



COLLÈGE
DE FRANCE
—1530—

CHAIRE DE PARTICULES ÉLÉMENTAIRES, GRAVITATION ET COSMOLOGIE

Année académique 2012-2013

M. Gabriele Veneziano, Professeur

Colloque de clôture : Un boson nommé Higgs

Vendredi 24 mai 2013 à 9h00

Implications et perspectives théoriques



Riccardo Barbieri
SNS and INFN, Pisa

Particle Physics in one page

\mathcal{L}_{ST} \leftarrow \mathcal{L}_{SM}

$$\mathcal{L}_{SM} = -\frac{1}{4}F_{\mu\nu}^a F^{a\mu\nu} + i\bar{\psi} \not{D}\psi$$

The gauge sector 1

$$+ |D_\mu h|^2 - V(h)$$

The EWSB sector 2

$$+ \psi_i \lambda_{ij} \psi_j h + h.c.$$

The flavour sector 3

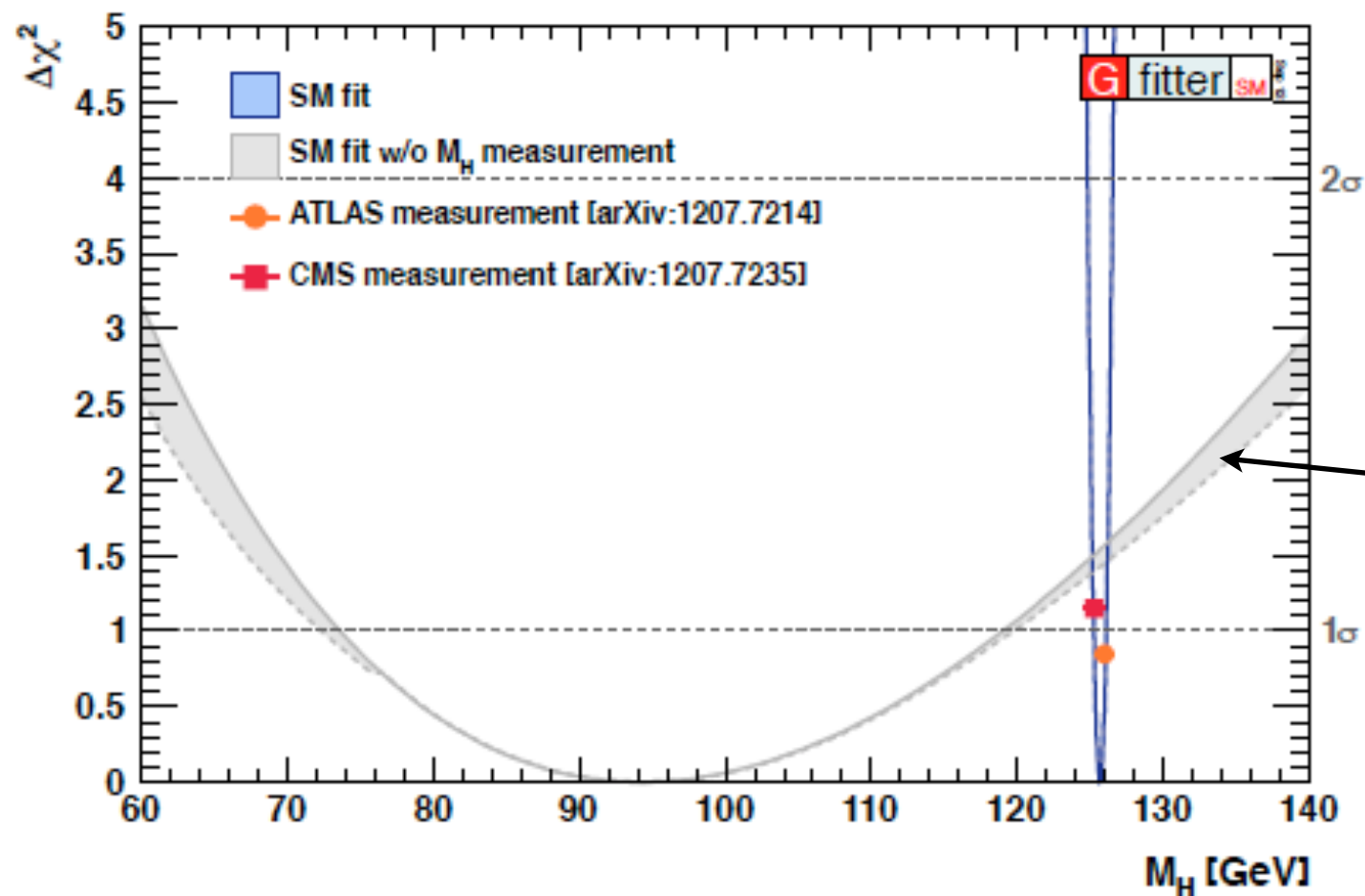
$$+ N_i M_{ij} N_j$$

The ν -mass sector 4
(if Majorana)

+
Dark Matter ✓
Baryon Asymmetry ✓
Dark Energy

(not included in current knowledge of particle physics)

The great empirical evidence for (and from) the gauge sector (for extension, precision, diversity)



the χ^2 distribution of m_h
w/o the m_h measurements

(with a similar story for the top discovery in 1994)
(and a partially similar story for the W-Z in 1884)

Is it the coronation of the SM or a step on a road still largely unexplored?

$$\mathcal{L}_{ST} = |D_\mu h|^2 - m^2 h^2 - \lambda h^4 + \lambda_{ij} \Psi_i \Psi_j h (+\Lambda^4)$$

how natural?

which dynamics, if any?

how about the flavour puzzle?

(Note: no physical inconsistency!)

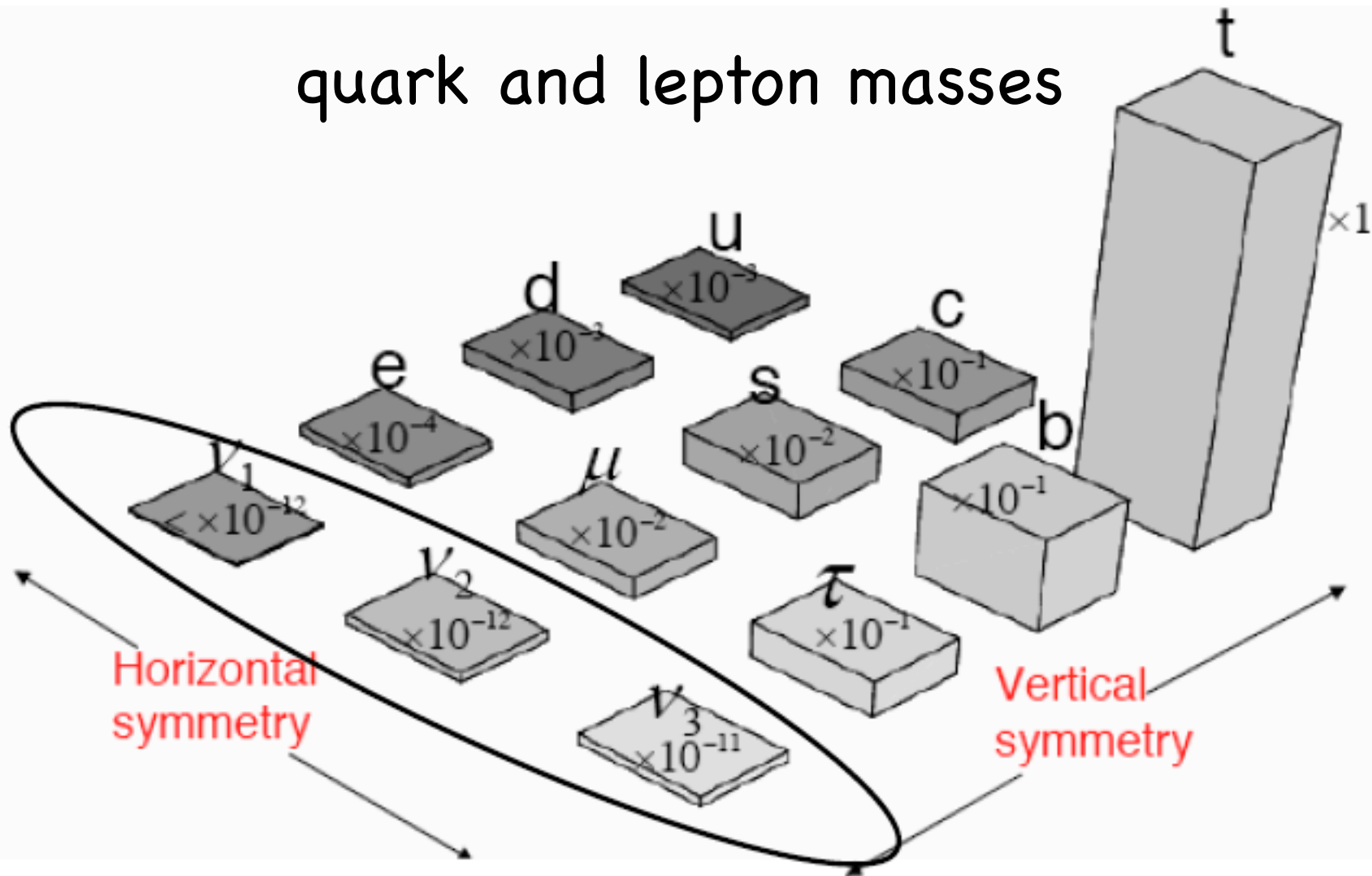
A paradoxical answer: yes to both alternatives

The flavour puzzle

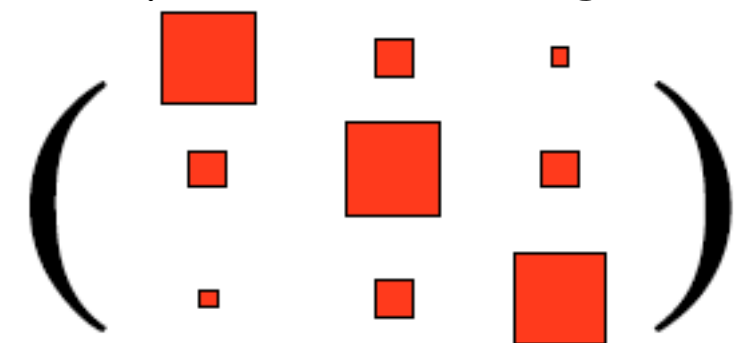
$$\lambda_{ij} \Psi_i \Psi_j h$$

(not touched in this lecture)

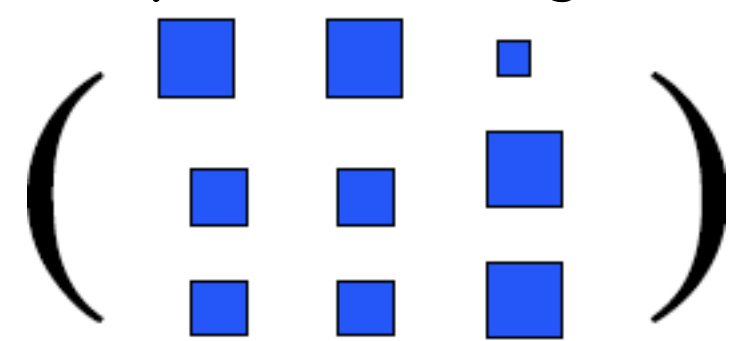
quark and lepton masses



quark mixings



lepton mixings



Every element in these pictures accounted for by an *ad hoc* parameter among the λ_{ij}

What determines this structure?

Not easy without observed deviations from the SM

About naturalness

a dominant paradigm in the last thirty years

naturalness 1:

$$m_{Pl} = (\hbar c / G_N)^{1/2} \approx 10^{19} \text{ GeV}$$

$$l_{Pl} = \hbar / (m_{Pl} c) \approx 10^{-33} \text{ cm}$$

In the current field theory framework:

Why there is a large universe ($\Lambda \approx 10^{-3} \text{ eV} \ll m_{Pl}$)?

Why there are large objects in it ($m_h \ll m_{Pl}$)?

naturalness 2:

Can we do physics at different scales without knowing the details at shorter distances?

Atomic
physics

Nuclear
physics

EW
physics

?
physics

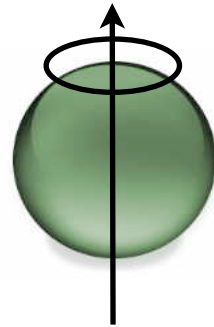
gravity

Apparently not at the moment!

naturalness 3:

Among the many examples that beautifully work so far

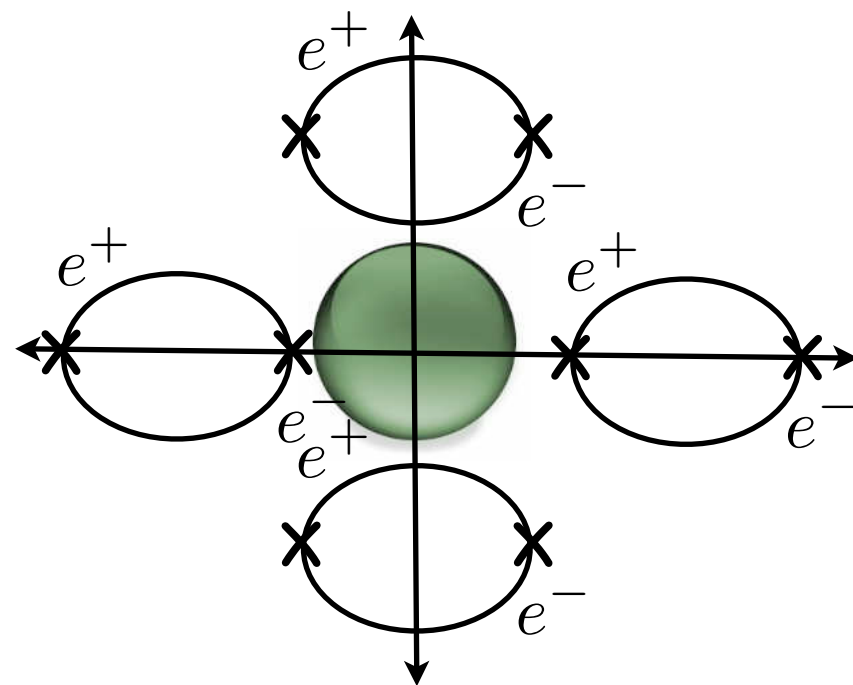
the electron self energy:



electric $E_{el} \approx \frac{e^2}{r_e} \lesssim m_e c^2 \Rightarrow \Lambda_e \equiv \frac{\hbar}{r_e c} \lesssim \frac{m_e}{\alpha} \approx 70 \text{ MeV}$

magnetic $E_{mag} \approx \frac{\mu^2}{r_e^3} \lesssim m_e c^2 \Rightarrow \Lambda_e \lesssim \frac{m_e}{\alpha^{1/3}} \approx 3 \text{ MeV} \quad (\mu = \frac{e\hbar}{2m_e c})$

the positron (a doubling of the d.o.f. at $\Lambda_e \sim m_e$) solves the problem



Weisskopf 1939

$$(M_{\pi^+}^2 - M_{\pi^0}^2 \Rightarrow m_\rho \lesssim 800 \text{ GeV})$$

$$(M_{K_L^0} - M_{K_S^0} \Rightarrow m_c \lesssim 2 \text{ GeV})$$

Back to the Higgs boson

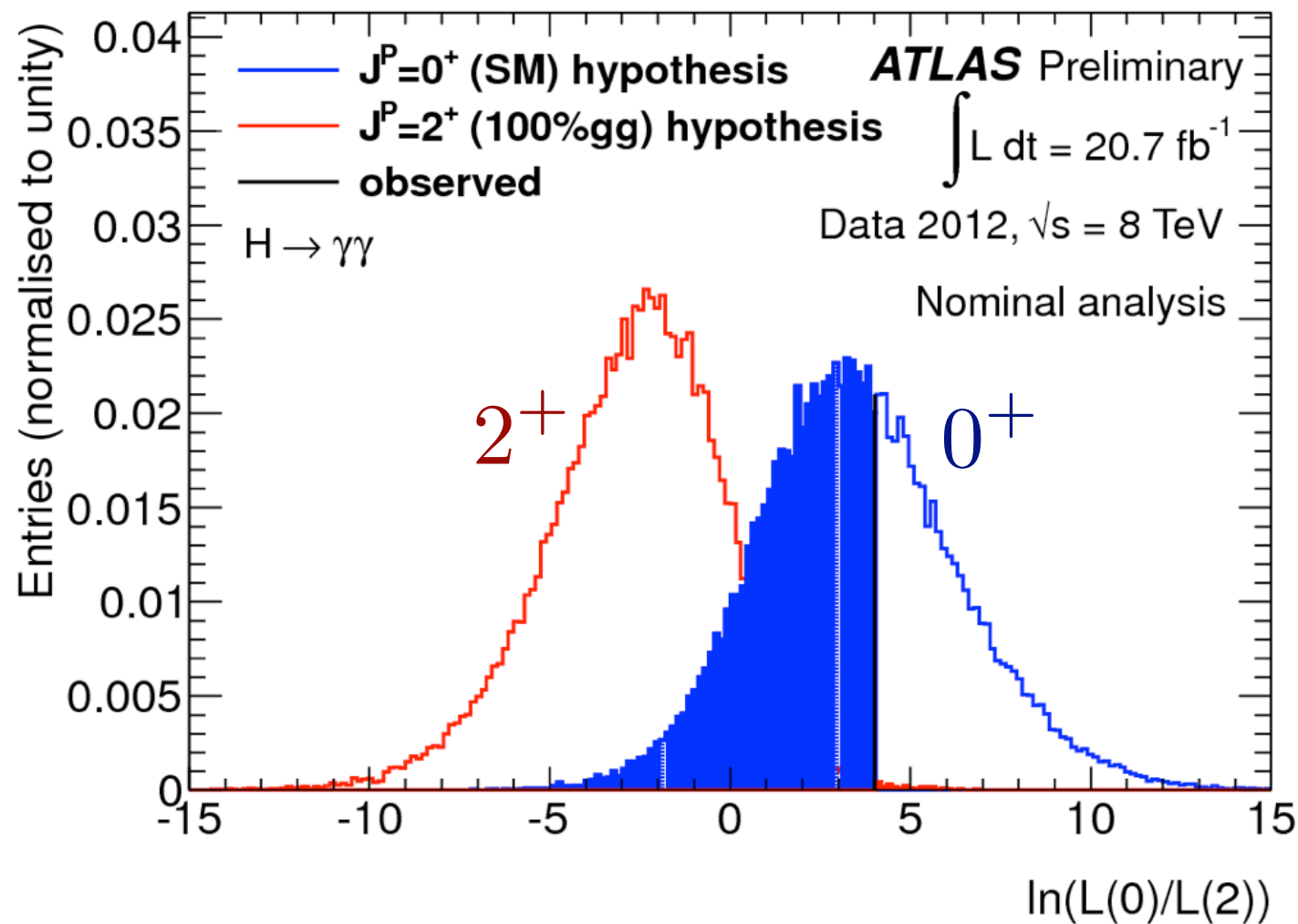
What one needs to know:

- ⇒ Its quantum numbers: $J^{PC} = 0^{++}$, gauge q.n.s
- ⇒ The strength of its interactions with all other particles and with itself
- ⇒ Is it “natural”?
- ⇒ Is it “elementary” or “composite”?
- ⇒ Is it alone or accompanied?

