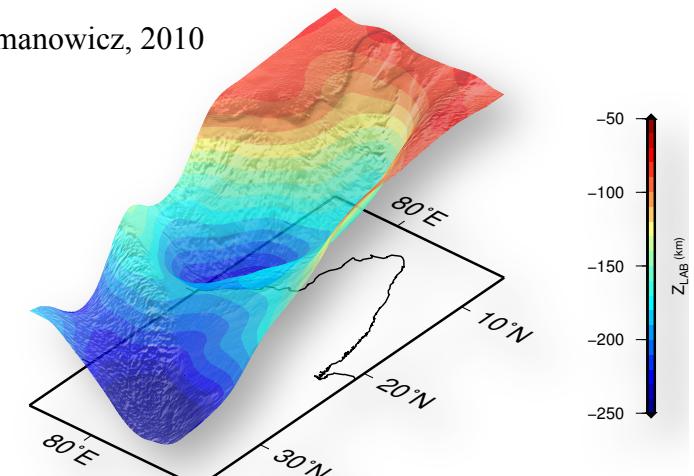
- Seismic Anisotropy can be mapped in different depth ranges
- Interpretation of seismic anisotropy is non-unique (intrinsic L.P.O. versus extrinsic anisotropy)
- Imaging of geological objects such as LAB (Lithosphere-Asthenosphere Boundary)
- New findings from anisotropy :

Oceans: MLB

Continents: variable thickness, Mid-LVZ
Keel

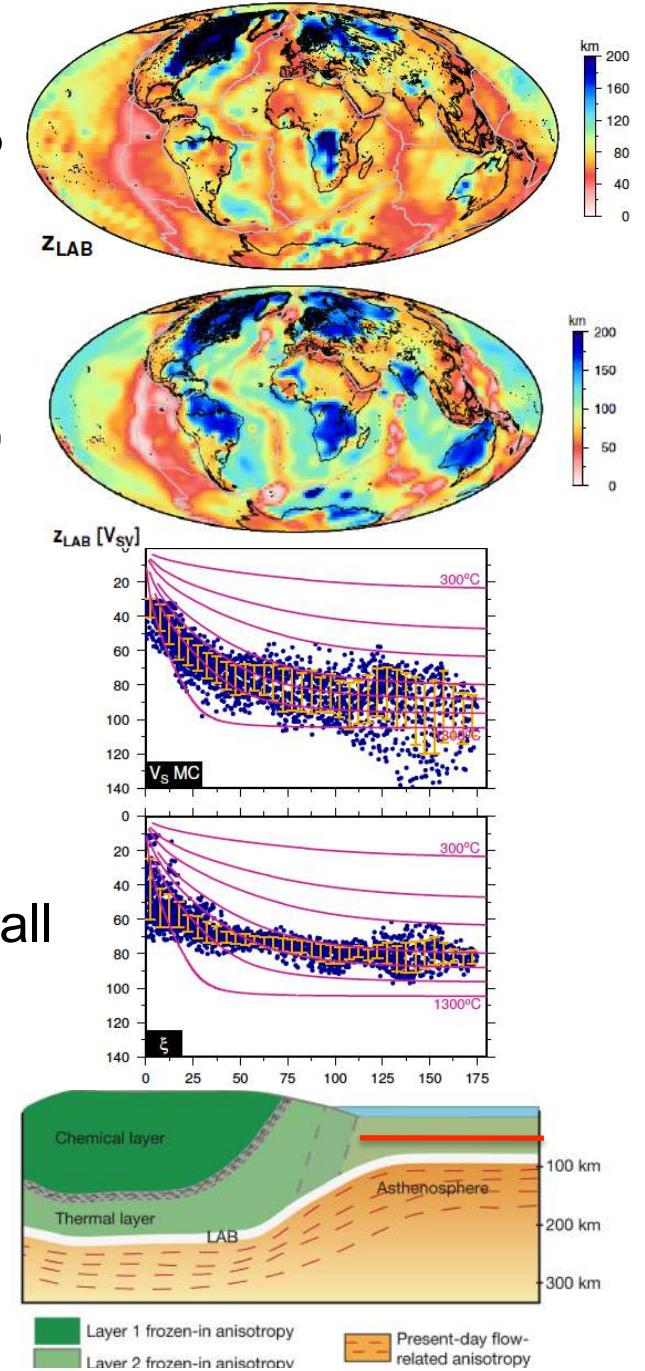


Yuan and Romanowicz, 2010



Oceanic and Continental plates

- LAB topography derived from surface wave data on a global scale (for oceans) and regional scale (India)
- For oceans: The ocean lithosphere not so simple! Stratification, existence of a Mid-Lithospheric Discontinuity (MLD: strong gradient of ξ between 60-80km)
- For continents: Large variability of craton thickness Stratification: MLB, low velocity zone (not present in all blocks). Relationship with MLD?
- The model of formation of lithosphere must be revisited in view of results from radial and azimuthal anisotropies in oceans and continents.



Yuan and Romanowicz, 2010

3-D Indian LAB model

