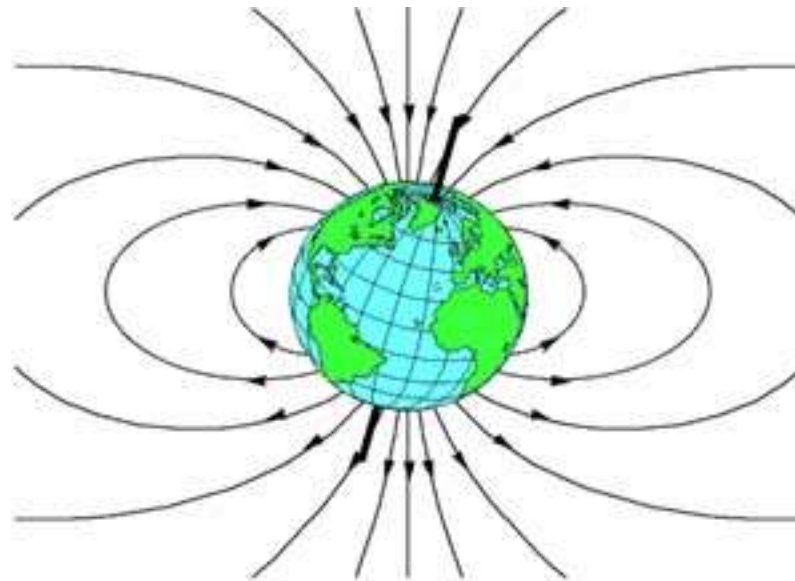


# Dynamics of Earth's Interior and Geodynamo



Structure and Dynamics of Earth-like Planets, Collège de France

Bruce Buffett, Earth & Planetary Science, UC Berkeley



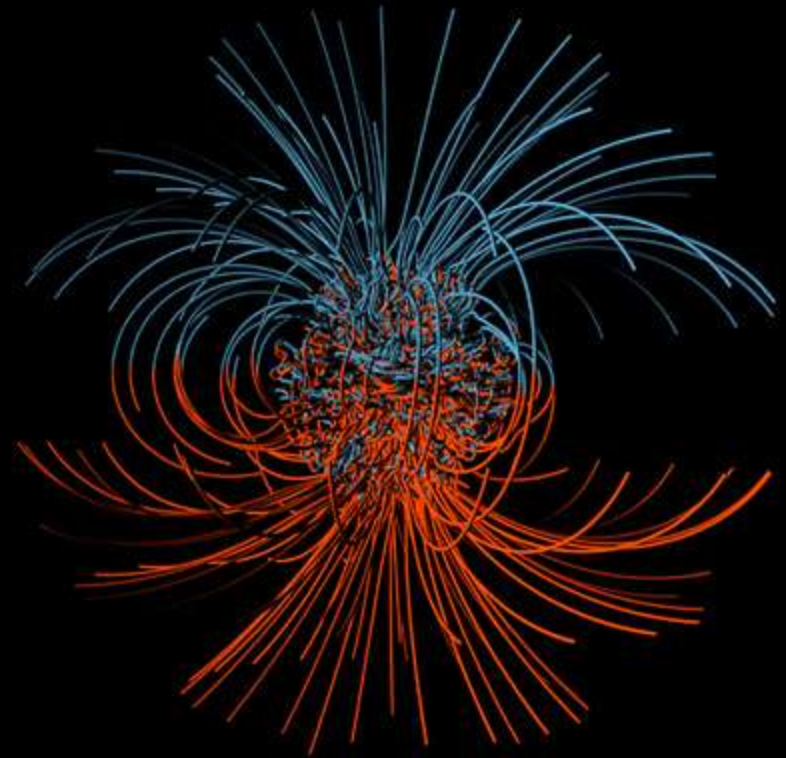
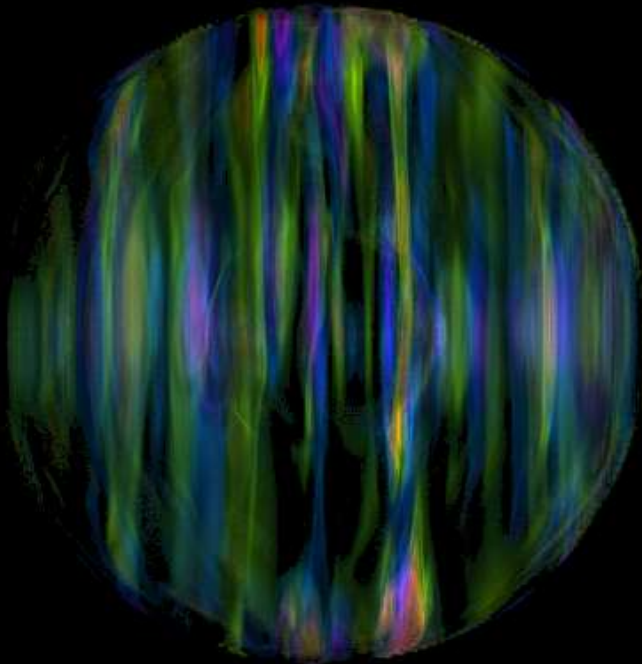
# Terrestrial Planets