$J^P = ?$ (0^+ expected)

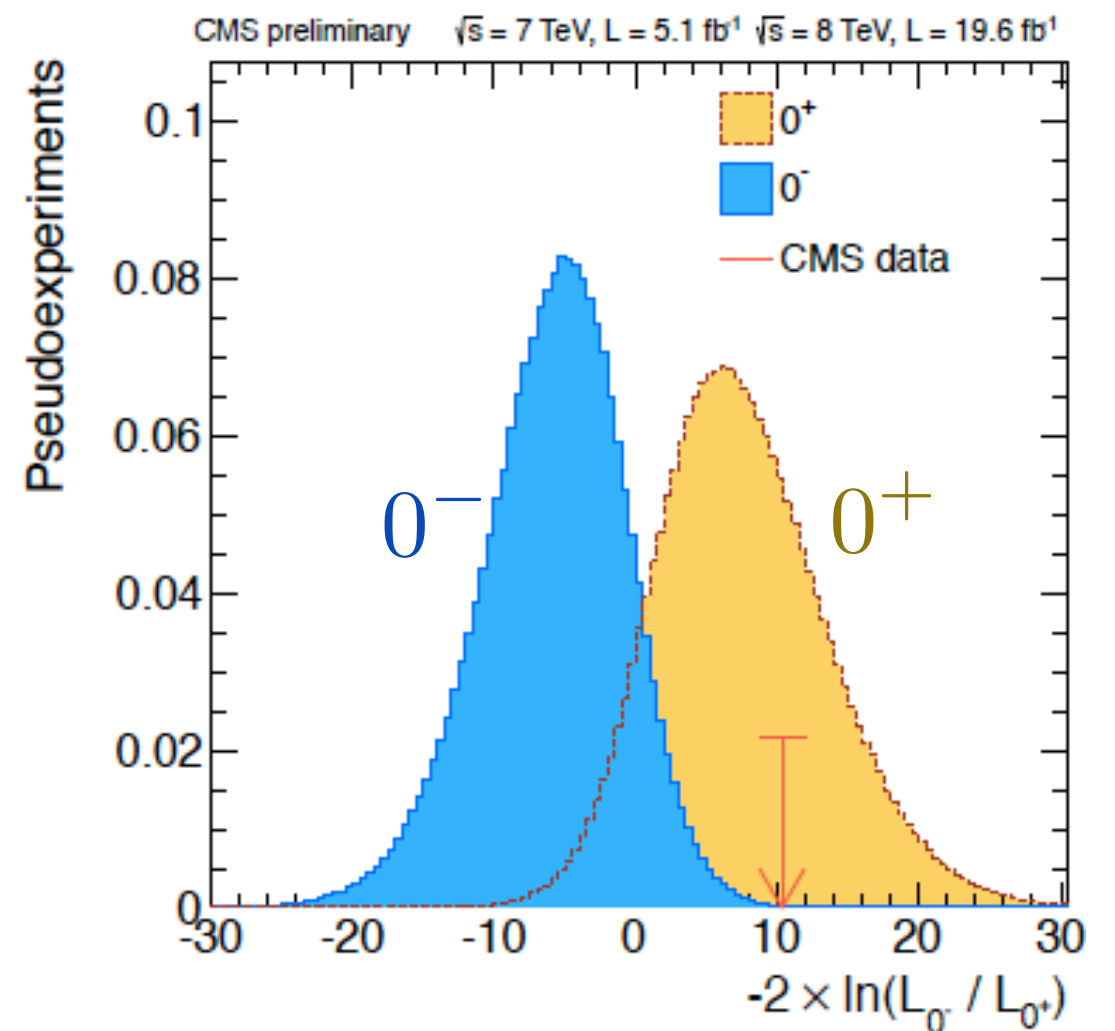
Parity and angular momentum discrimination by angular distribution in decays (pairwise hypothesis tests)

$$h \rightarrow \gamma\gamma$$



the angular momentum looks right

$$h \rightarrow ZZ^* \rightarrow 4l$$



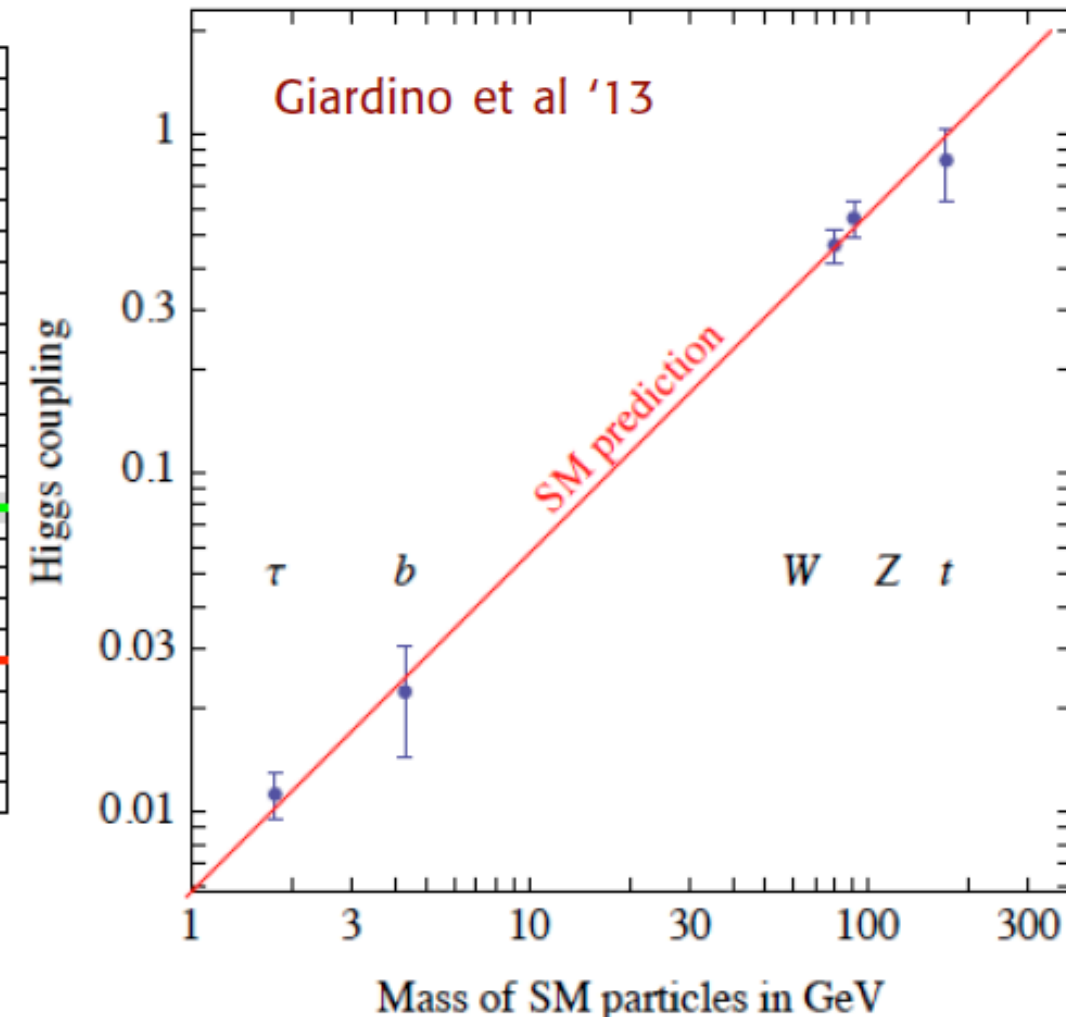
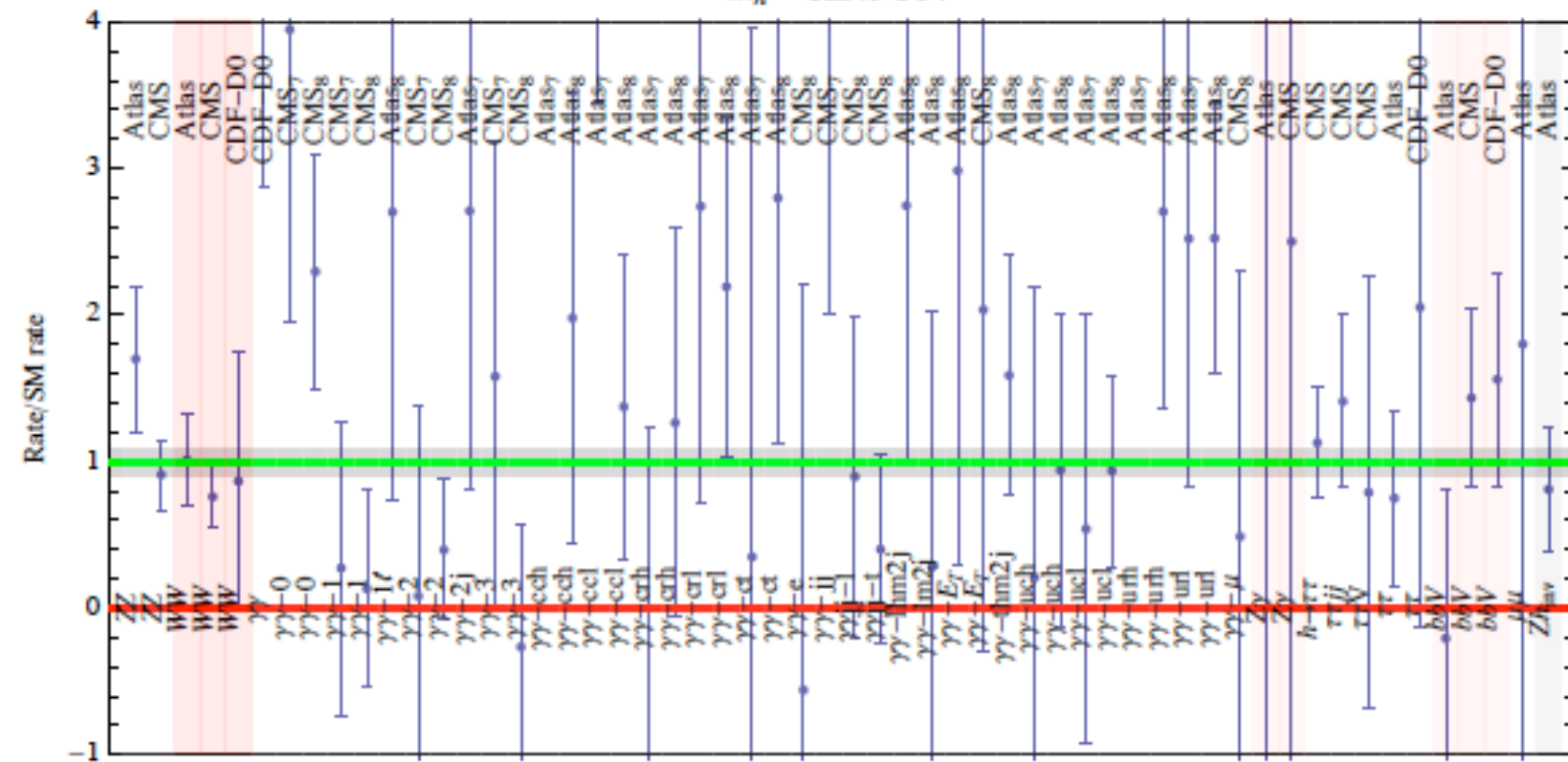
the parity looks right

The couplings to other particles

From a theorist's informal combination of ATLAS&CMS data

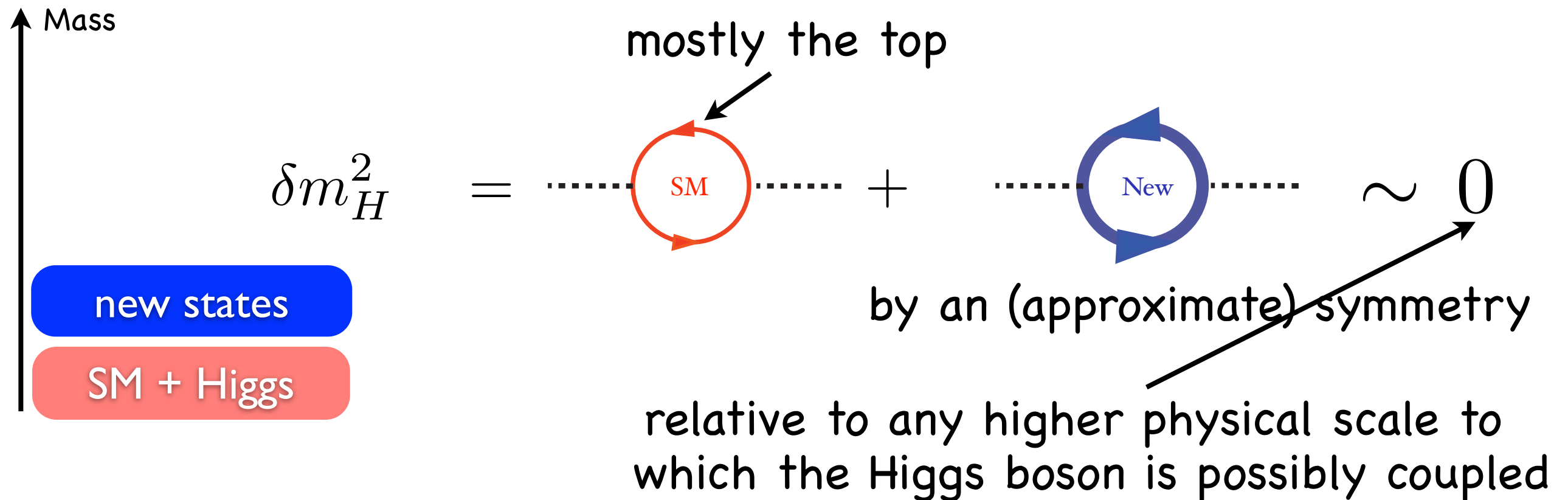
Giardino, Kannike, Masina, Raidal, Strumia
(as many others)

$m_h = 125.6 \text{ GeV}$



The coupling-versus-mass linear relation is
 an absolute prediction of the ST (not exhaustive: $gg, \gamma\gamma$)
 No Clebsch distortion: the Higgs boson is (close to) a doublet

A "natural", not Fine Tuned Higgs boson



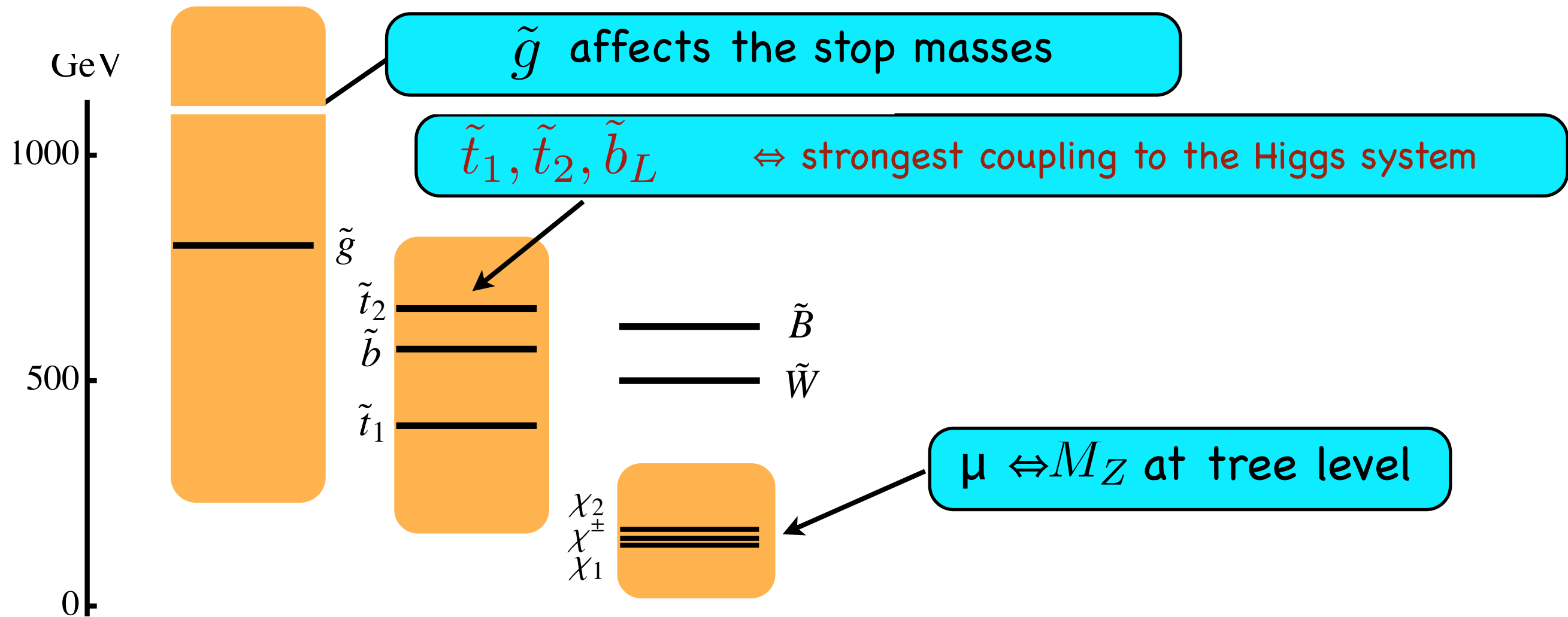
If so, explain why the great empirical success of the SM does not depend on unknown short distance physics

The "crucial" configuration of supersymmetry

(introduced well before the LHC)

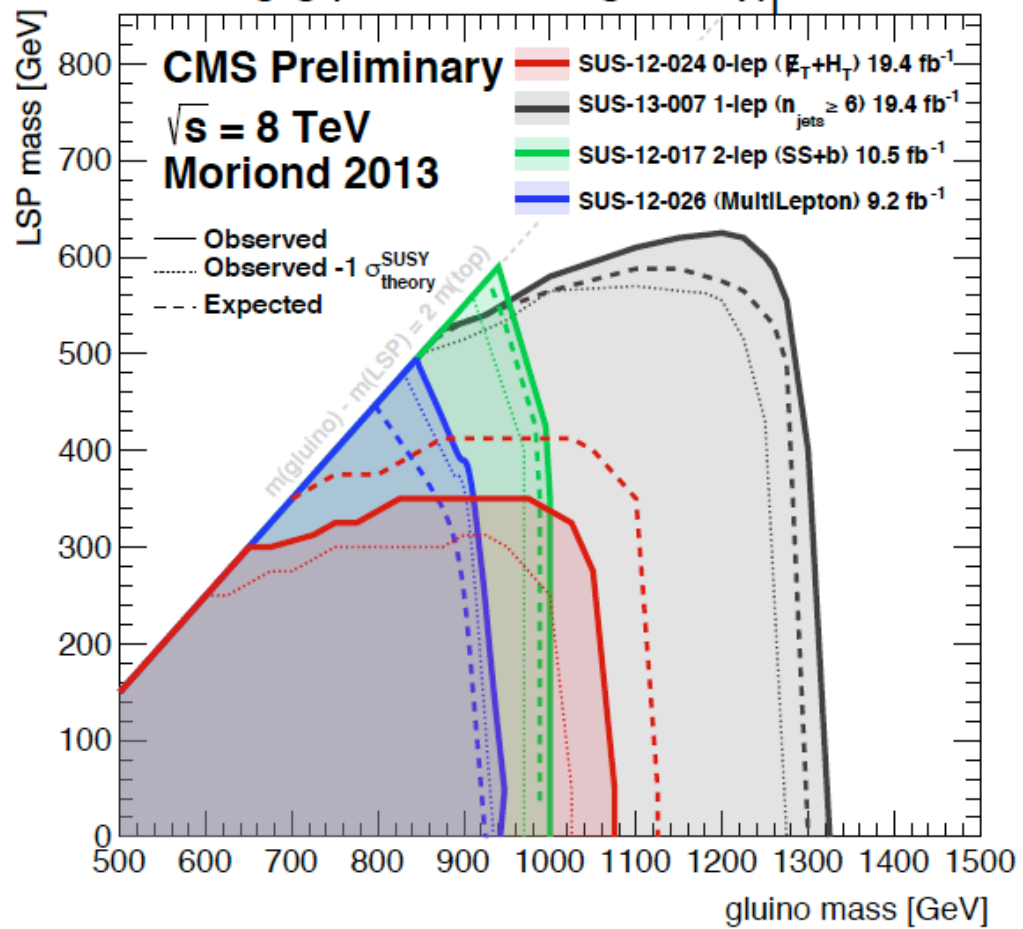
$$\delta m_H^2 = \text{SM} + \text{New} \sim 0$$

s-particles

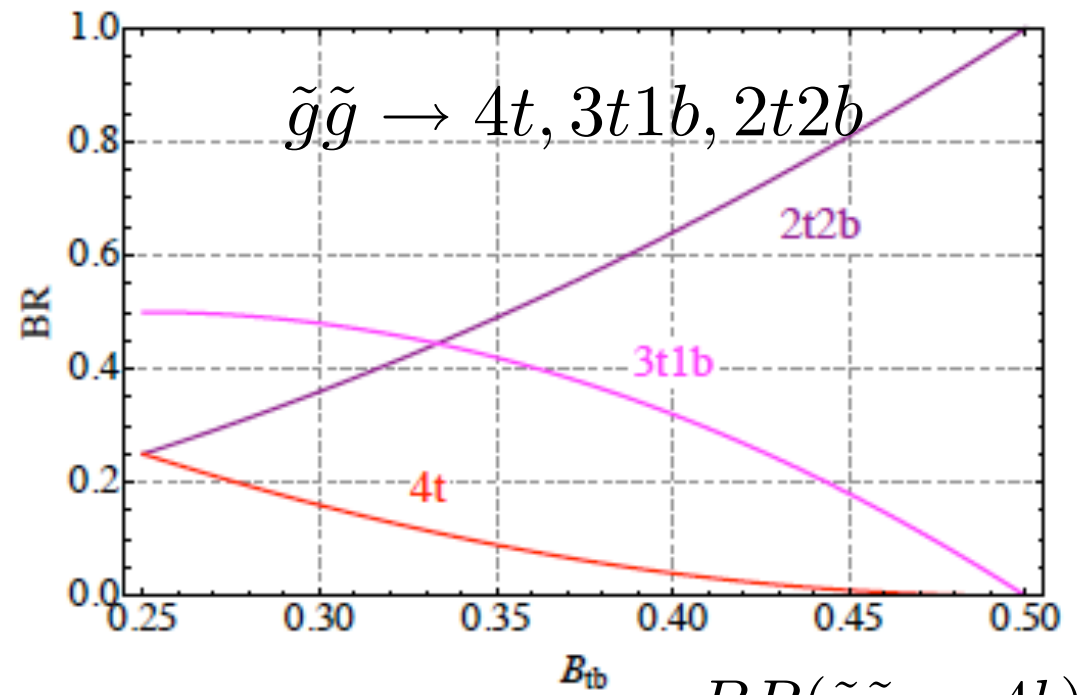
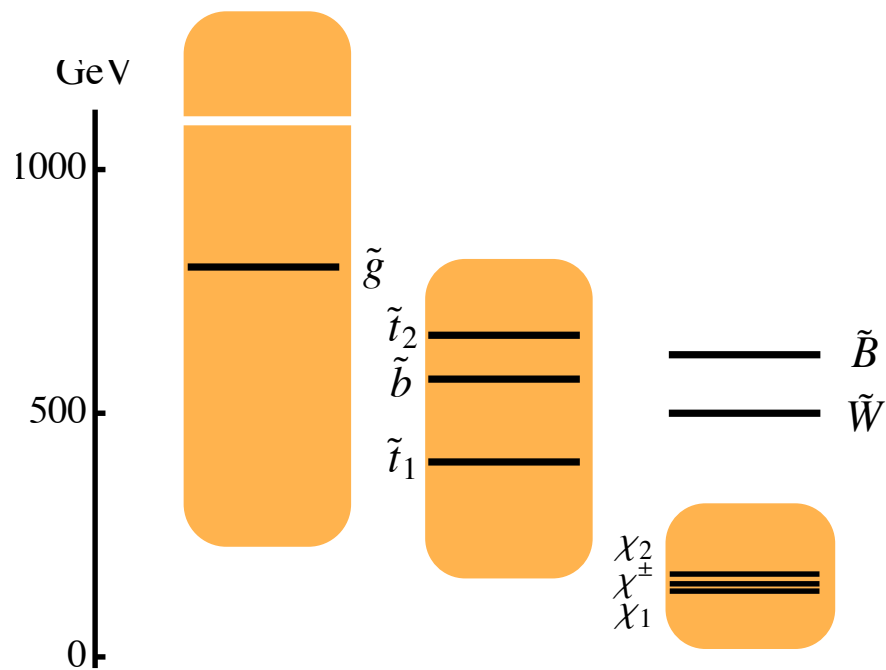
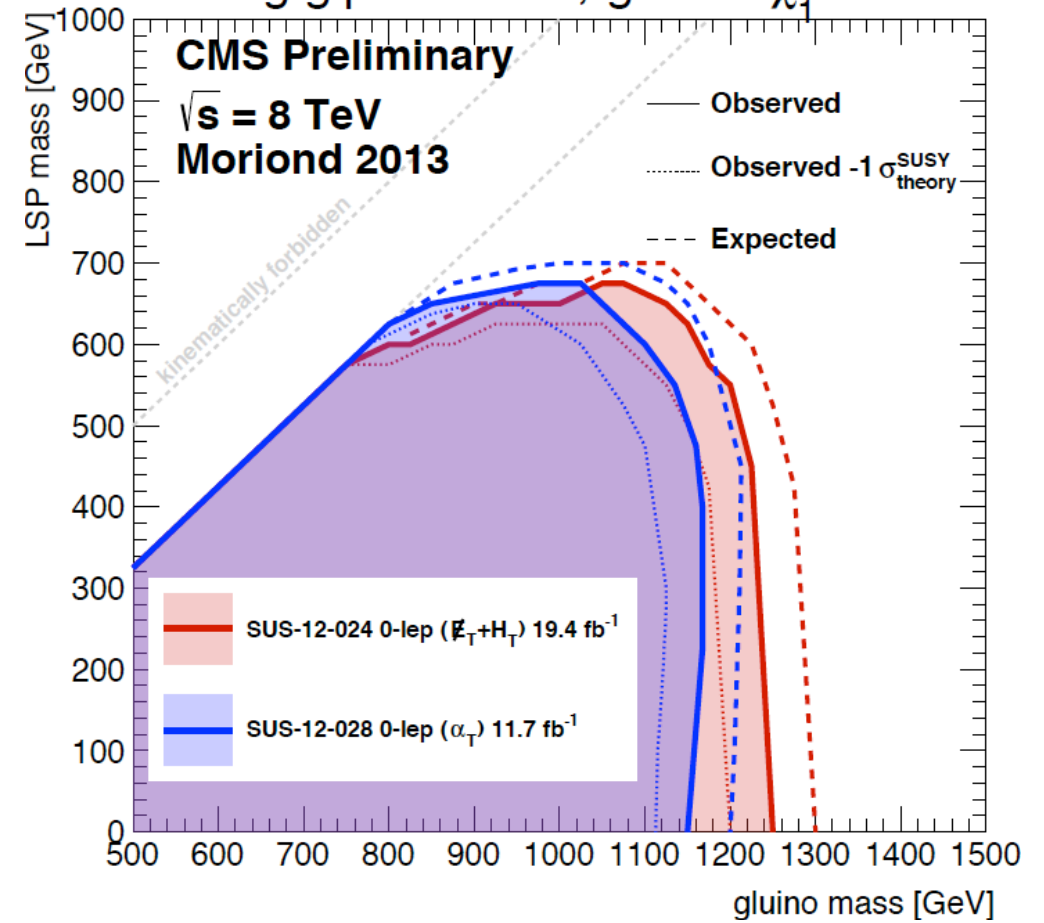


orange areas indicative and dependent
on how the Higgs boson gets its mass

$\tilde{g}\text{-}\tilde{g}$ production, $\tilde{g} \rightarrow t \bar{t} \tilde{\chi}_1^0$

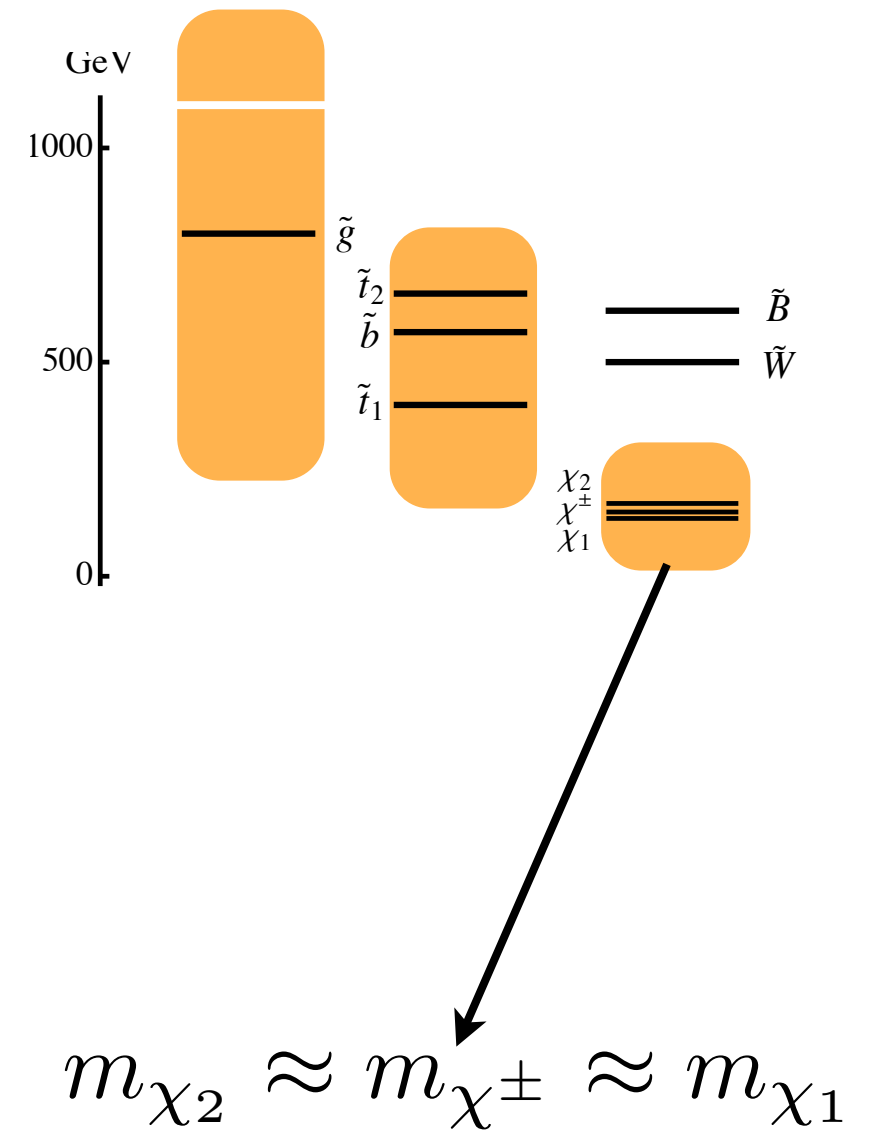
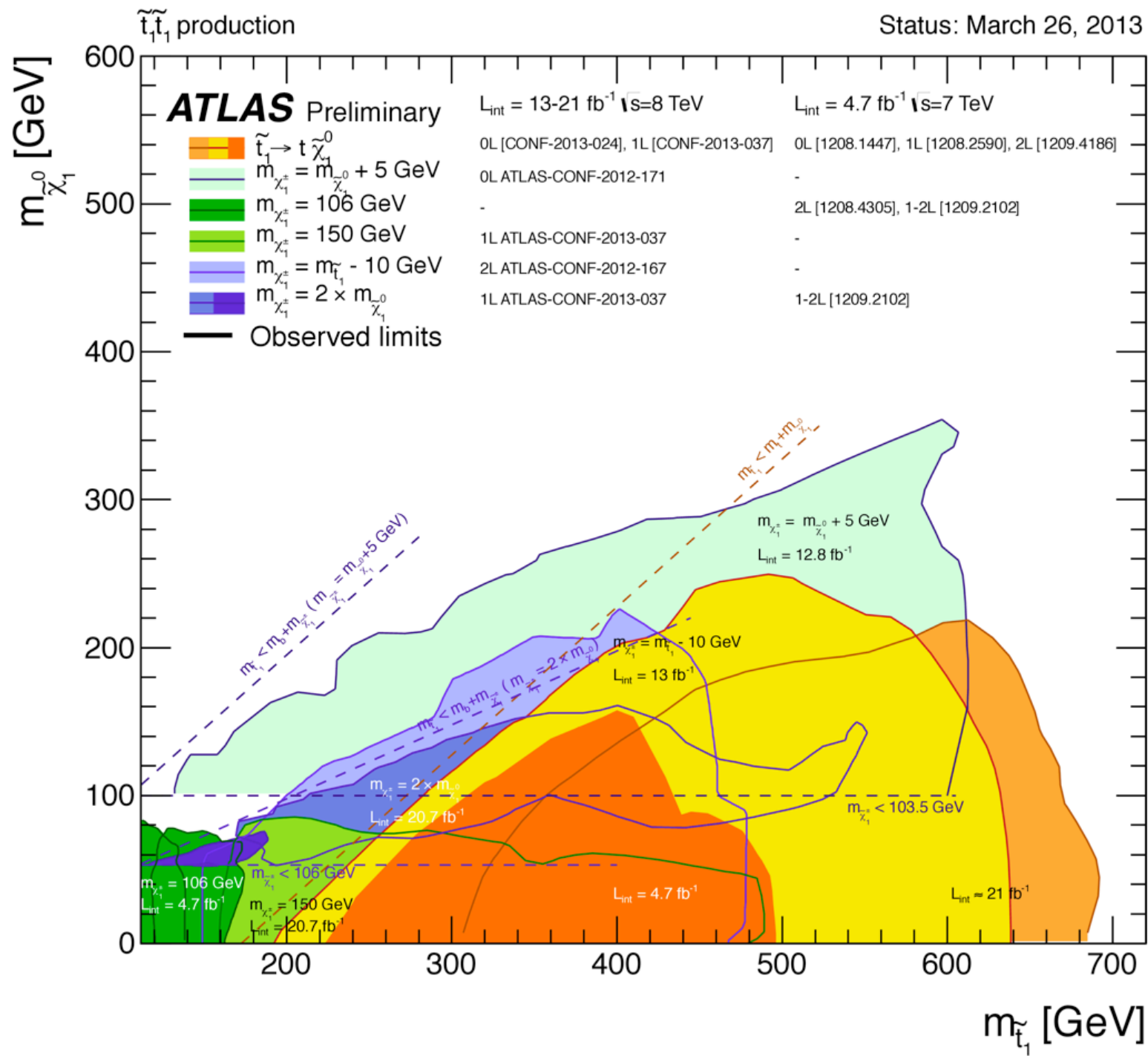


$\tilde{g}\text{-}\tilde{g}$ production, $\tilde{g} \rightarrow b \bar{b} \tilde{\chi}_1^0$



“naively” $m_{\tilde{g}} \gtrsim 1300 \text{ GeV}$
 “conservatively” $m_{\tilde{g}} \gtrsim 1000 \text{ GeV}$

$BR(\tilde{g}\tilde{g} \rightarrow 4b) \lesssim 4\%$
 ($\tan \beta \lesssim 10$)



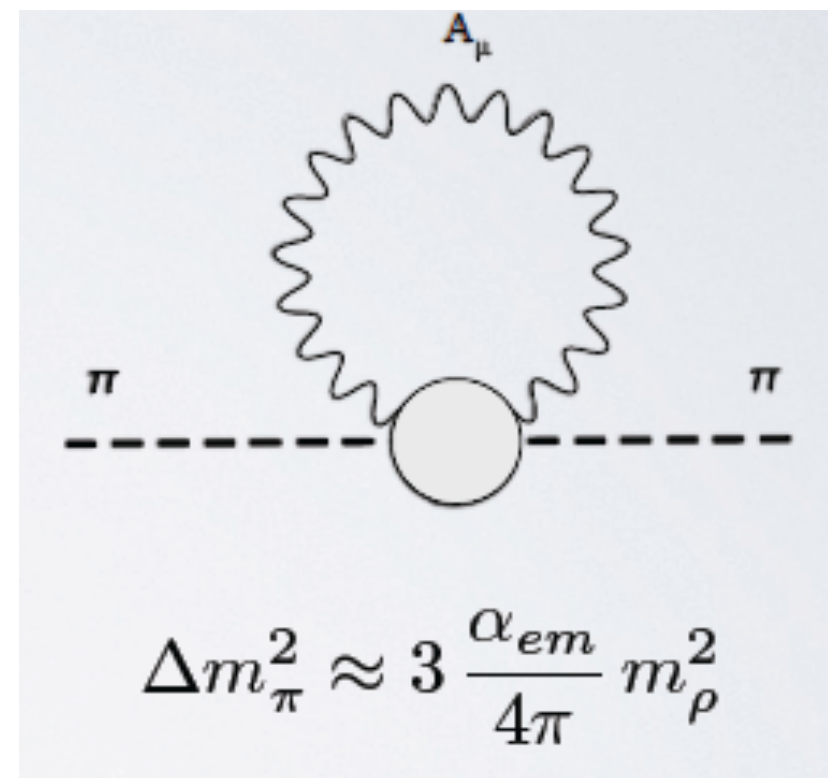
“naively” $m_{\tilde{t}_1} \gtrsim 700 \text{ GeV}$

“conservatively” $m_{\tilde{t}_1} \gtrsim 200 \div 300 \text{ GeV}$ (with $m_{\tilde{\chi}} = 150 \div 250 \text{ GeV}$)

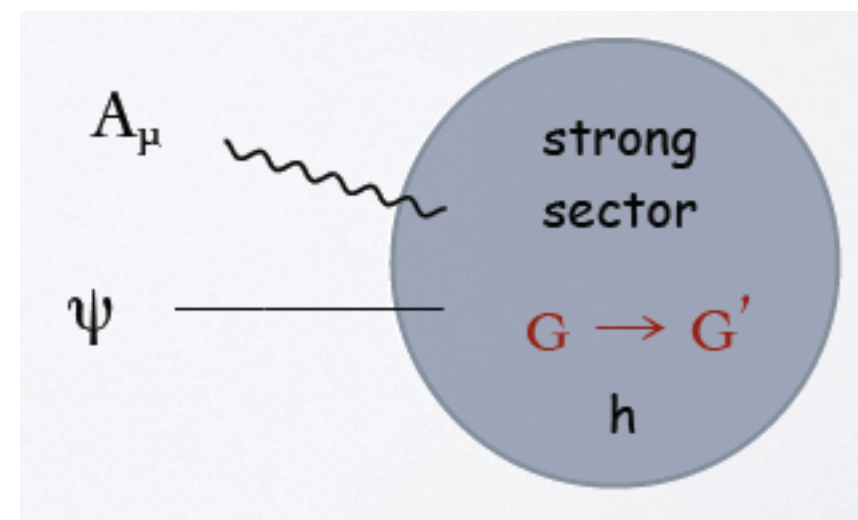
The Higgs boson as a pseudo-Goldstone boson (hence a composite rather than a "fundamental" object)

The pion as an analogy:

$$SU(2)_L \times SU(2)_R \Rightarrow SU(2)_I$$
$$\Delta m_\pi^2 = m_{\pi^+}^2 - m_{\pi^0}^2$$