# Gas and Ice Giants

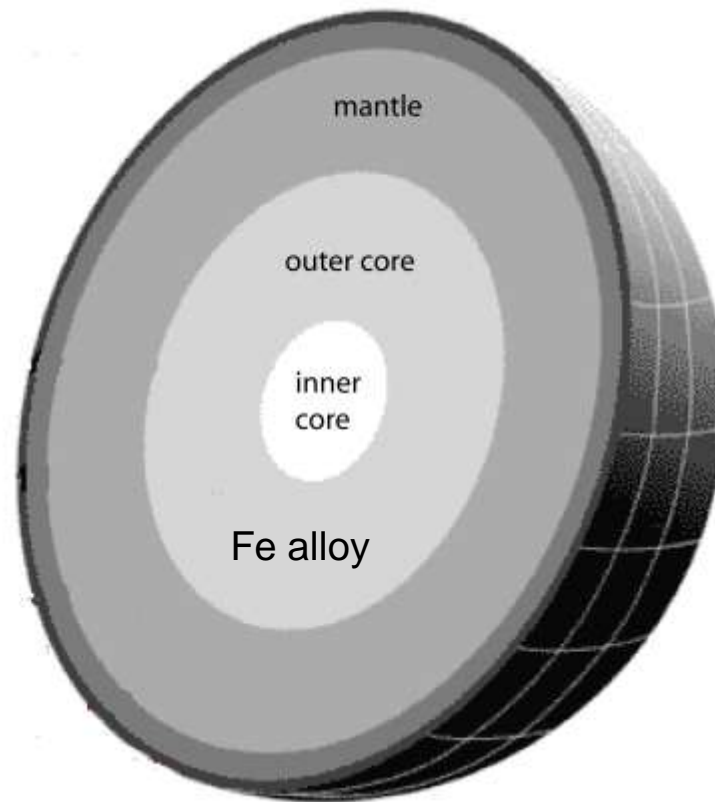


# Generation Mechanism

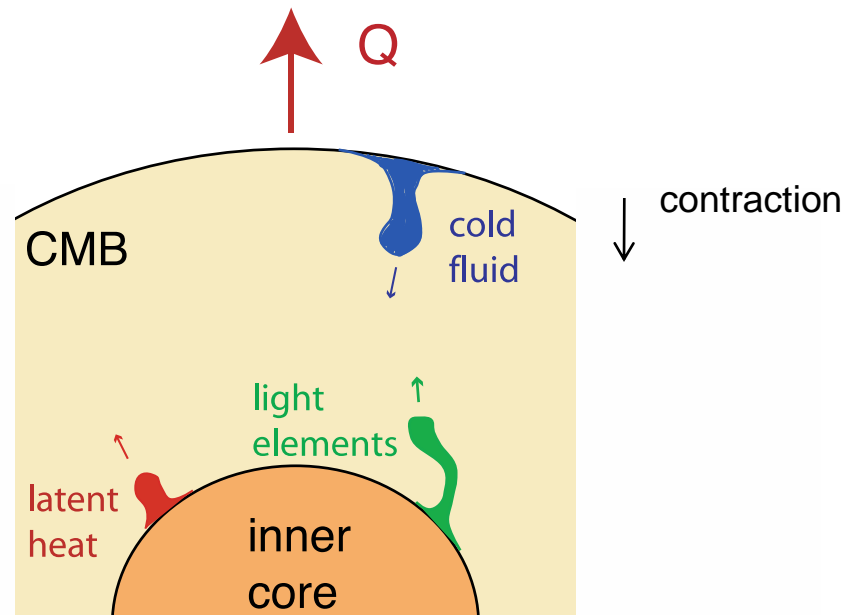


fluid motions in an electrical conductor generate the magnetic field

# Internal Structure

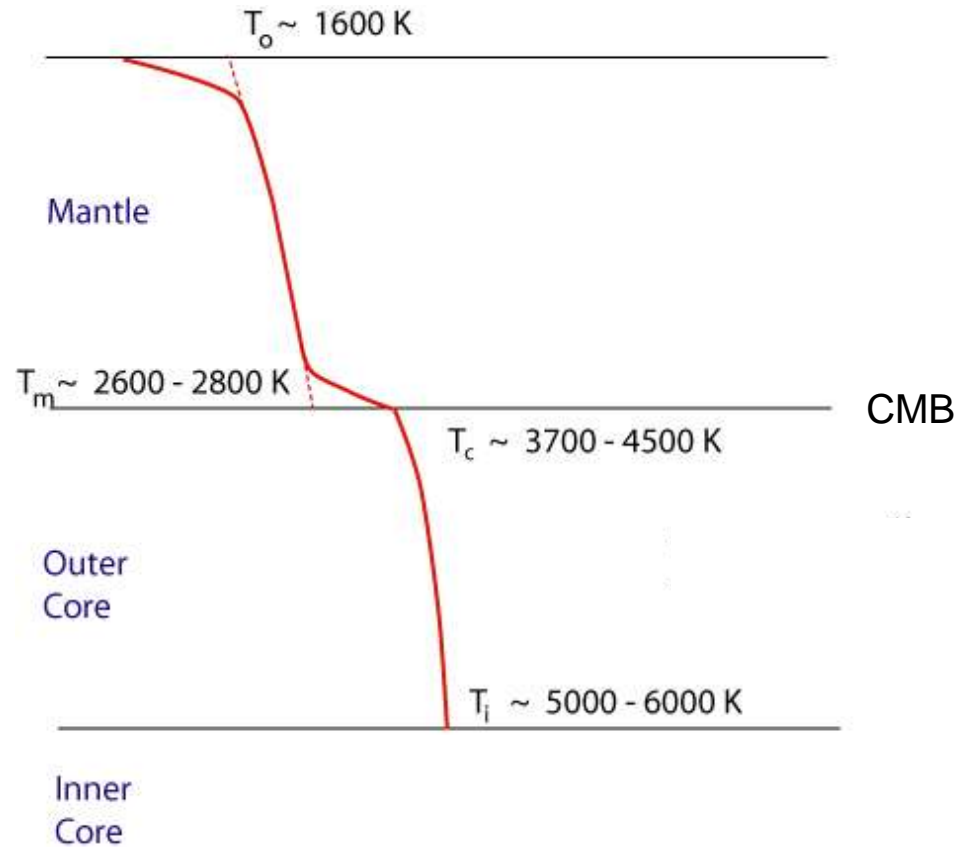


# Physical Processes



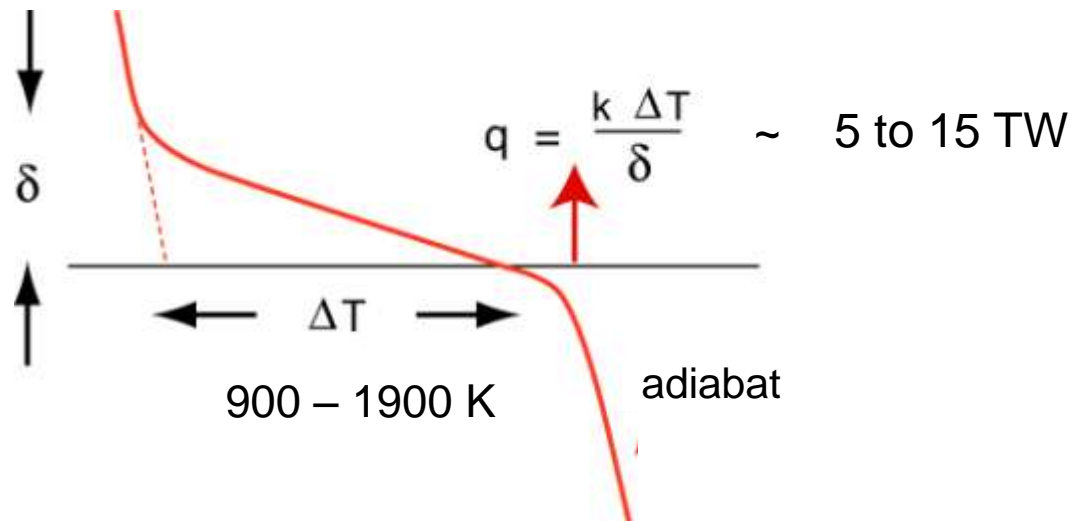
Cooling of the core is controlled by mantle convection