A new strong sector
at the TeV scale

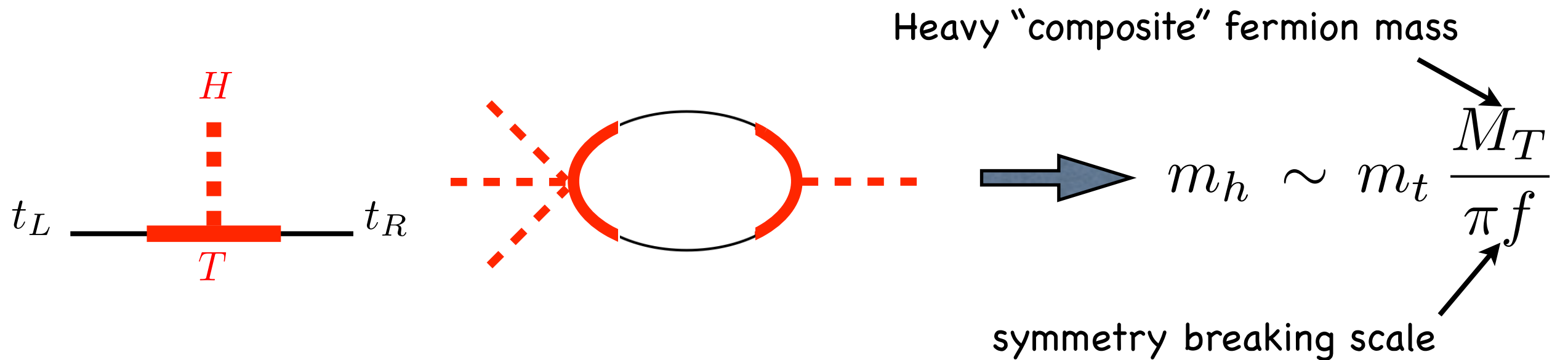


Like the pion in QCD, the Higgs boson as a quasi GB
of a spontaneously broken global symmetry at the TeV

More in detail

$$\delta m_H^2 = \text{---} \circlearrowleft \text{SM} \text{---} + \text{---} \circlearrowleft \text{New} \text{---} \sim 0$$

Heavy "composite" fermions



Most common:

$$Q = \begin{pmatrix} T \\ B \end{pmatrix}, \quad X = \begin{pmatrix} X_{5/3} \\ T_{2/3} \end{pmatrix}, \quad \tilde{T}$$

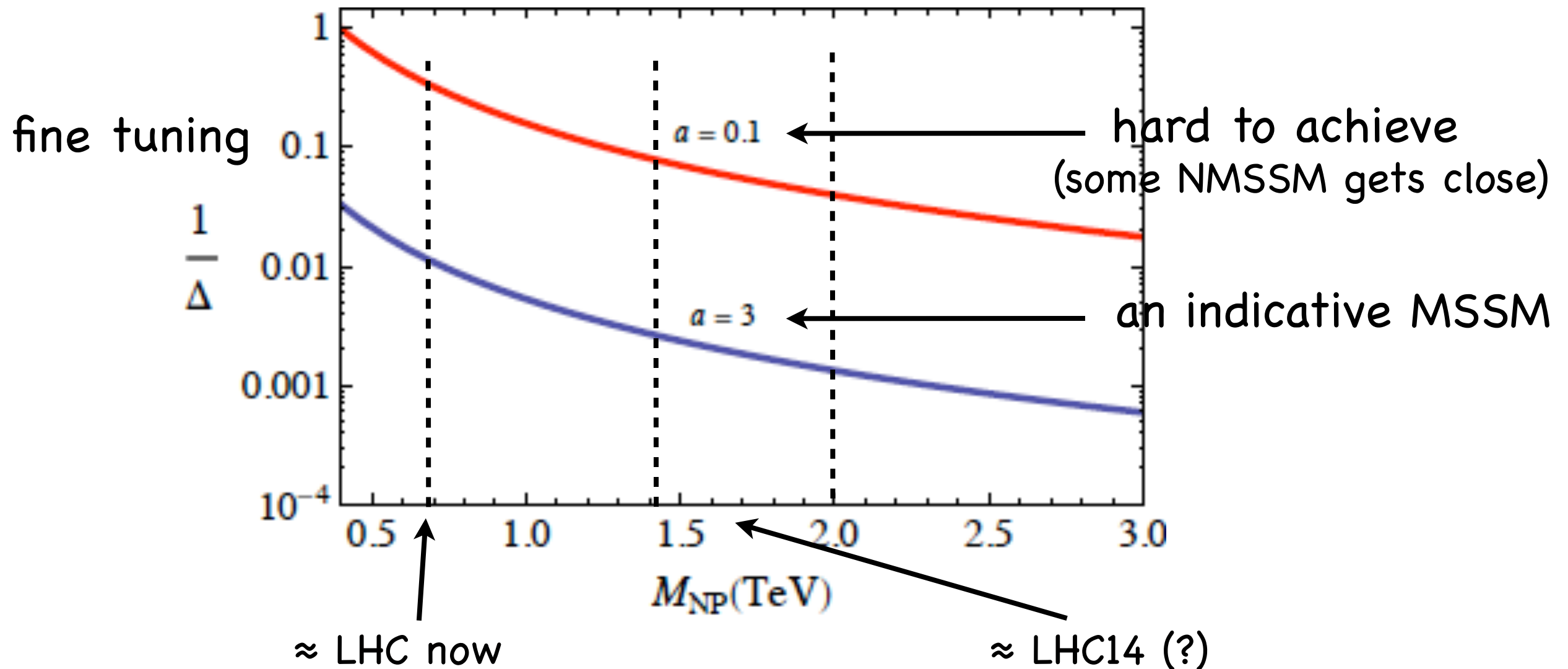
Current searches exclude masses below 500-800 GeV, depending on the charge

A quantitative measure (!?) of naturalness

$$\delta m_h^2 \approx a M_{NP}^2 < \Delta m_h^2$$

model dependent

a measure of fine tuning
(which exist in nature)



Last but not least: one or more Higgs bosons?

The pro's for one:

1. simplicity

How about the 12 (18) matter and the 12 (3) vector states?

2. electromagnetism always preserved

From 2 to 3 phases only

3. flavour

No big reason to be proud of the λ_{ij}

4. a single tuning, in case

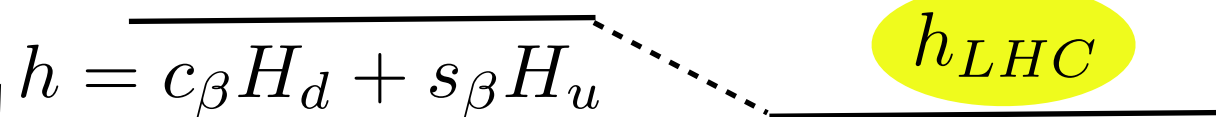
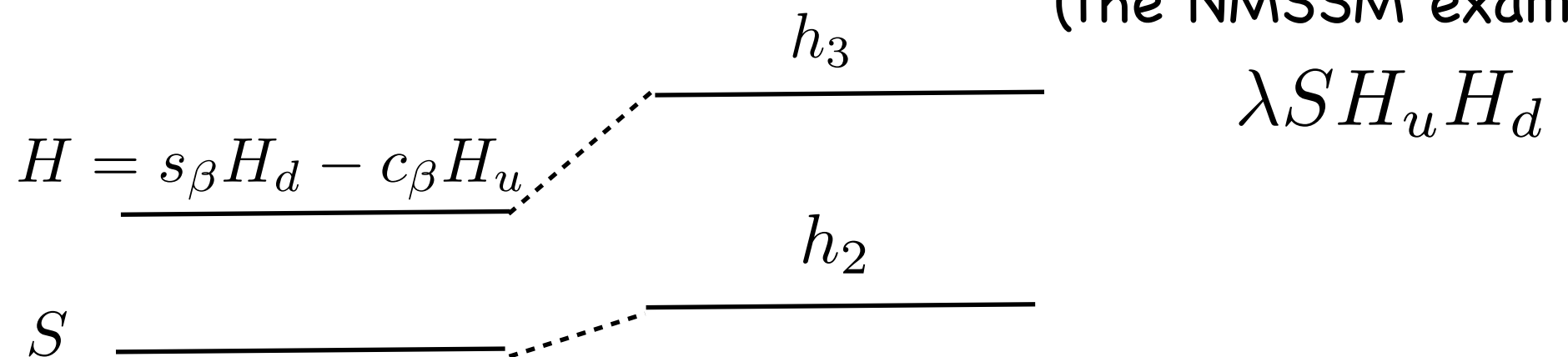
None is better, which often demands more Higgs boson

Two ways to attack the problem

⇒ By direct search $pp \rightarrow h_{\neq LHC} + X$
 ↓
 → decay products

⇒ By precision measurements of the couplings of
 the 125 GeV (quasi-standard) Higgs boson

(the NMSSM example)



has SM properties

Current status of the MSSM

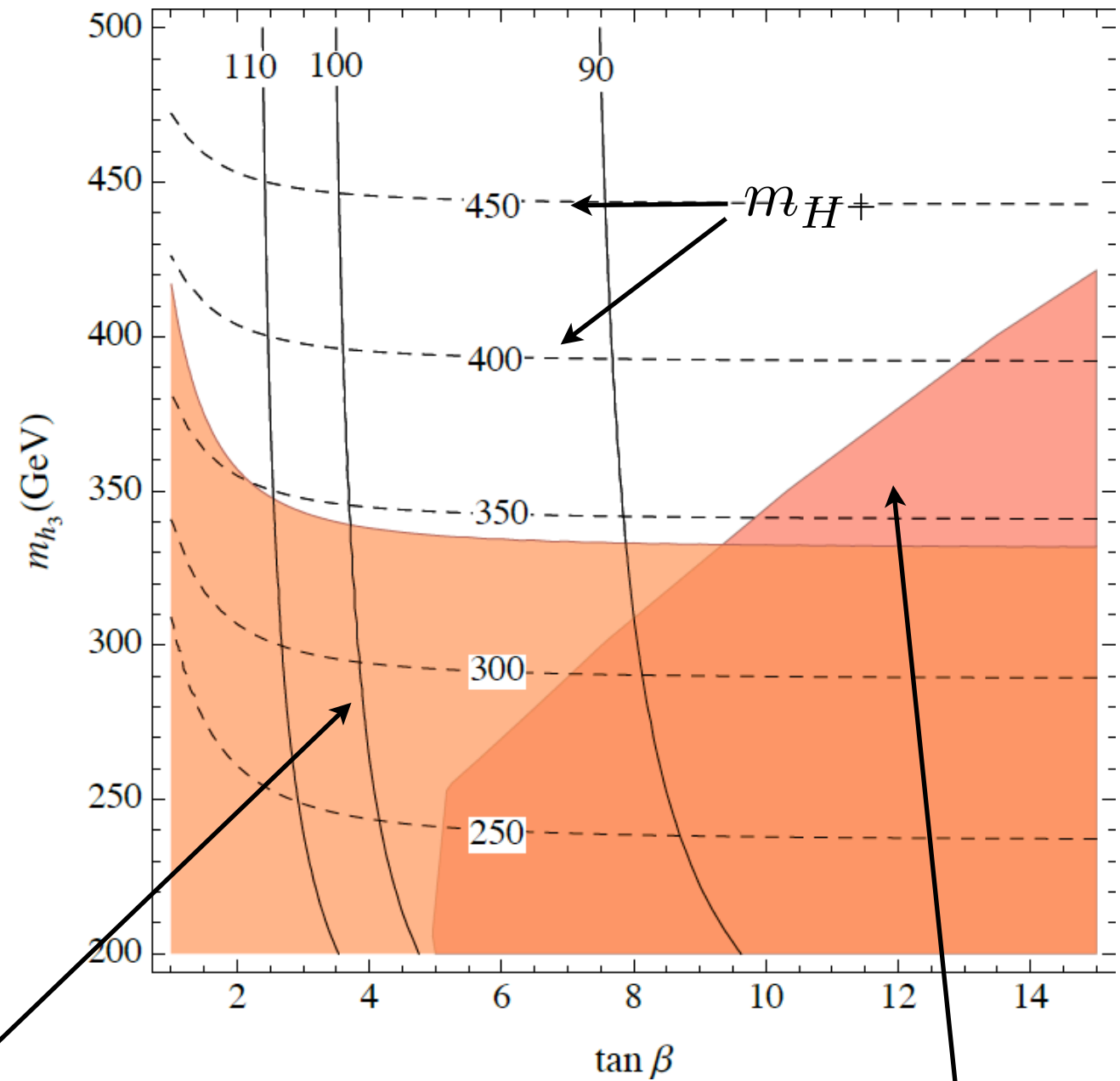
$$H = s_\beta H_d - c_\beta H_u$$

$$h = c_\beta H_d + s_\beta H_u$$

h_3

h_{LHC}

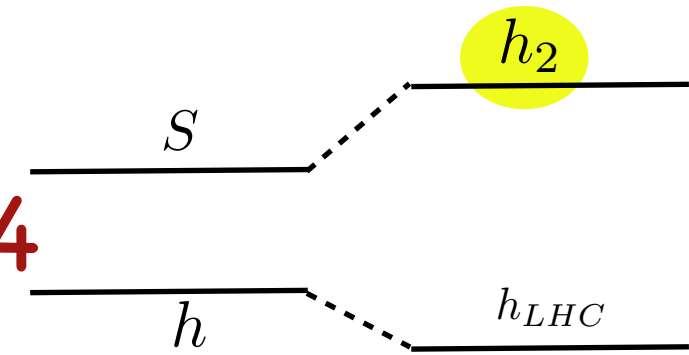
MSSM at variable Δ_t



"excluded" by h_{LHC} -signal strenghts

"excluded" by $h_3, A \rightarrow \tau\tau$

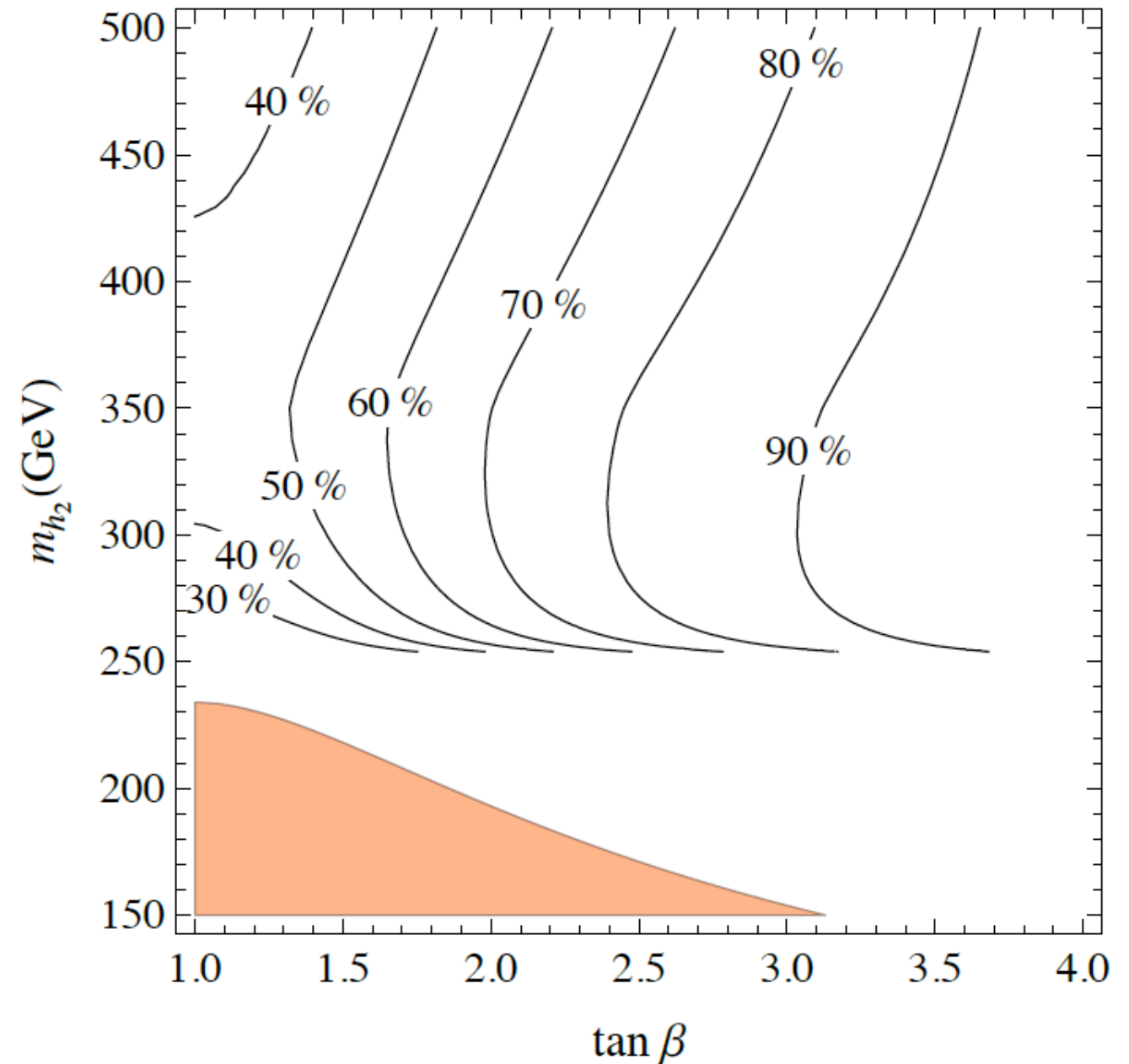
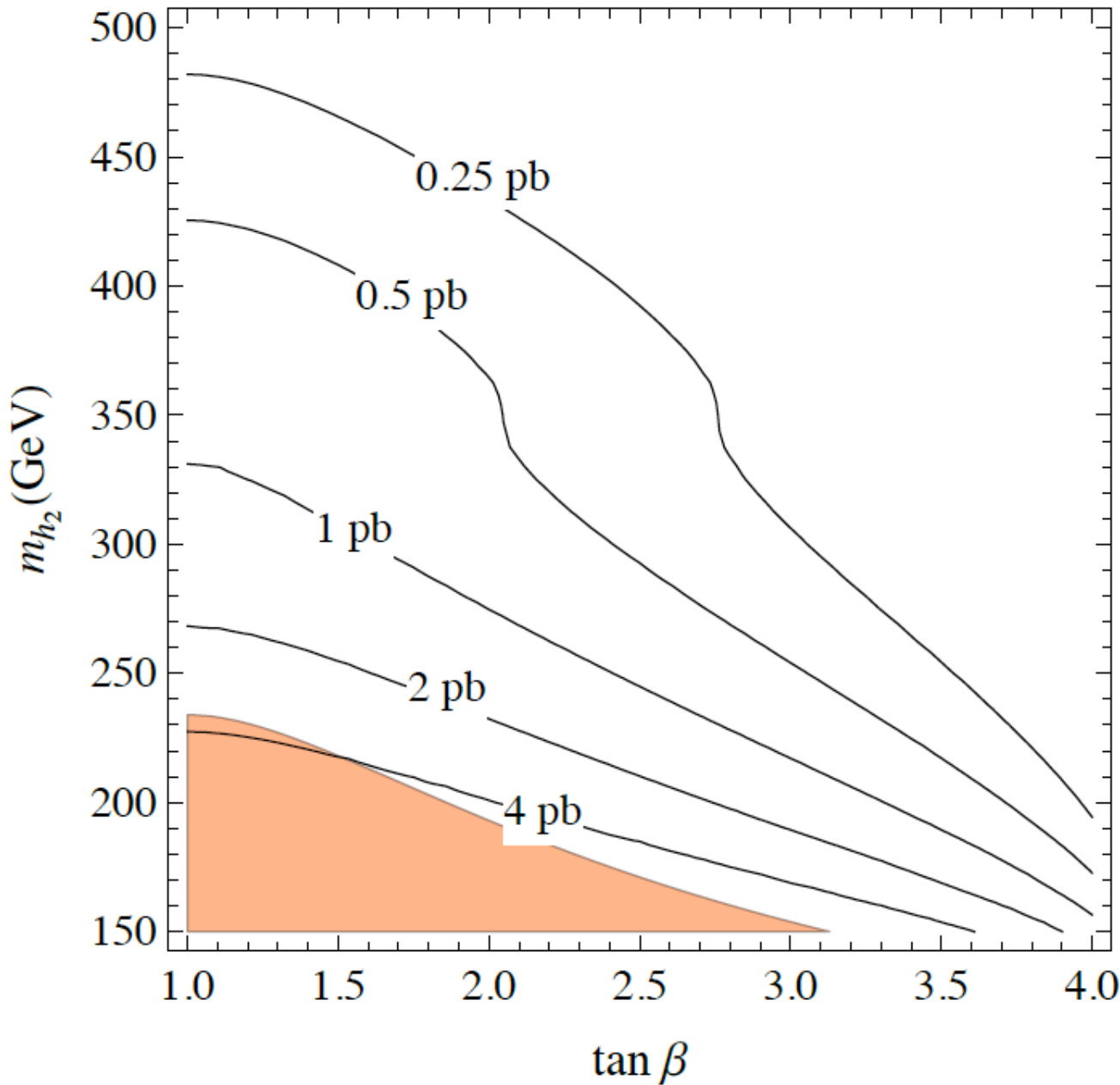
NMSSM: Direct search at LHC14



$\sigma(gg \rightarrow h_2)$

$\lambda = 0.8$

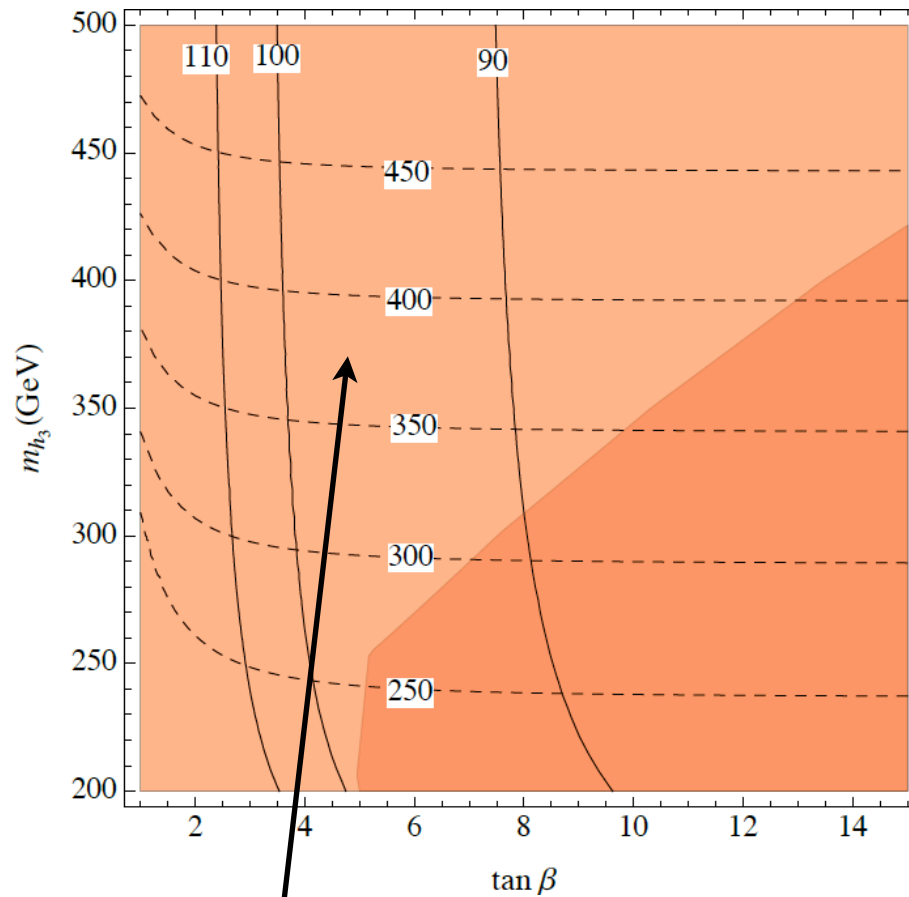
$BR(h_2 \rightarrow h_1 h_1)$



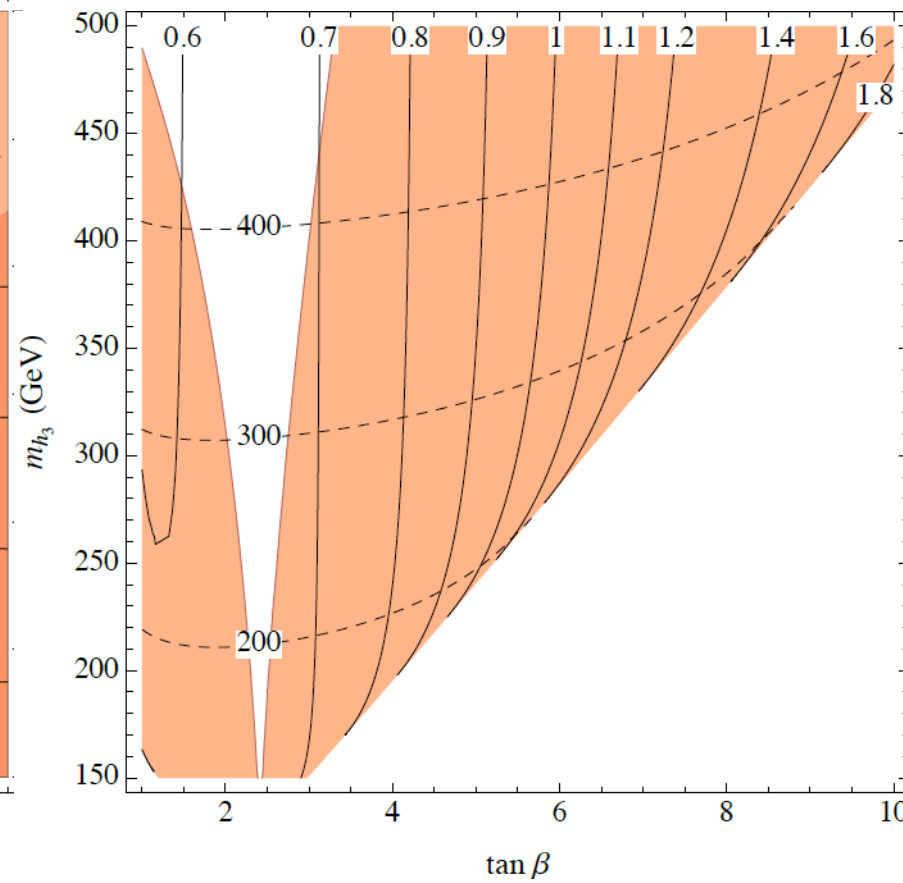
(changes in the h_1 self coupling by factors 3-4 possible)

A projection from the measurements of the signal strengths of h_{LHC}

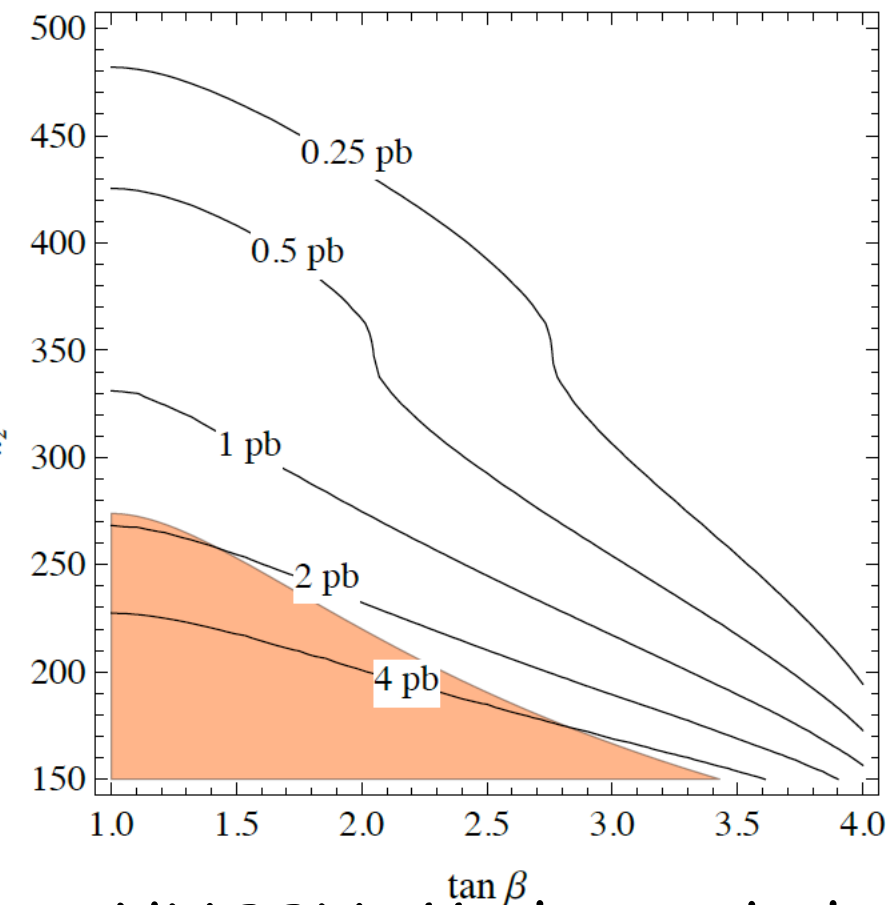
LHC14 at $300 fb^{-1}$ with ATLAS/CMS projected errors



MSSM



NMSSM, S-decoupled

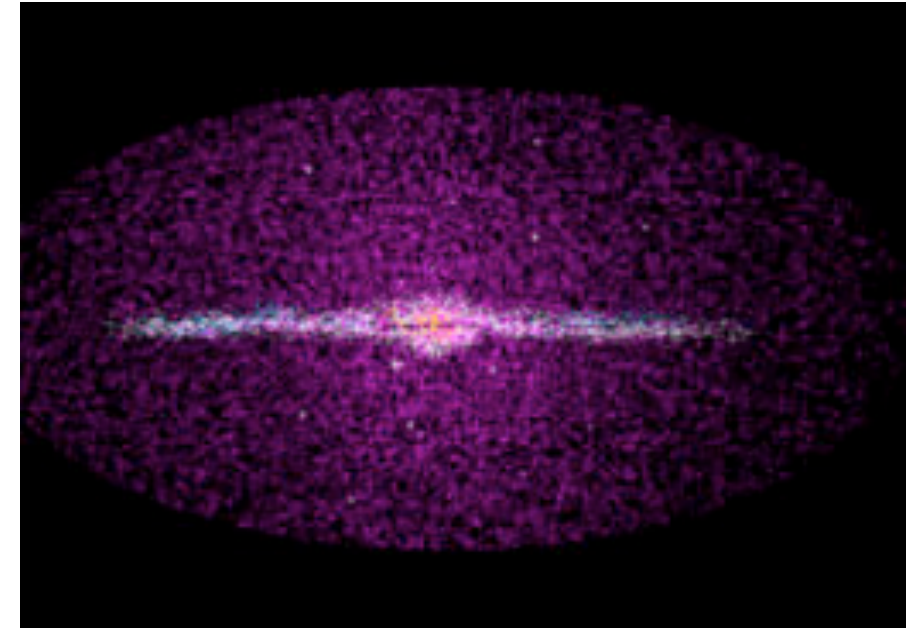
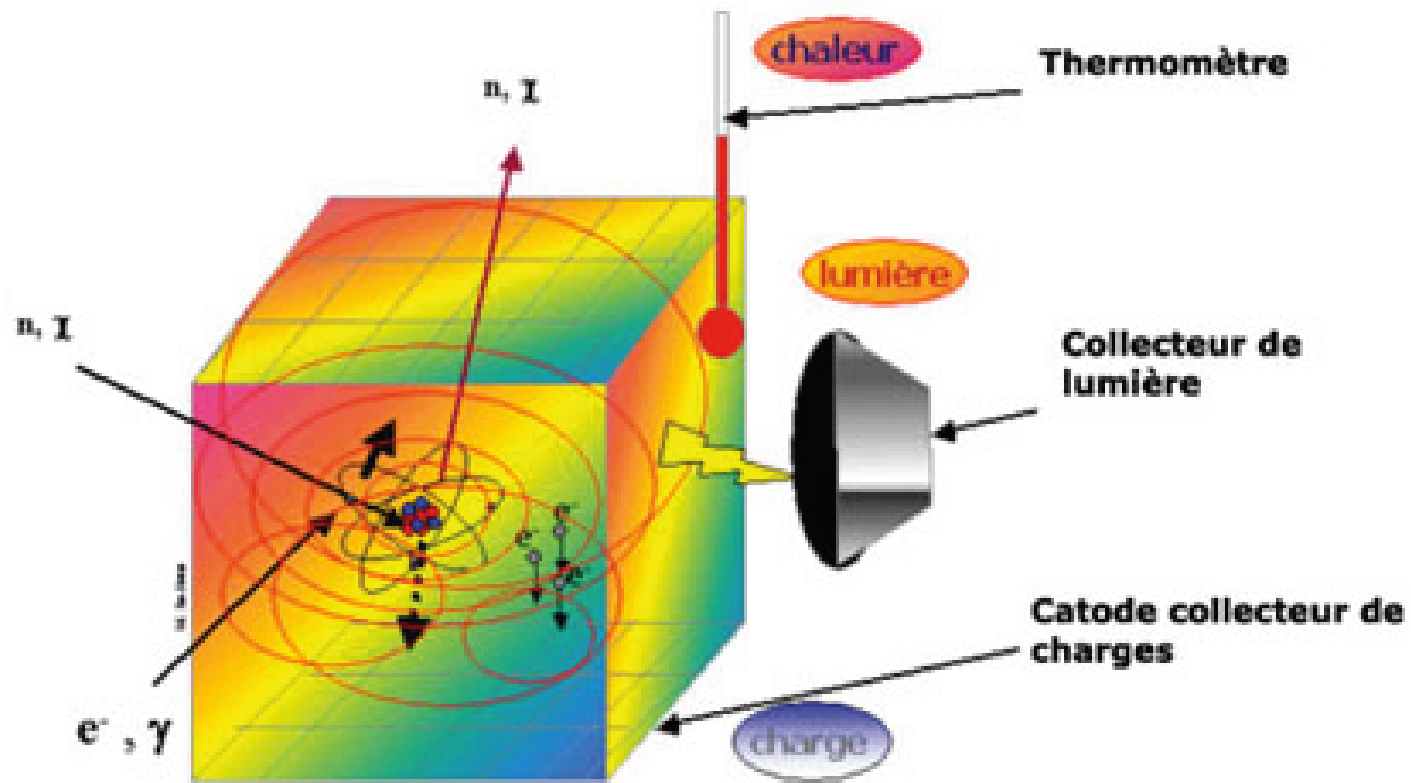


NMSSM, H-decoupled

The sensitivity region extends up to about 1 TeV for m_{h_2}

The direct search for Dark Matter

$$\chi N \rightarrow \chi N$$



$$Z_\mu \bar{\Psi}_M \gamma_\mu \Psi_M = 0$$

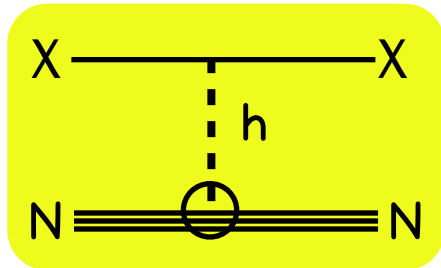
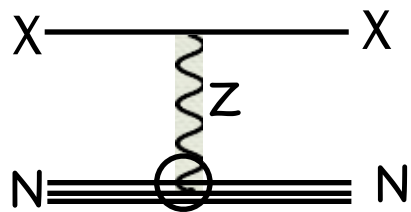
$$Z_\mu \bar{\Psi}_M \gamma_\mu \gamma_5 \Psi_M \neq 0$$

DM searches and the Higgs boson

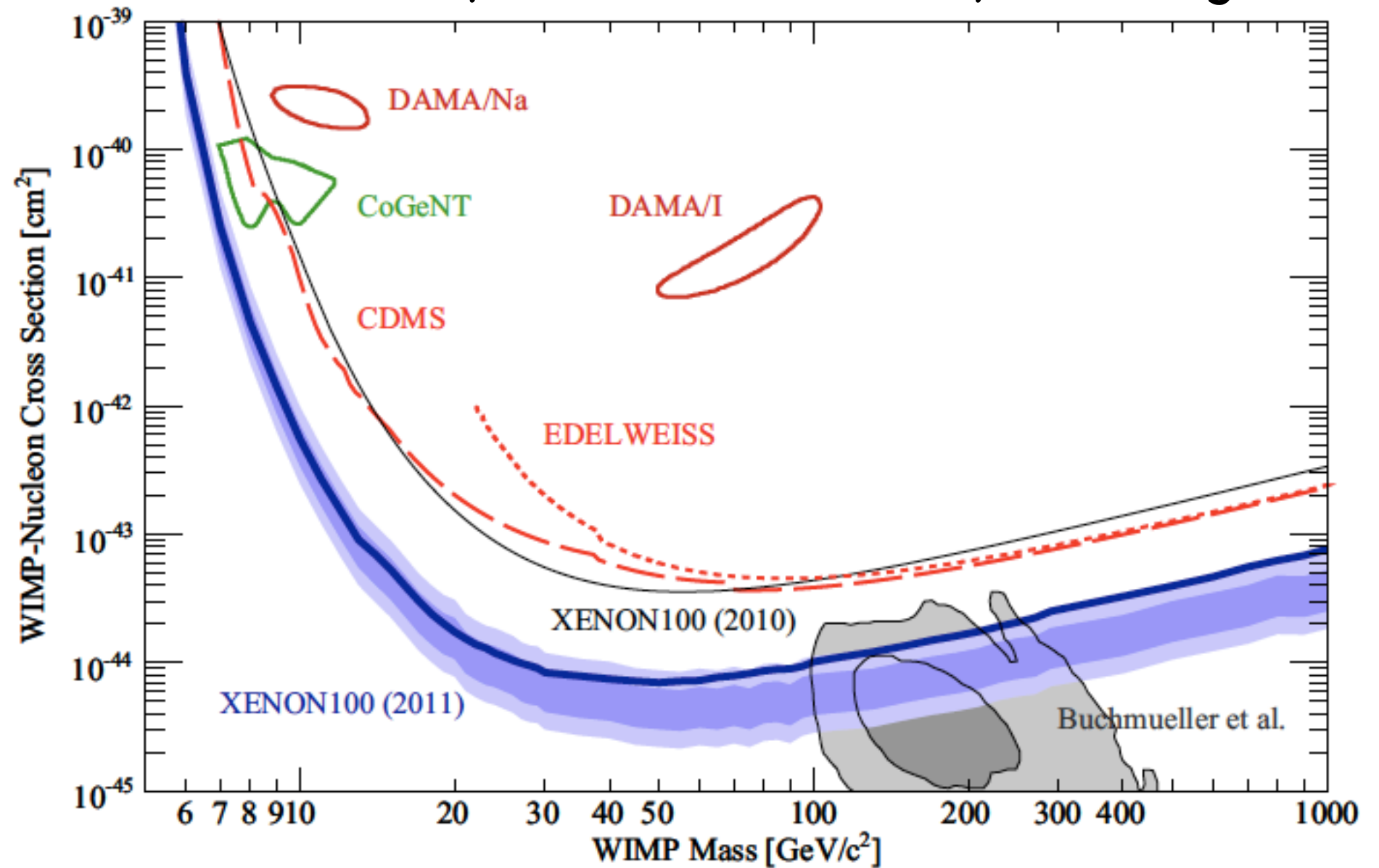
$$\chi N \rightarrow \chi N$$

3 events/1.8 backgd

$\sigma_Z(\chi N)$ spin indep.
excluded since
long time



exclusion by XENON100 (100 days x 48 kgs)



Higgs boson exchange being probed now for $m_h = 125 \text{ GeV}$

$$\sigma_h(\chi N) \approx 10^{-43} \text{ cm}^2 \left(\frac{\lambda}{0.1}\right)^2 \left(\frac{100 \text{ GeV}}{m_\chi}\right)^2 \left(\frac{100 \text{ GeV}}{m_h}\right)^4$$

Conclusion

1. The discovery of the Higgs boson:

might be BOTH the coronation of the Standard Theory
AND
a first step towards unexplored territory

2. Natural or unnatural theories?

before accepting a shift of paradigm,
useful to be patient and careful

3. One or more Higgs bosons?

could be the lightest new particle(s) around

4. What about the flavour puzzle?

$m's, V_{CKM} \Leftrightarrow \lambda_{ij}^{Yukawa}$: a great embarrassment,
unlikely to be solved without new key data

Conversation with Mrs Thatcher: 1982

What do you do?