# Present-Day Temperature



temperature drop across boundary layer:  $\Delta T = 900 - 1900 \text{ K}$

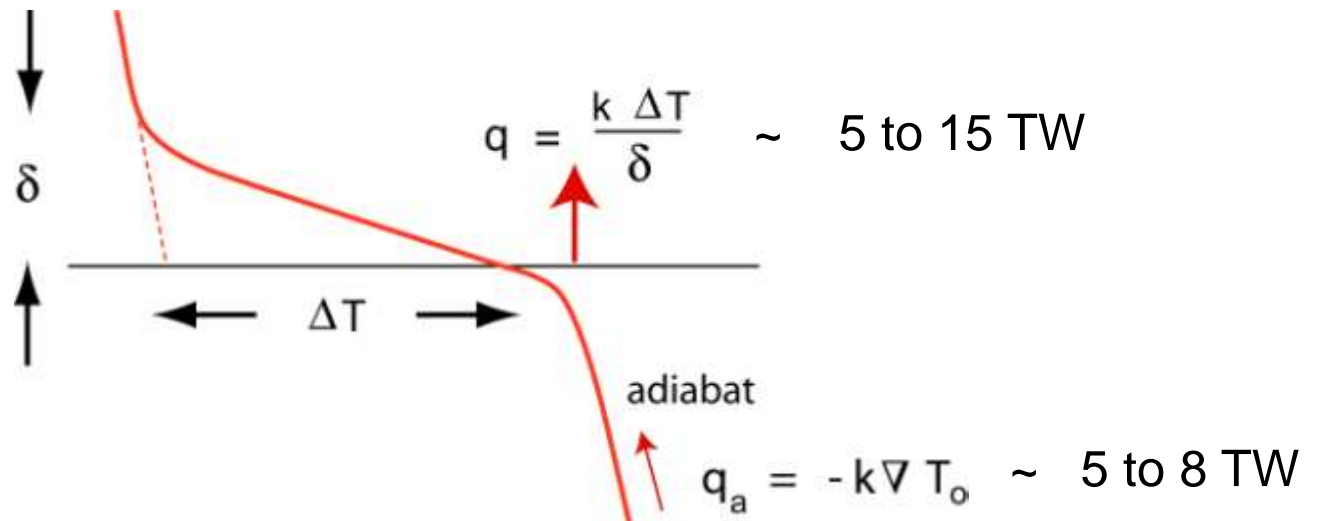
# Core Heat Flow



thermal boundary layer on the core side?

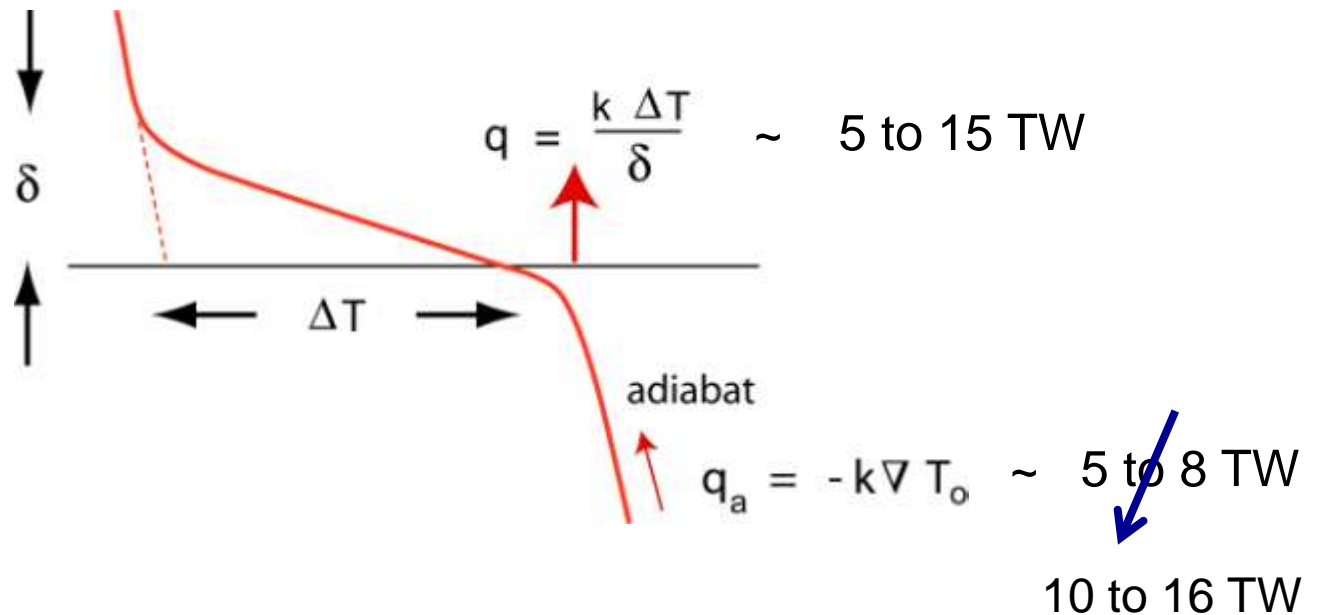


# Core Heat Flow



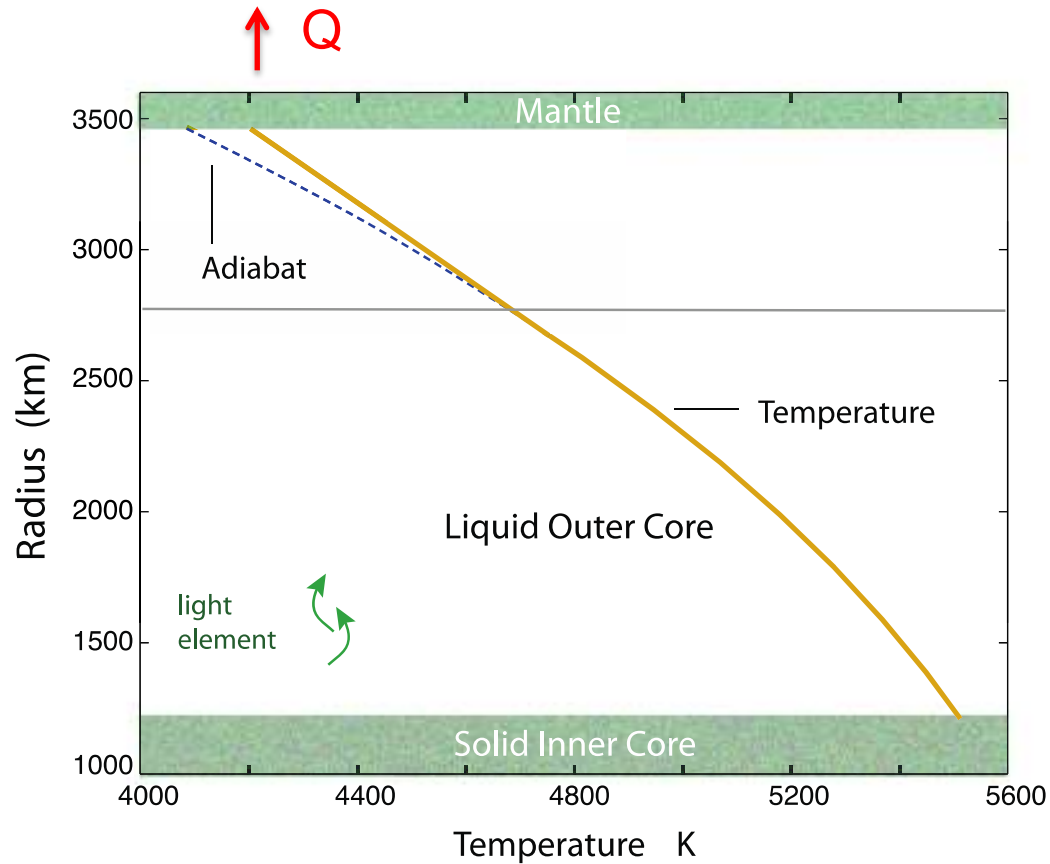
conduction along adiabat is comparable to total heat flow

# Core Heat Flow



conduction along adiabat is comparable to total heat flow

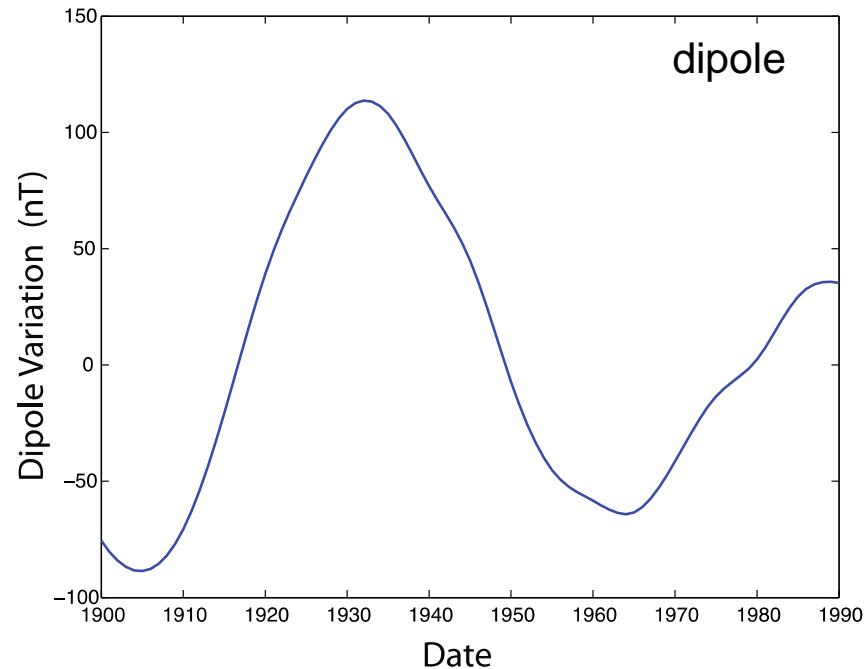
# Thermal Stratification



stratification supports waves with periods of several decades

# Long-Period Fluctuations

A dominant 60-year period in magnetic field

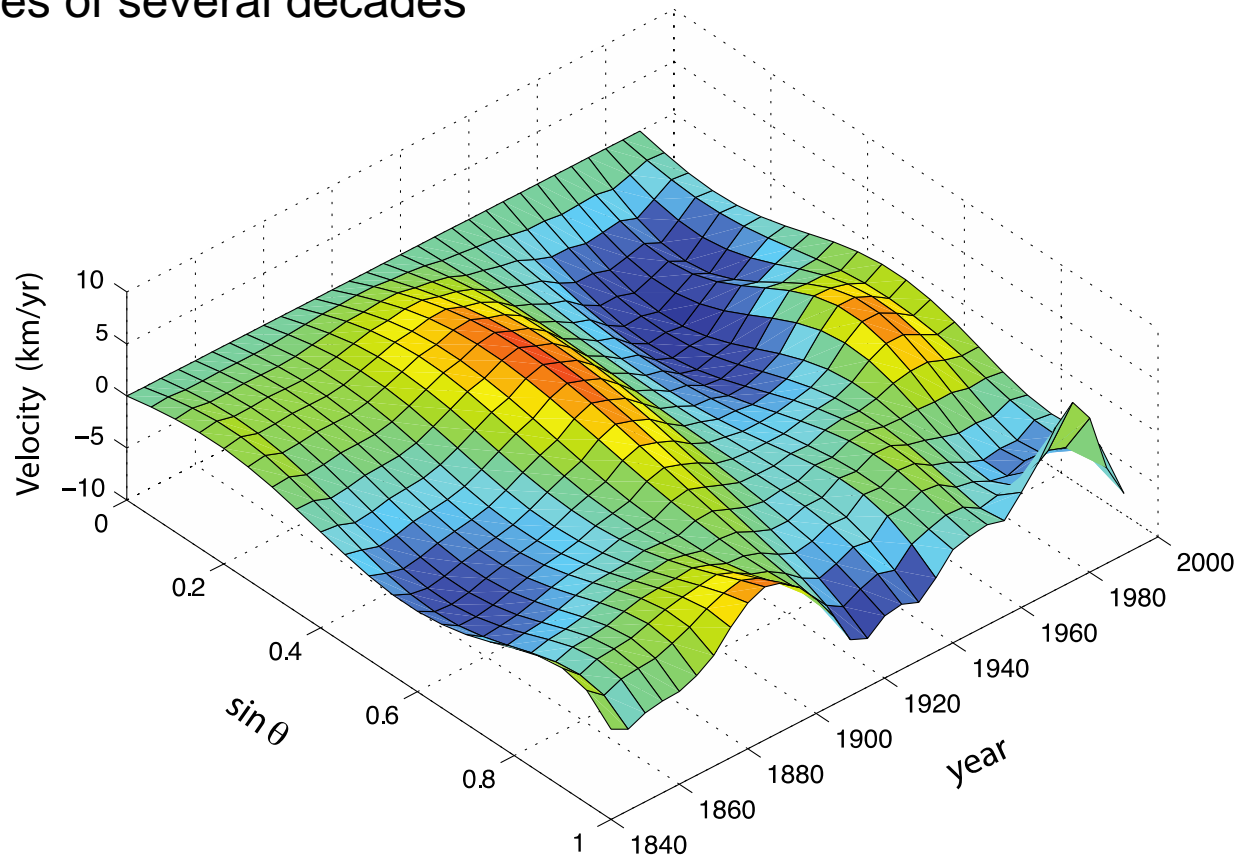
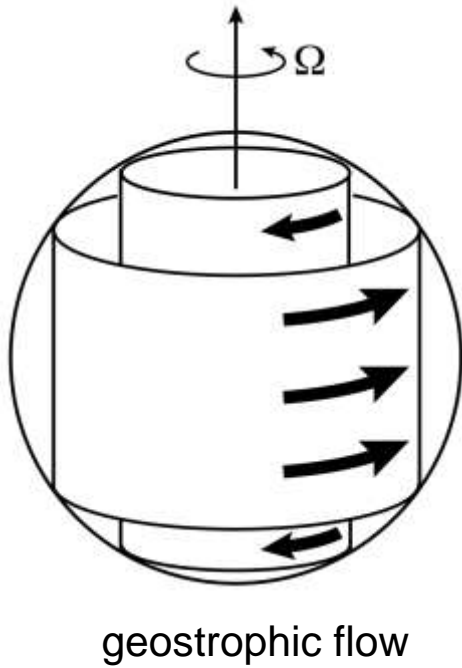


60-year signal also detected in magnetic declination

(e.g. Roberts et al. 2007)