Think of things for the experiments to look for, and hope they find something different

Wouldn't it be better if they found what you predicted?

Then we would not know how to proceed!

Guess who is Mrs. Thatcher and who is the other guy

NMSSM

$$\Delta f = \lambda H_u H_d$$

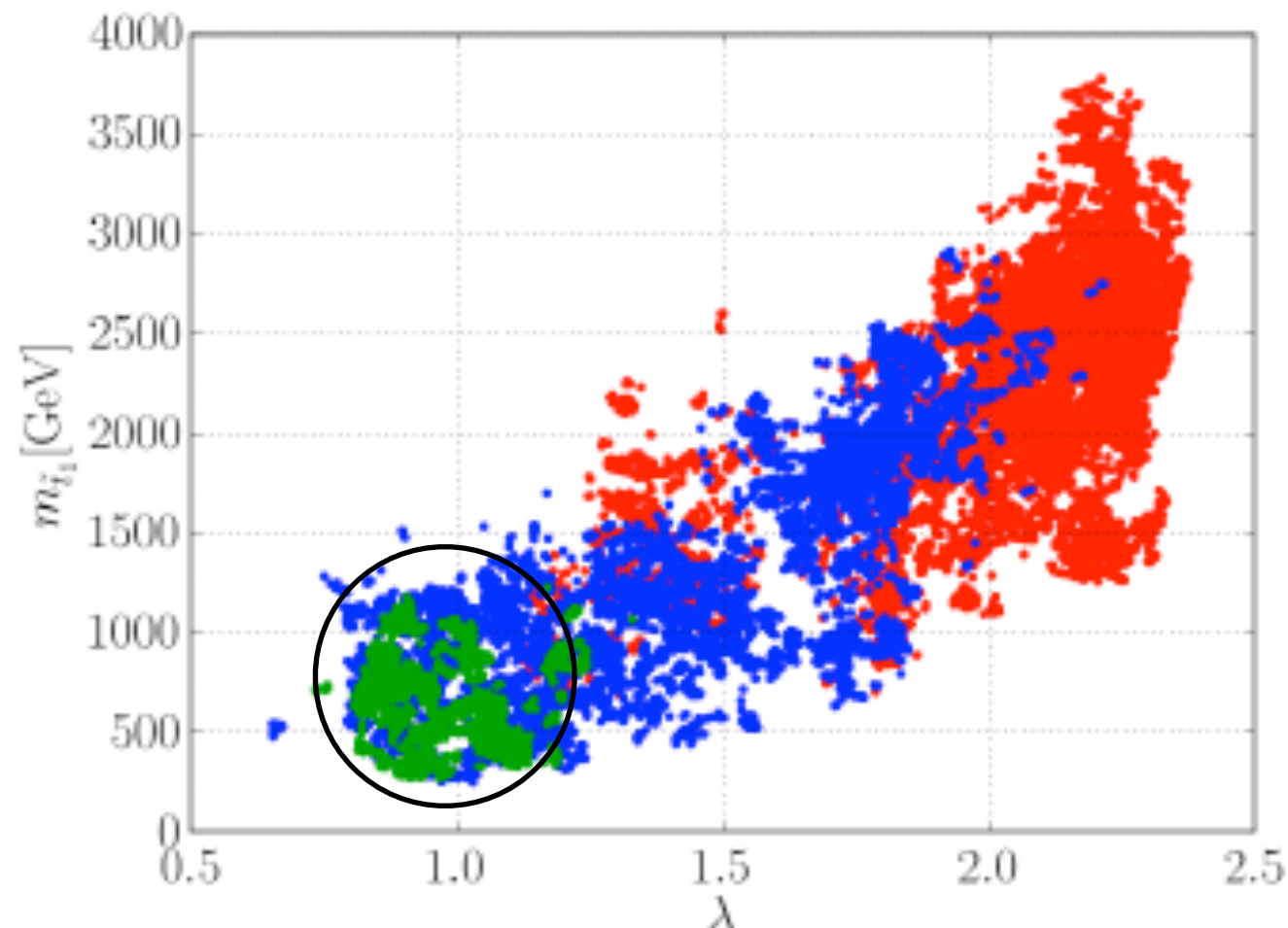
Fayet 1975

Two independent reasons to consider it:

1. Add an extra contribution to $m_{hh}^2 = m_Z^2 c_{2\beta}^2 + \Delta_t^2 + \lambda^2 v^2 s_{2\beta}^2$ thus allowing for lighter stops

2. Alleviates fine tuning in v for $\lambda \gtrsim 1$ and moderate $\tan \beta$

$$\left. \frac{dv^2}{dm_{H_u}^2} \right|_{NMSSM} \approx \frac{\kappa}{\lambda^3} \cot 2\beta \quad \text{versus} \quad \left. \frac{dv^2}{dm_{H_u}^2} \right|_{MSSM} \approx \frac{4}{g^2}$$



green points have better than 5% "combined" fine-tuning and $\Lambda_{mess} = 20 \text{ TeV}$ in the scale invariant NMSSM

$$m_{\tilde{t}_1} < 1.2 \text{ TeV}$$

$$m_{\tilde{g}} < 3 \text{ TeV}$$

Gherghetta et al 2012

Can the extra Higgs bosons of the NMSSM be the lightest new particles around?

⇒ Assume a negligibly small CPV in the Higgs sector

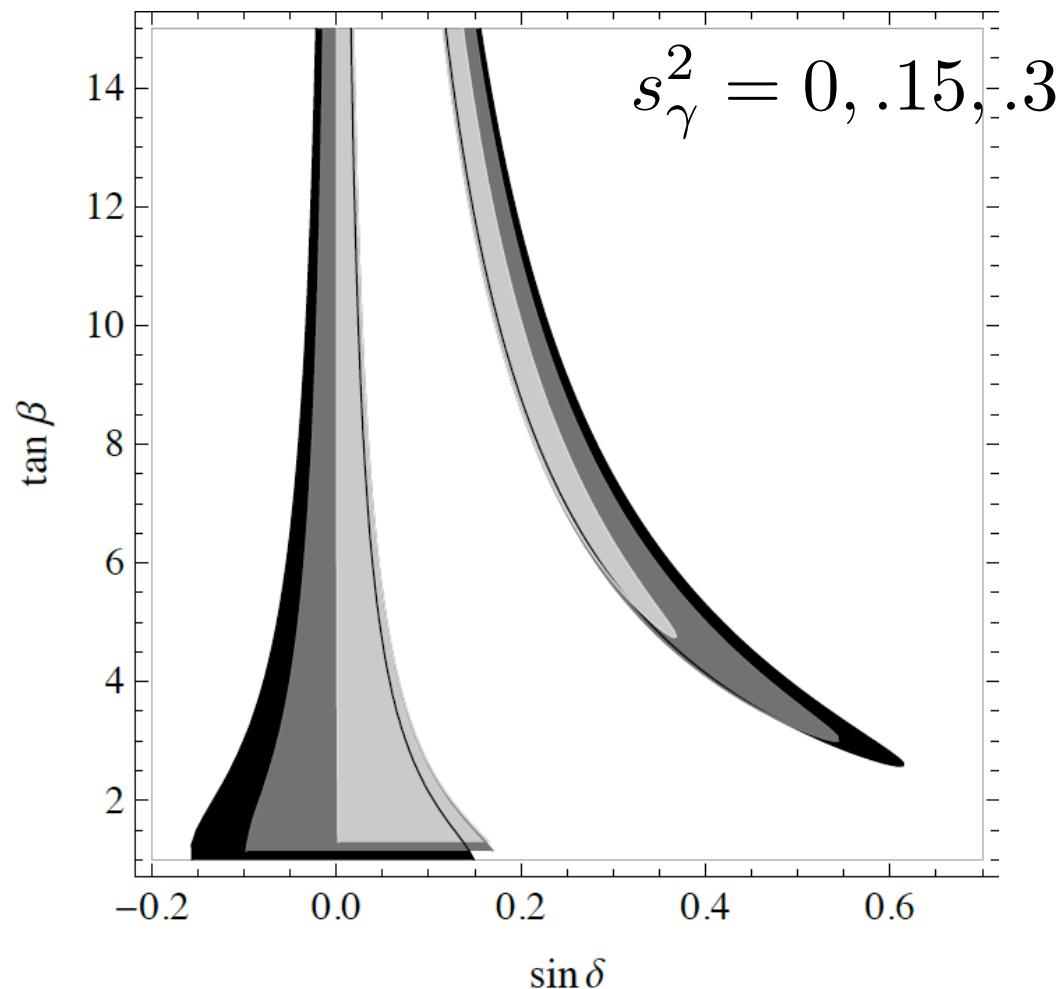
$$\mathcal{H} \equiv (H_d, H_u, S)^T = R_\alpha^{12} R_\gamma^{23} R_\sigma^{13} (h_3, h_1, h_2)^T \equiv R \mathcal{H}_{\text{ph}}$$

⇒ Take $h_1 = h_{LHC}$ with $m_{h_1} > m_{h_2}, m_{h_3}$

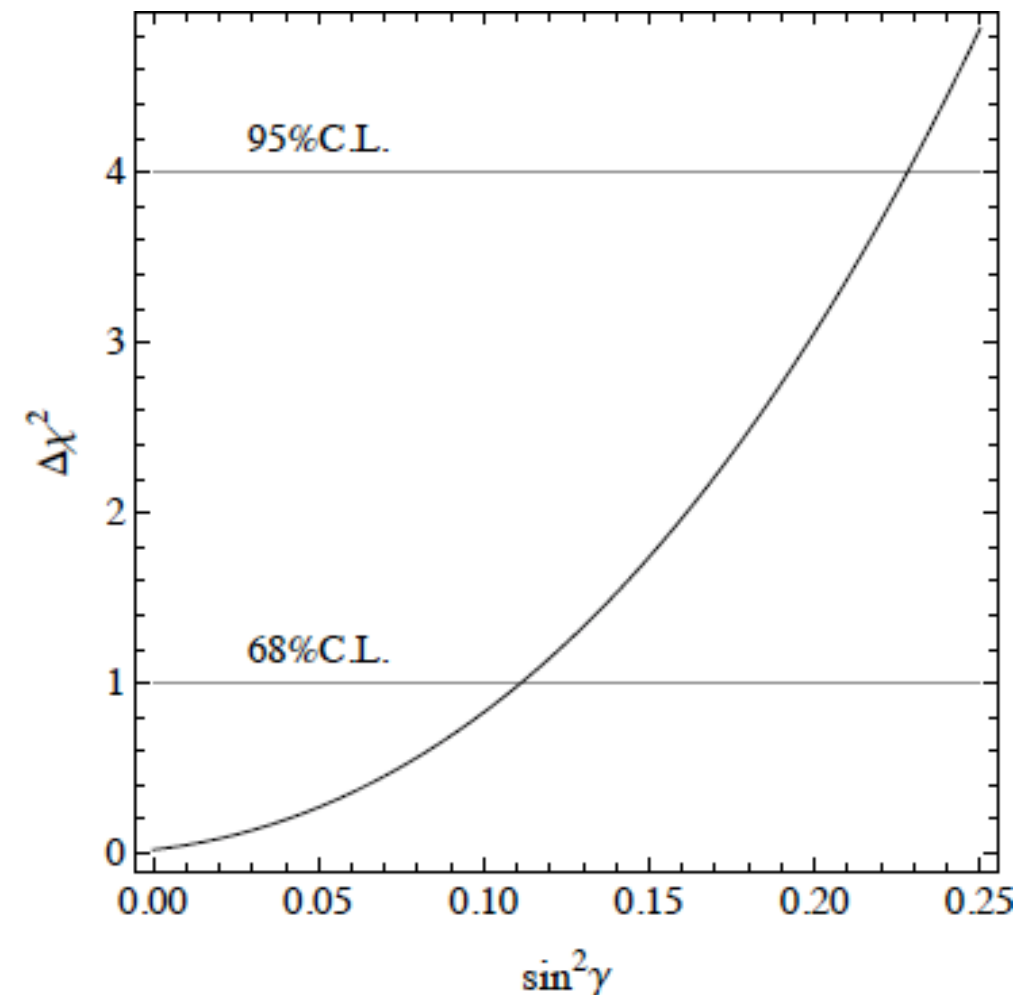
$$h_1 = c_\gamma(-s_\alpha H_d + c_\alpha H_u) + s_\gamma S$$

⇒ No susy loops nor invisible decays, like $h_1 \rightarrow \chi\chi$

95%CL on $\delta = \alpha - \beta + \pi/2$



95%CL on γ ($\delta = 0$)