# Surface Core Flow

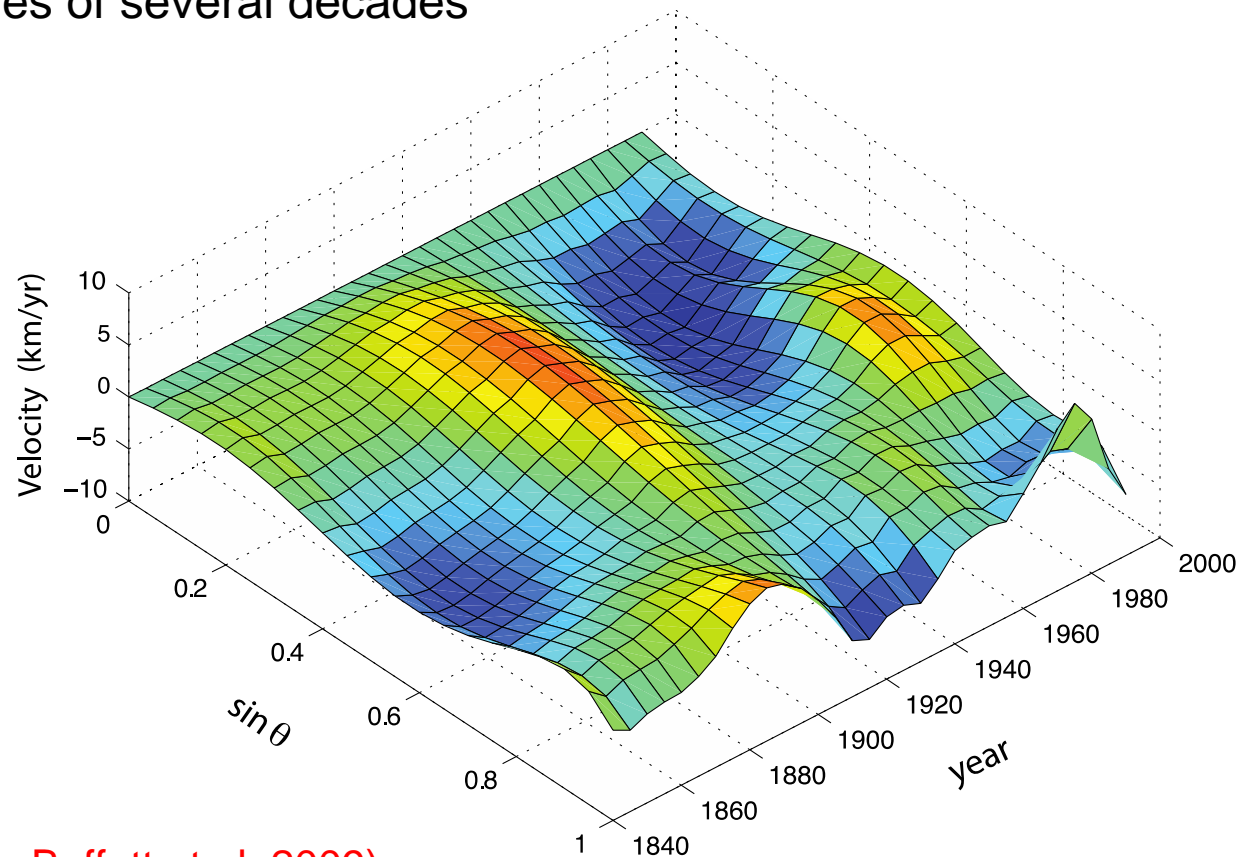
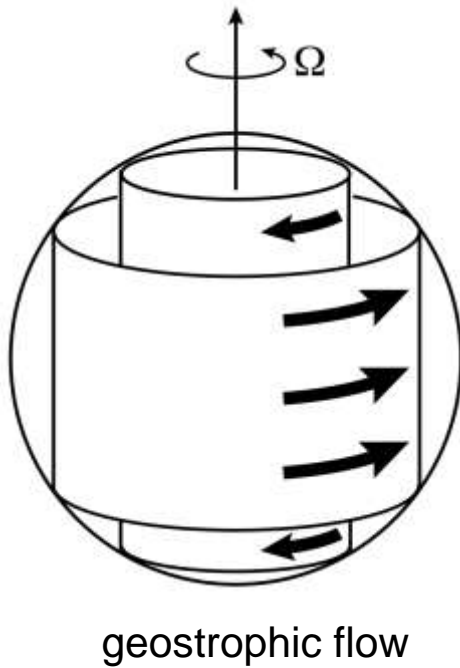
Fluctuations on timescales of several decades



Jackson (1997)

# Time-Dependent Flow

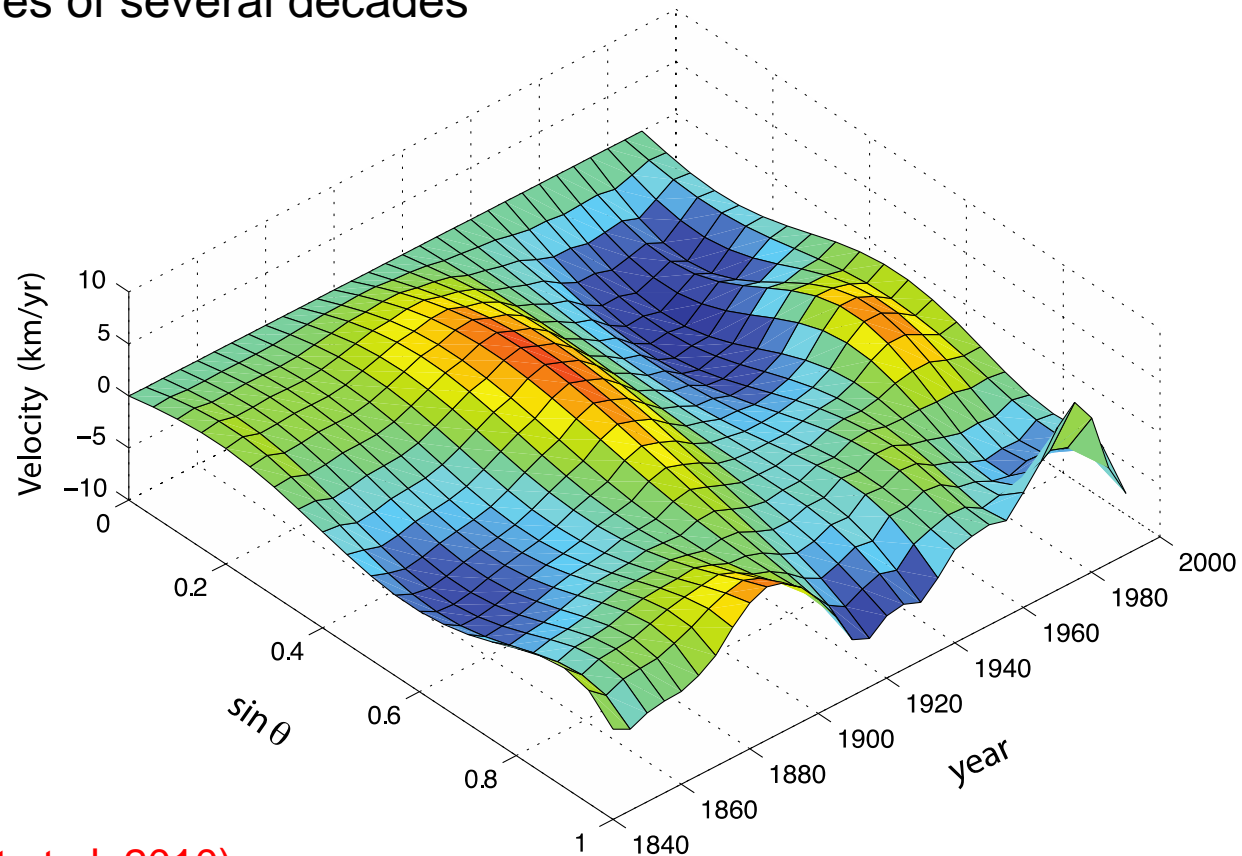
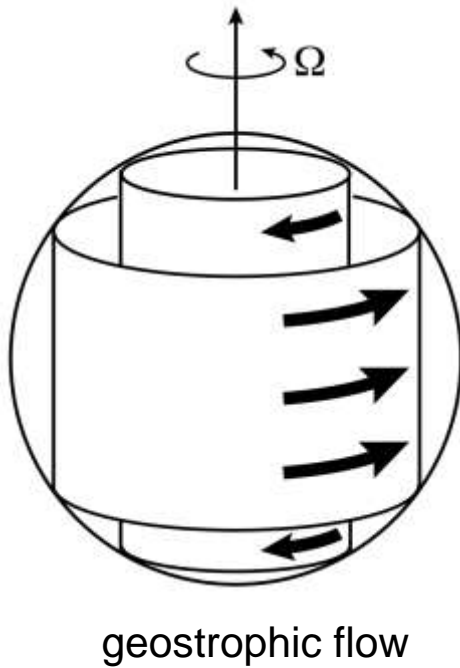
Fluctuations on timescales of several decades