Current status

$$H = s_\beta H_d - c_\beta H_u$$

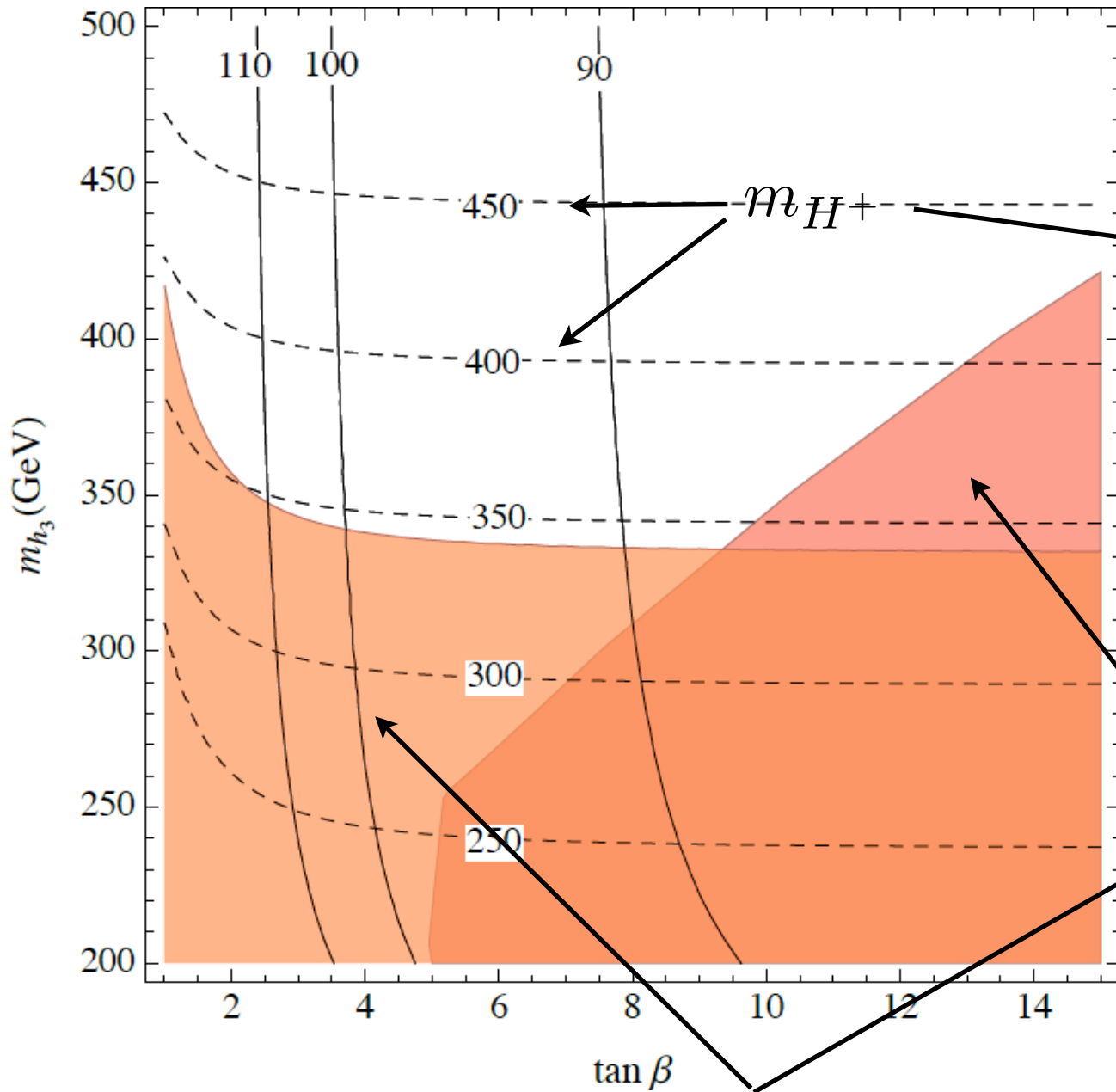
$$h = c_\beta H_d + s_\beta H_u$$

h_3

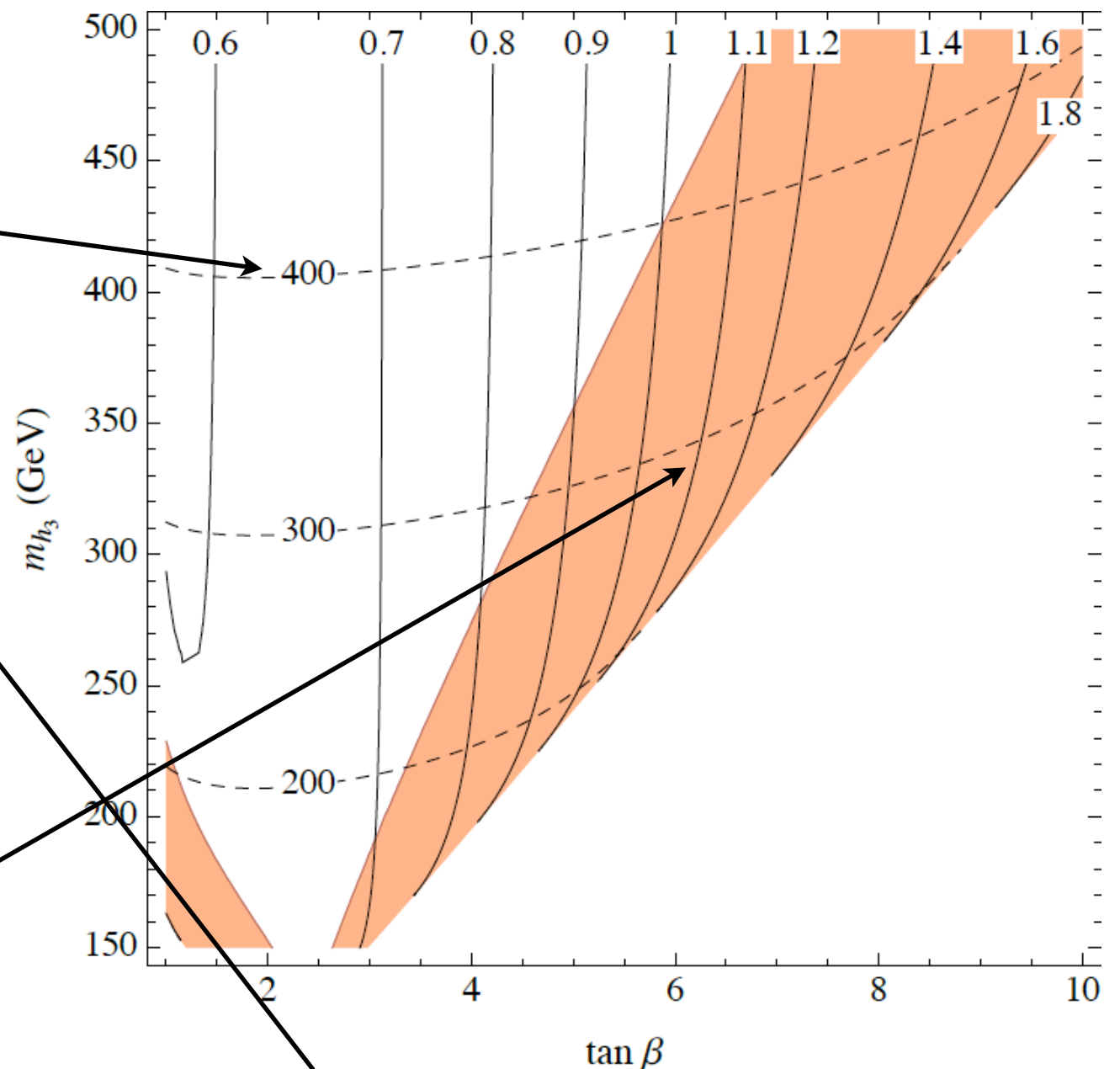
h_{LHC}

MSSM at variable Δ_t

NMSSM at variable λ

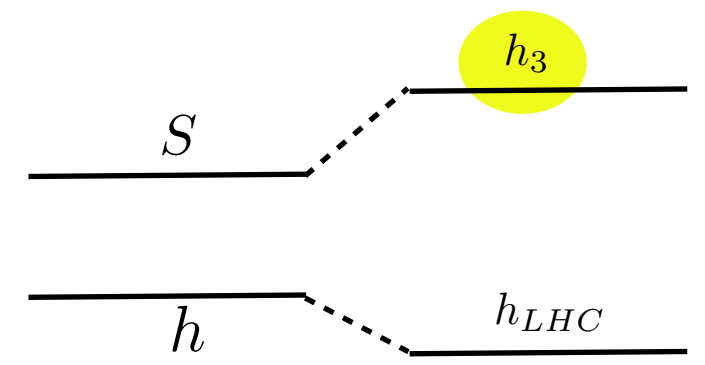


"excluded" by h_{LHC} -signal strenghts



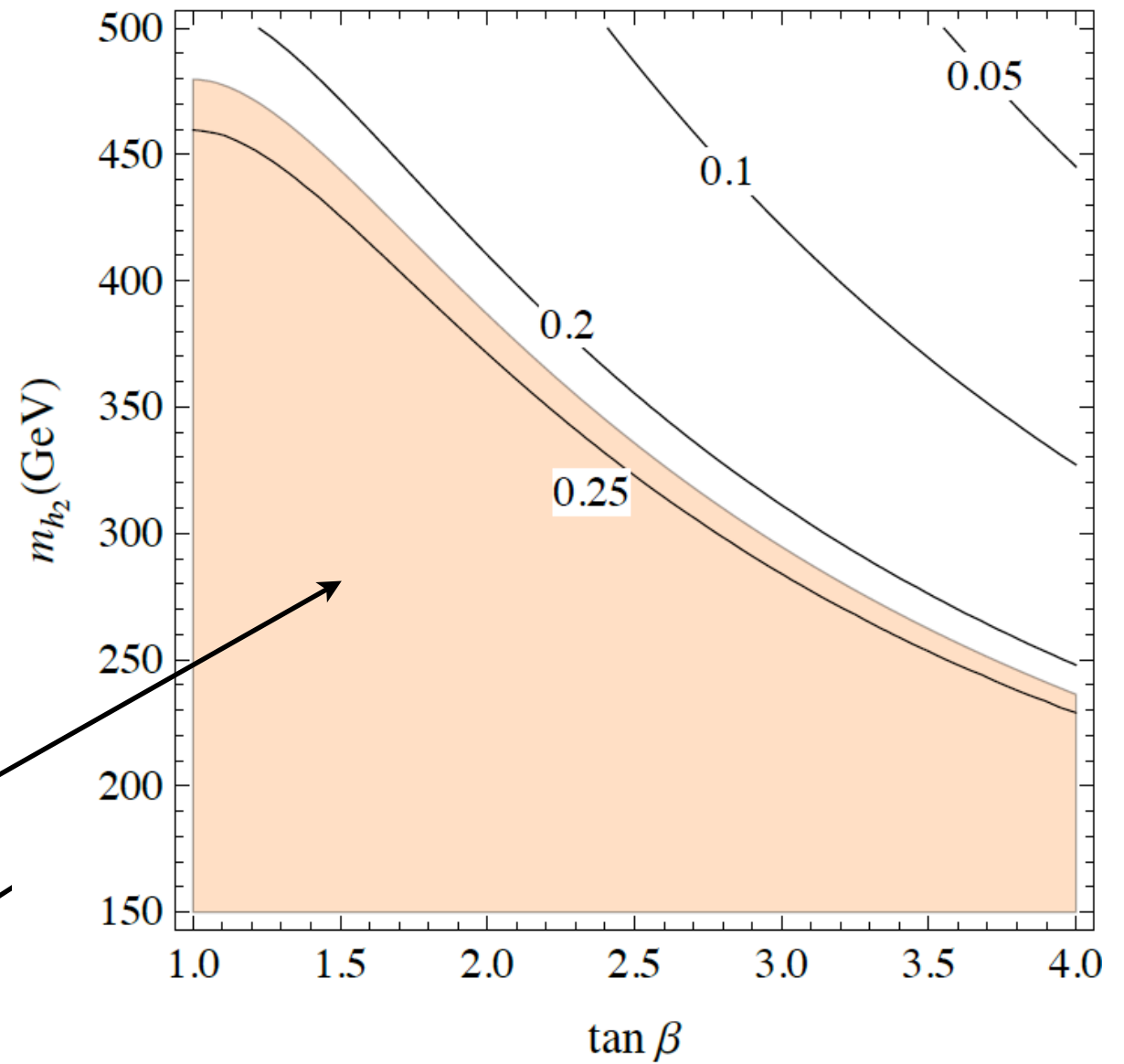
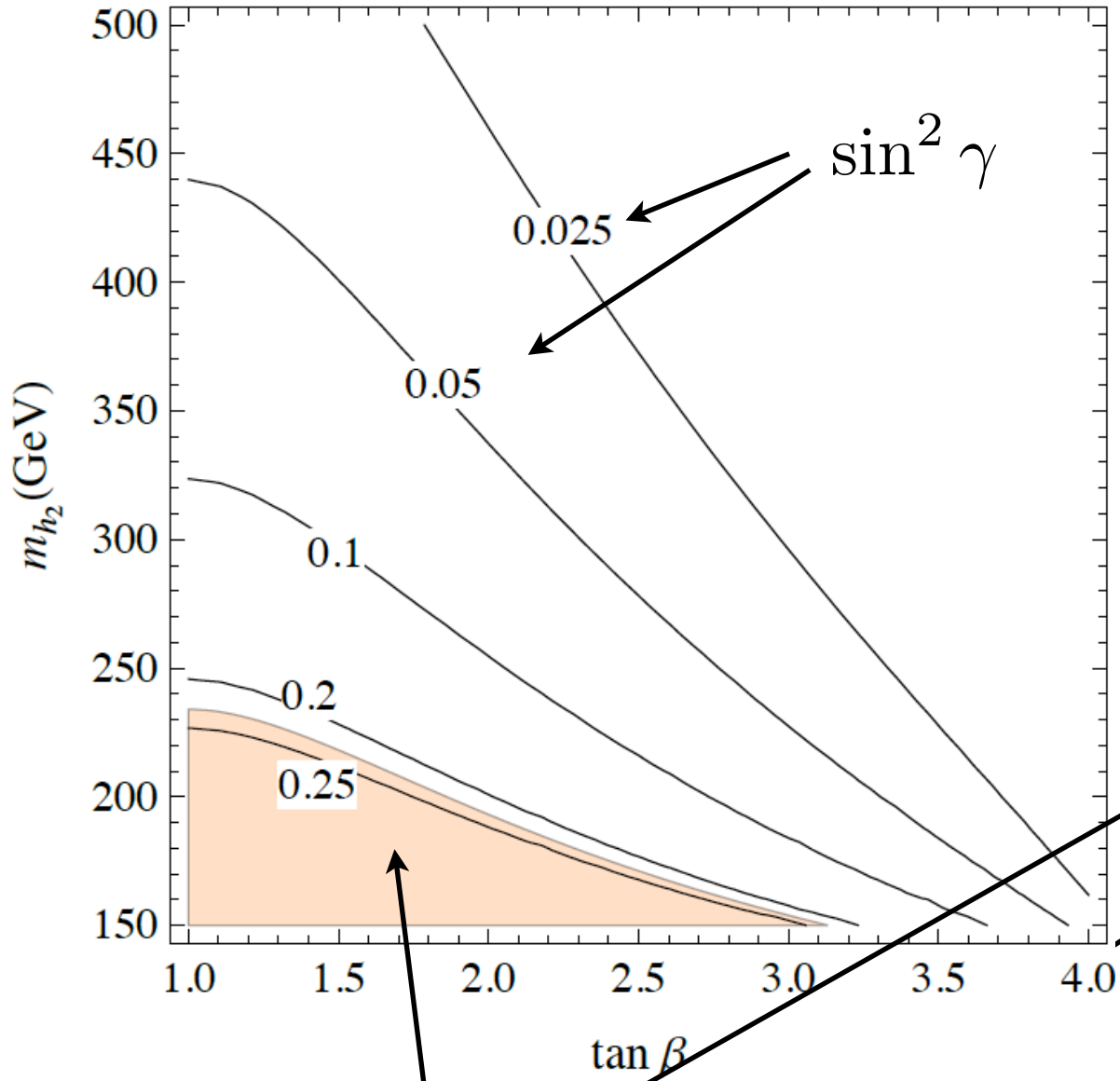
"excluded" by $h_3, A \rightarrow \tau\tau$