→  $B \sim 0.3 \text{ mT}$  (e.g. Buffett et al. 2009)

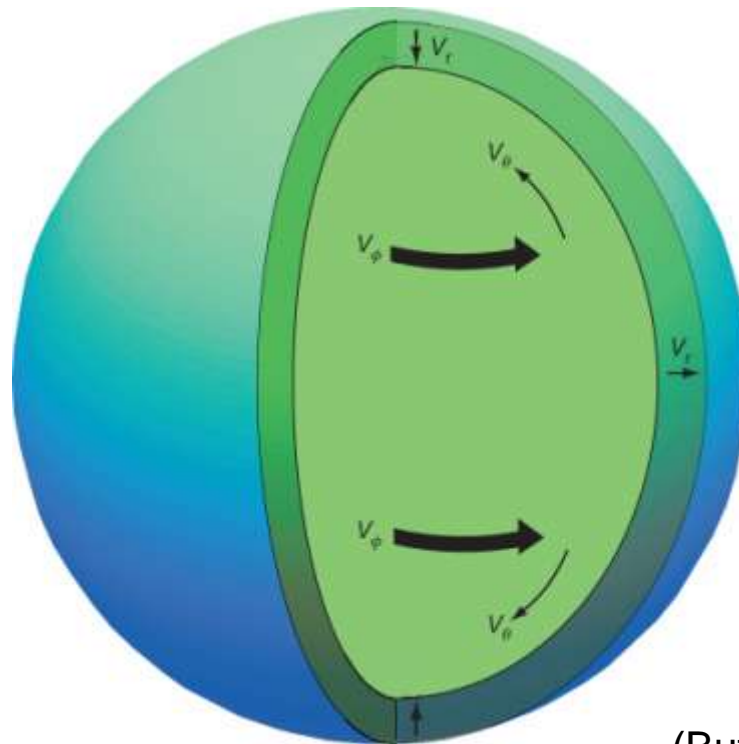
# Time-Dependent Flow

Fluctuations on timescales of several decades



  $B > 2 \text{ mT}$  (Gillet et al. 2010)

# MAC Waves

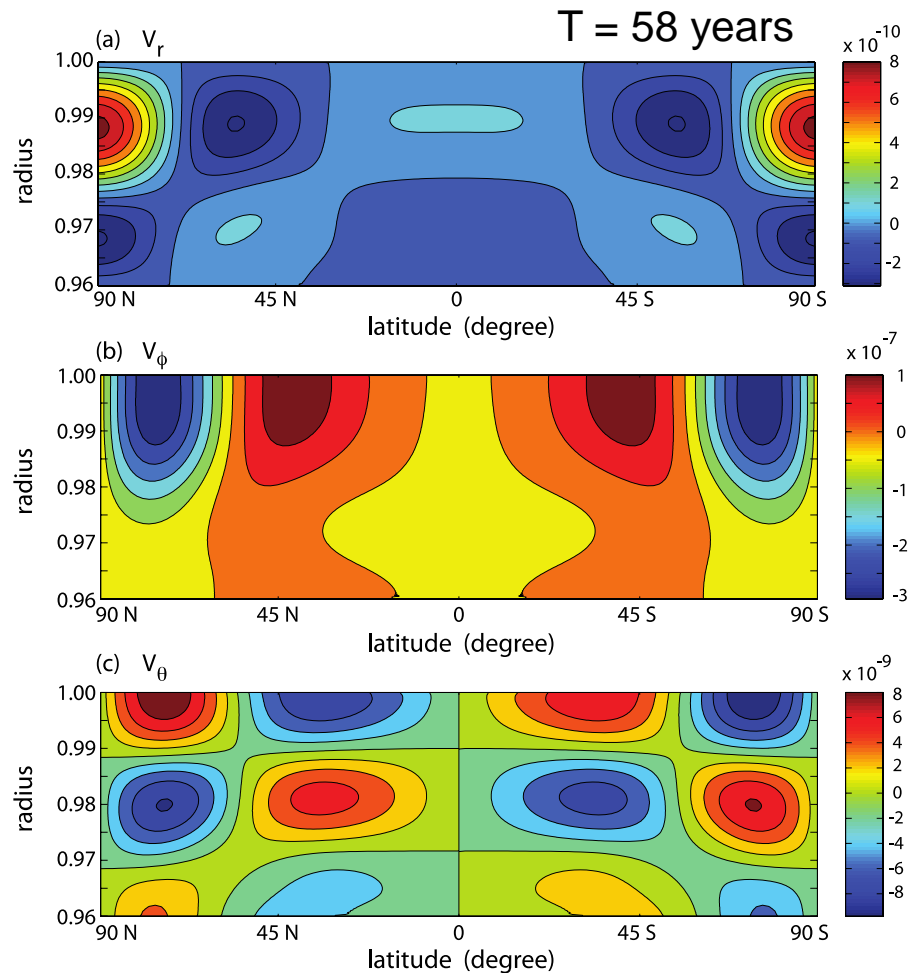


(Buffett, 2014)

A combination of Magnetic, Archimedes and Coriolis forces

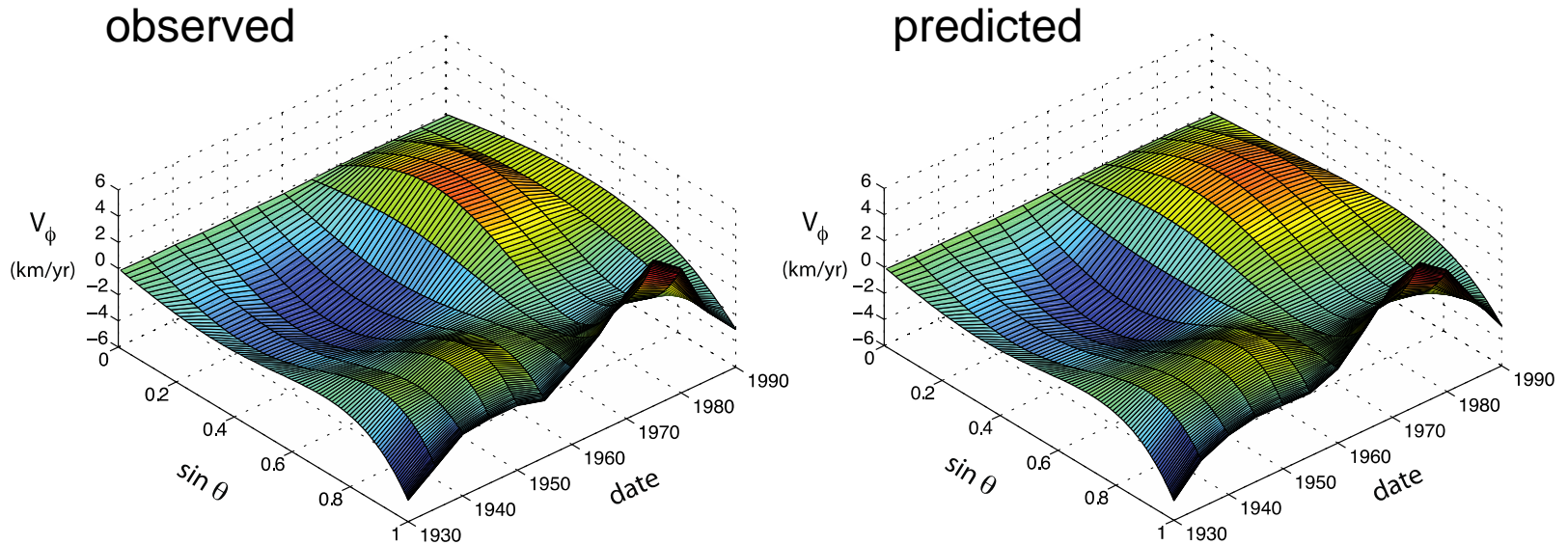


# Representative Wave



model parameters: thickness  $H$  and buoyancy frequency  $N$

# Fit to Observations



Vary  $H$  and  $N_{\max}$  to fit estimates of  $v_\phi$  using linear combination of 6 modes

best-fitting model parameters:  $H \sim 140$  km       $N_{\max} \sim \Omega$

# Consequences of Zonal Flow

Simple scaling suggests

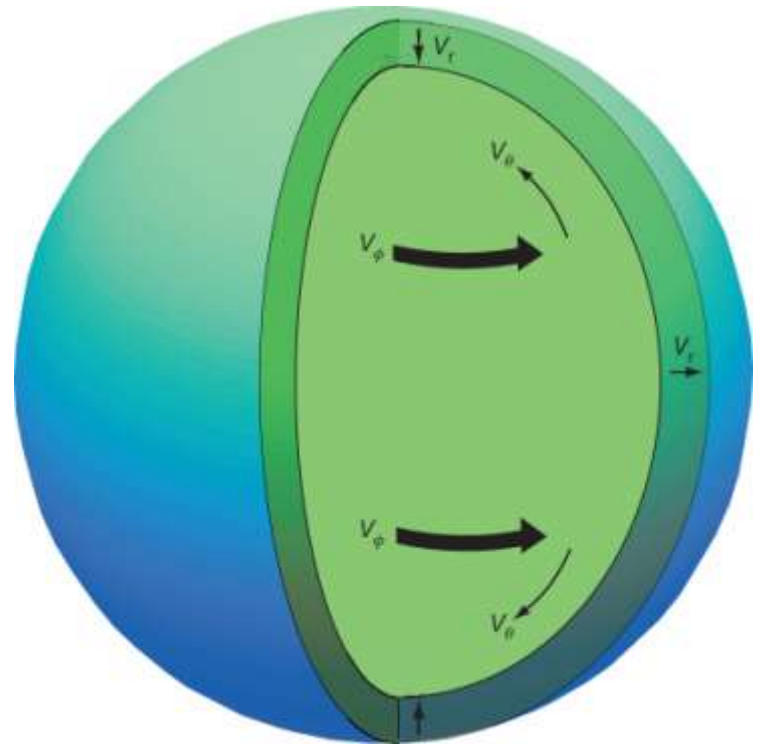
$$\frac{v_\theta}{v_\phi} \approx \frac{V_a^2 \pi^2}{H^2 \Omega \omega} \approx 0.08$$

where  $V_a = B_r / \sqrt{\mu \rho}$

Secular Variation of Magnetic Field

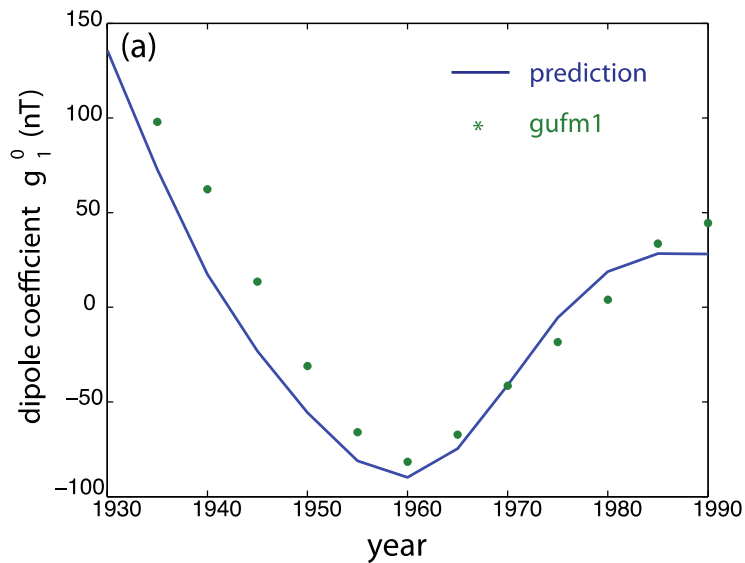
$$\partial_t b_r = -\nabla_H \cdot (\mathbf{v} B_r)$$

use waves to predict dipole fluctuations

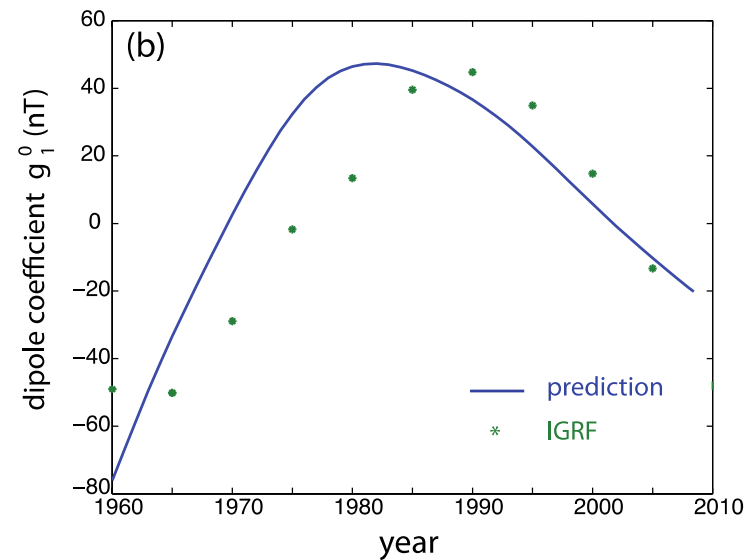


# Dipole Fluctuations

Fit to  $V_\phi$  of Jackson (1997)



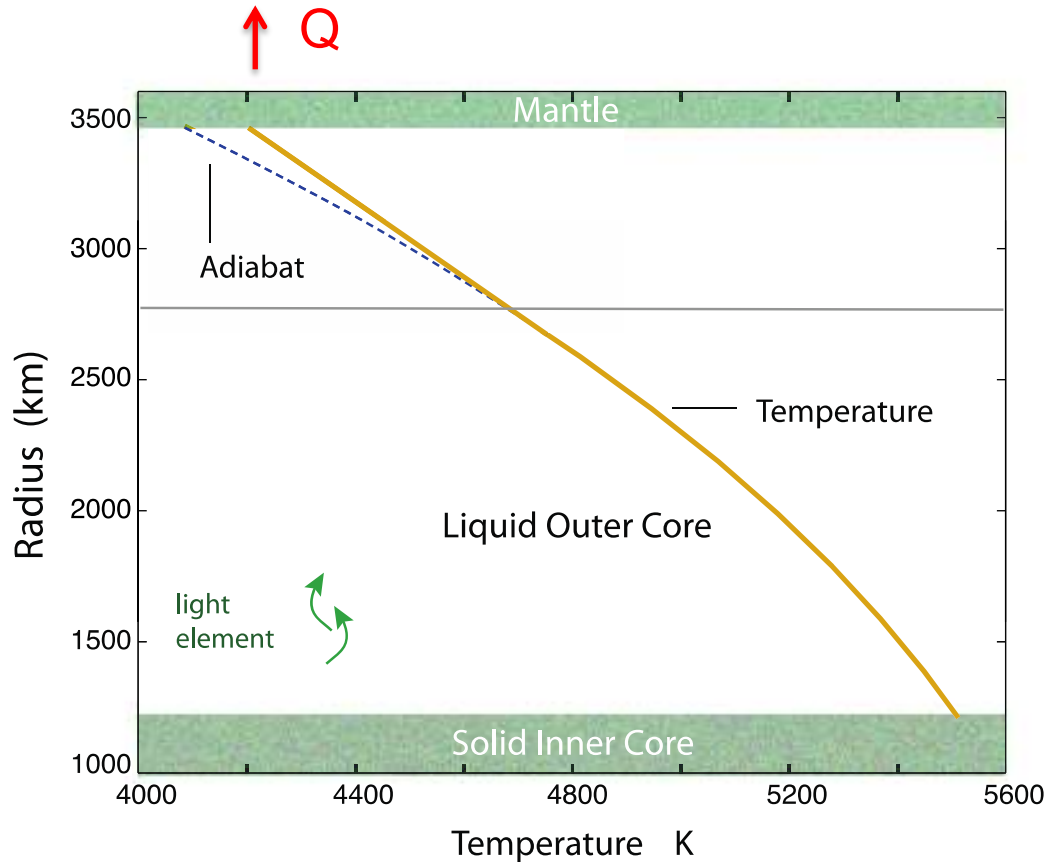
Fit to  $V_\phi$  of Wardinski and Lesur (2013)



Good fit to flow and dipole fluctuation requires  $\sigma \sim 10^6 \text{ S m}^{-1}$

$$\frac{\mathbf{v}_\theta}{\mathbf{v}_\phi} \approx \frac{V_a^2 \pi^2}{H^2 \Omega \omega} \approx 0.08$$

# Thermal Stratification



Model of Lister & Buffett (1998)

$$Q = 13 \text{ TW}$$

$$Q_a = 15 \text{ TW}$$

yields

$$H = 140 \text{ km}$$

$$N = 1.5 \Omega$$

(waves suggest  $H = 140 \text{ km}$  and  $N = 1.0 \Omega$ )

# Conclusions

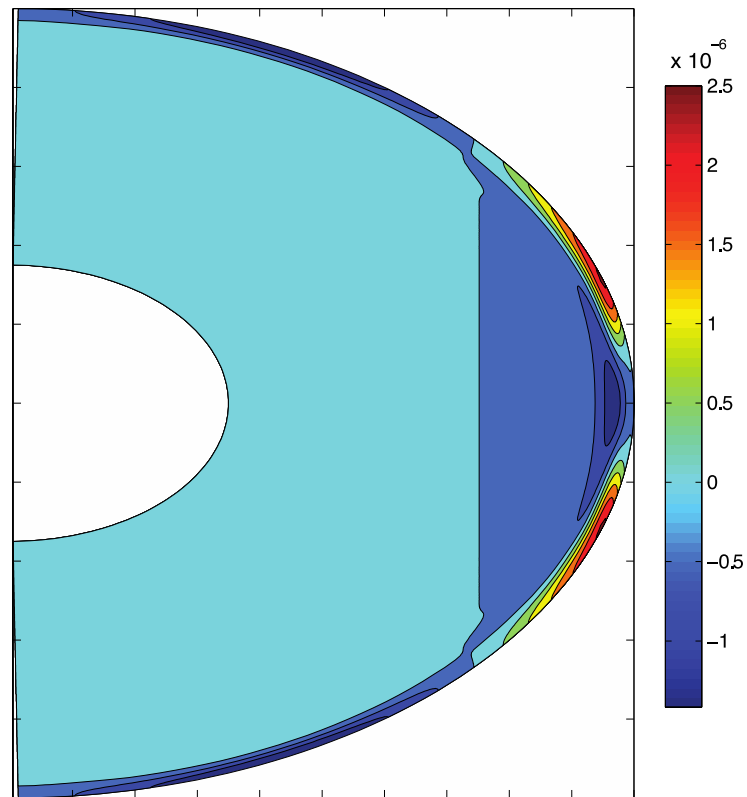
1. Energy for geodynamo depends on core cooling
  - thermal and chemical buoyancy today
  - thermal buoyancy prior to 500 Ma ?
  - existence of field at 3.4 Ga requires plate tectonics?
  
2. Field variations provide important constraints on dynamics
  - evidence for MAC waves at top of core
  - mantle heat flow is approximately  $Q = 13$  TW
  - higher  $Q$  is required to sustain field before 500 Ma

# Conclusions

3. Wave motion offers a new probe of core dynamics
  - non-zonal modes contribute to field variations
  - how are the modes excited ?
  - coupling to deeper flow required to explain LOD

# Flow in Interior

Azimuthal velocity  $v_\phi$



wave motion in stratified layer couples to geostrophic flow in interior