H-decoupled



$\lambda = 0.8$

$\lambda = 1.4$



$\Delta_t \leq 75 \text{ GeV}$ almost irrelevant

"excluded" by h_{LHC} -signal strengths

S-decoupled at LHC14

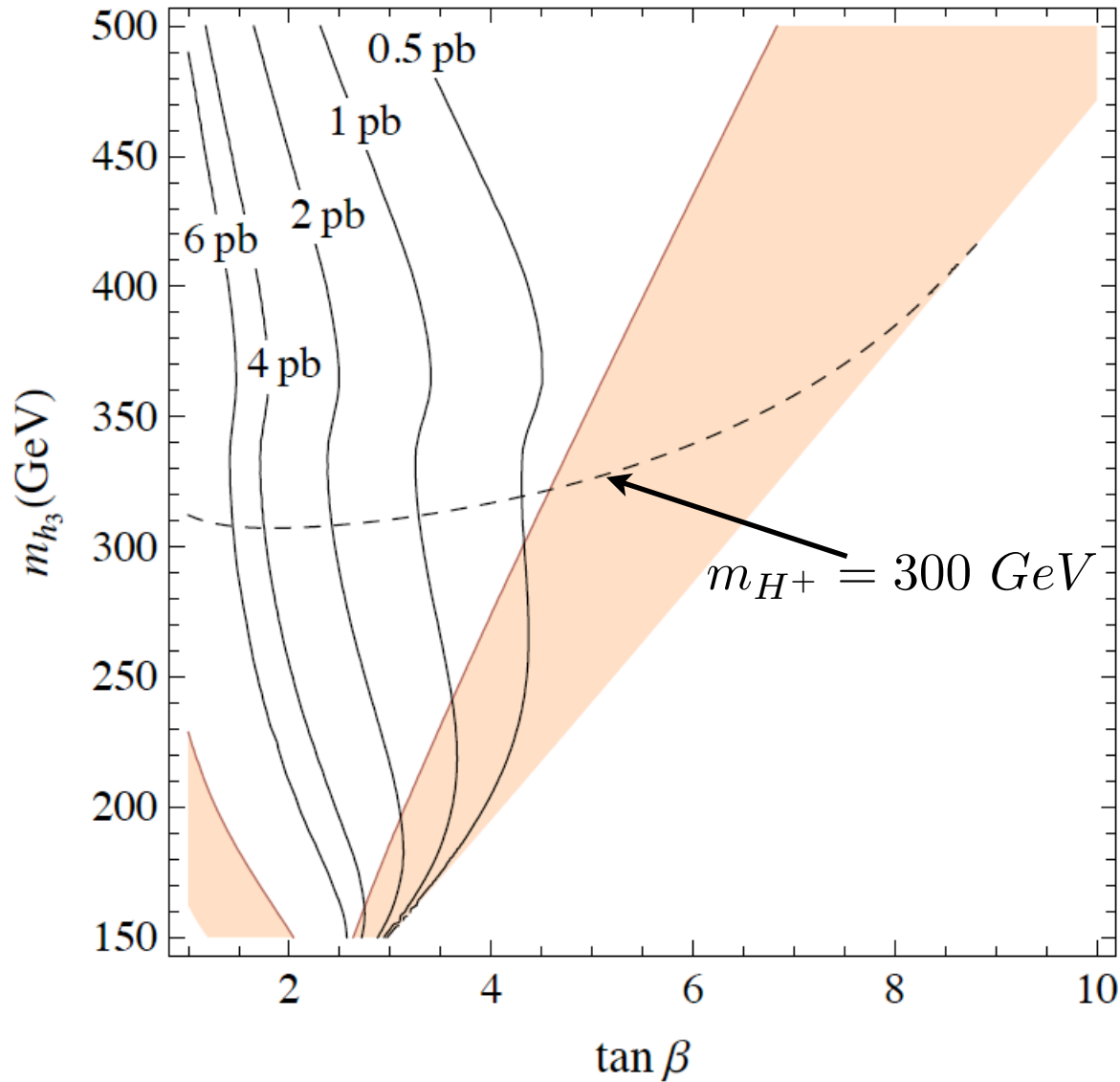
$$H = s_\beta H_d - c_\beta H_u$$

$$h = c_\beta H_d + s_\beta H_u$$

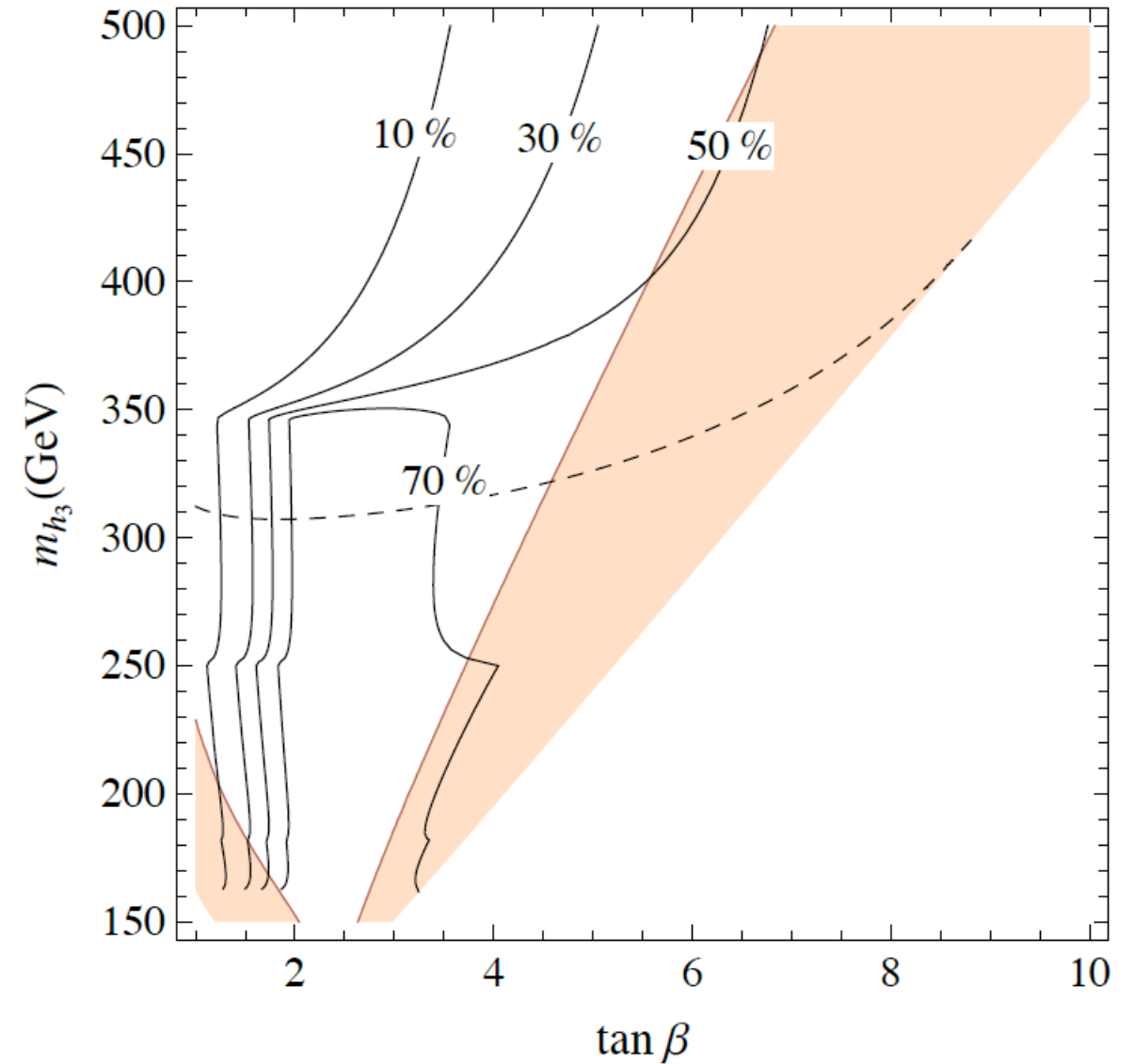
h_3

h_{LHC}

$\sigma(gg \rightarrow h_3)$



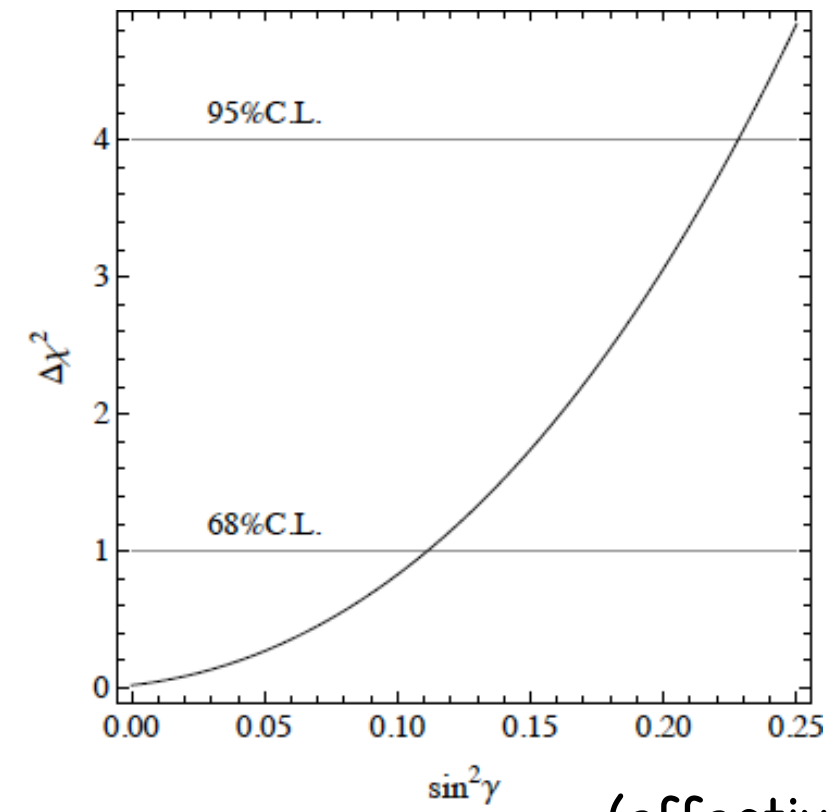
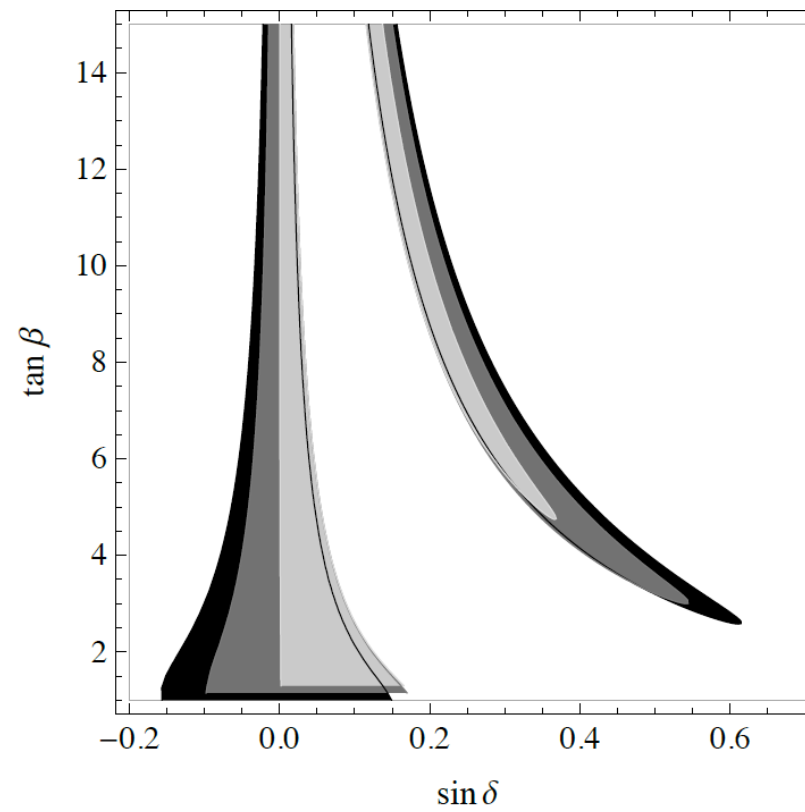
$BR(h_3 \rightarrow b\bar{b})$



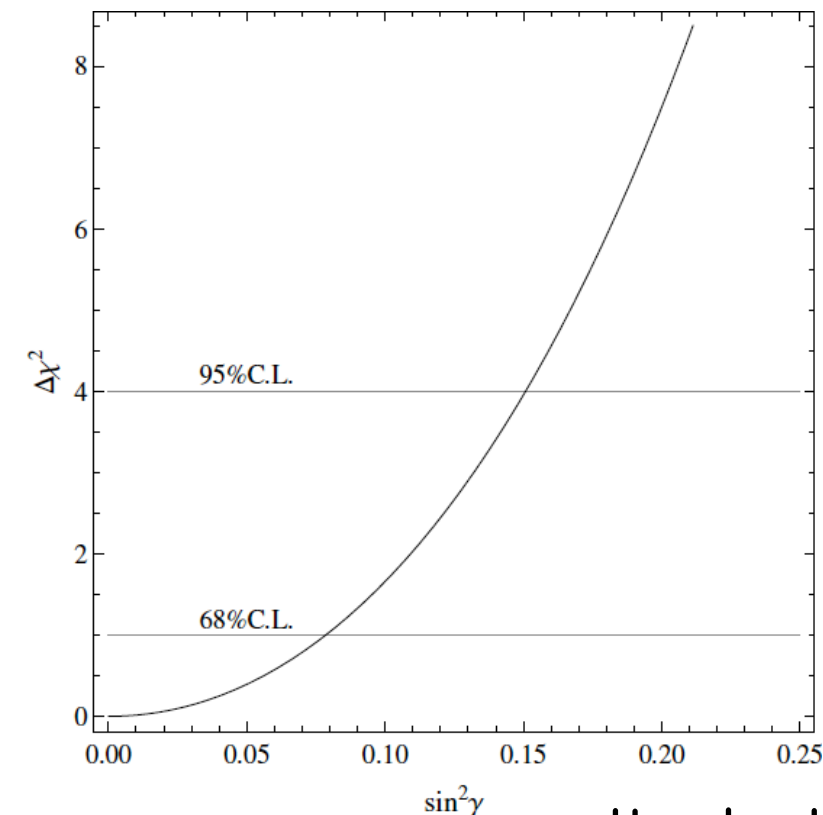
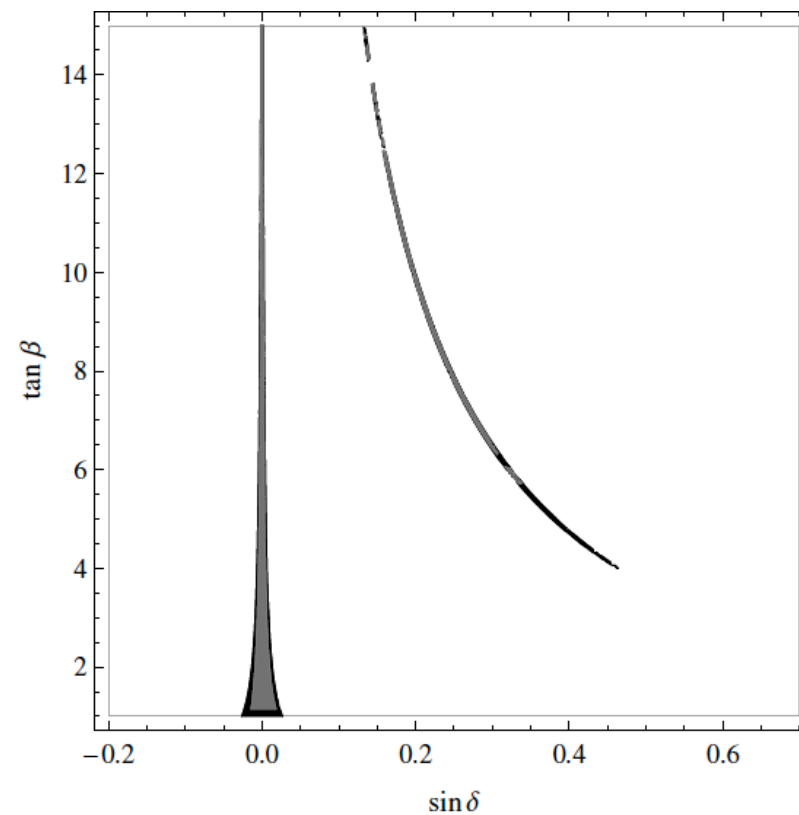
(and correspondingly $\tau\bar{\tau}$)

A projection from the measurements of the signal strengths (ATLAS and CMS preliminary) on the mixing angles

Now

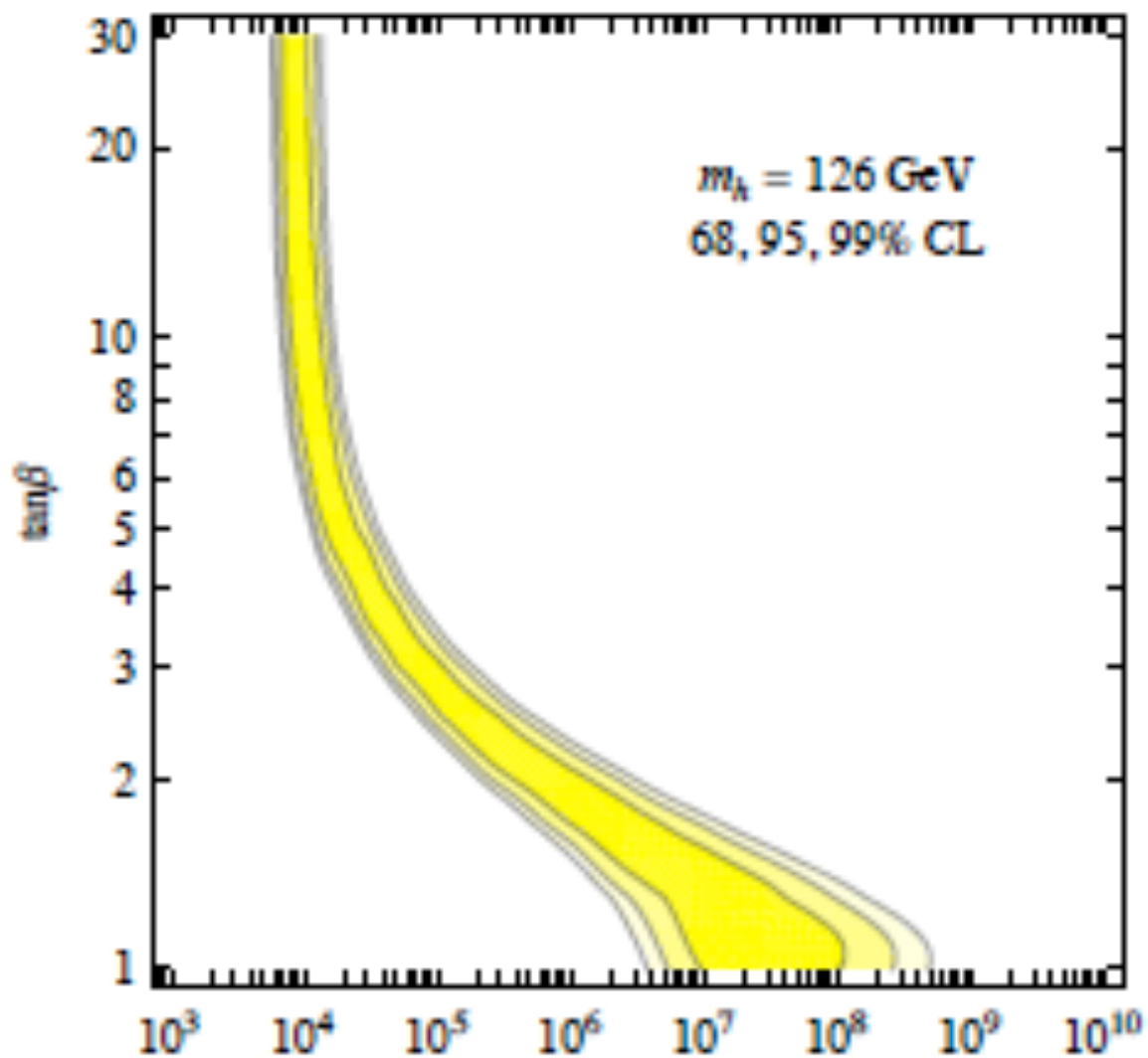


LHC14
at $300 fb^{-1}$
(central values
as in the SM)



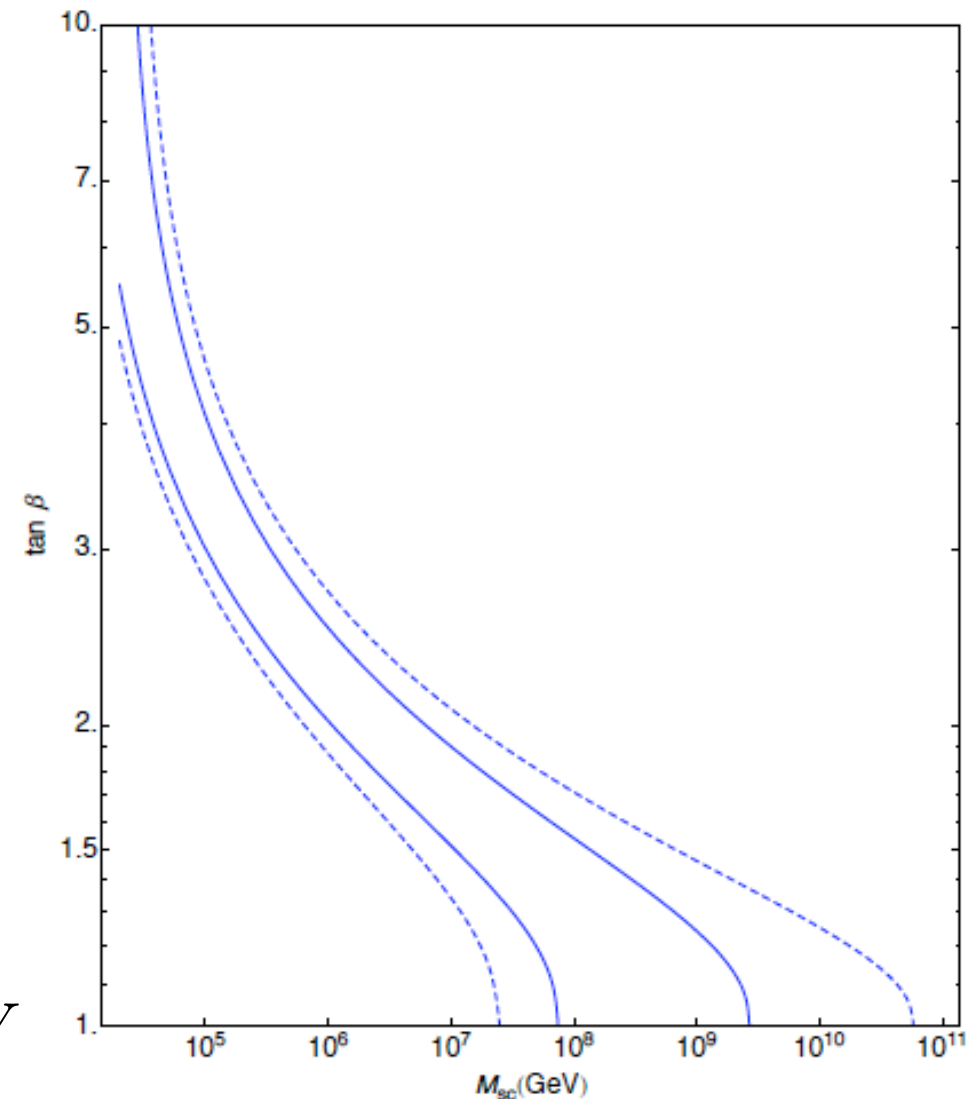
thanks to Kannike

An alternative supersymmetric view



SUSY scalars at m_S
SUSY fermions at $\sim \text{TeV}$

m_S/GeV



gauginos only at $\sim \text{TeV}$

Arkani-Hamed, Dimopoulos 2004

Giudice, Strumia 2011

Arkani-Hamed et al 2012

[If $m_S < T_R$, not to overclose the universe by a stable LSP, $m_S < 10 \div 100 \text{ TeV}$]

Hall et al, 2013

A less motivated (?) but simpler (?) picture

What can we expect from (and for) flavour physics?

$$\Delta\mathcal{L} = \sum_i \frac{1}{\Lambda_i^2} \mathcal{O}_i$$

In some cases $\Lambda_i \gtrsim 10^3 \div 10^4 \text{ TeV}$, unless some restriction operative

Is it possible that...

$$\Delta\mathcal{L} = \sum_i \frac{c_i}{\Lambda_i^2} \xi_i \mathcal{O}_i$$

with ξ_i controlled by symmetries or some dynamics and $c_i = O(1)$

and $\Lambda_i \approx 4\pi v \approx 3 \text{ TeV}$

strongly interacting EWSB

new weakly int. particle(s) at $\sim v$

Flavour \Leftrightarrow EWSB

Breaking of flavour symmetries embedded in few basic parameters

$$U(3)_Q \times U(3)_u \times U(3)_d \equiv U(3)^3$$

Chivukula, Georgi 1987 (TC)
Hall, Randall 1990 (SUSY)
D'Ambrosio et al 2002 (general)

$$Y_u = (3, \bar{3}, 1) \quad Y_d = (3, 1, \bar{3}) \quad (\text{MFV})$$

$$U(2)_Q \times U(2)_u \times U(2)_d \equiv U(2)^3$$

B, Isidori et al 2011 (general)

Y_u, Y_d split under $U(2)^3$ - representations

$$Y_u = \lambda_t \begin{pmatrix} \Delta_u & V_Q \\ V_u^T & 1 \end{pmatrix} \quad Y_d = \lambda_b \begin{pmatrix} \Delta_d & V'_Q \\ V_d^T & 1 \end{pmatrix}$$

Requiring a small breaking of $U(2)^3$: $V = V_Q \propto V'_Q$ $\|V\| = O(V_{cb})$
and, by consistency with flavour data, $\|V_u\|, \|V_d\| \ll \|V\|$

$$\left[U(3)^3 \text{ at large } \tan \beta \rightarrow U(2)^3 \quad \begin{array}{l} \text{Feldmann, Mannel 2008} \\ \text{Kagan et al 2009} \end{array} \right]$$

The $\Delta F = 2$ case

$$U(3)^3$$

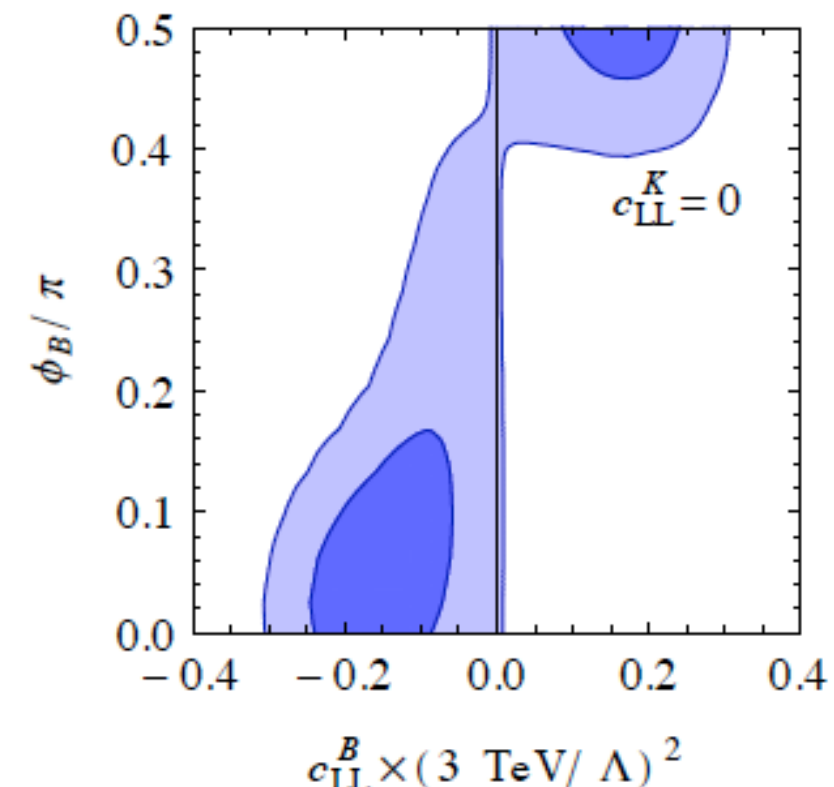
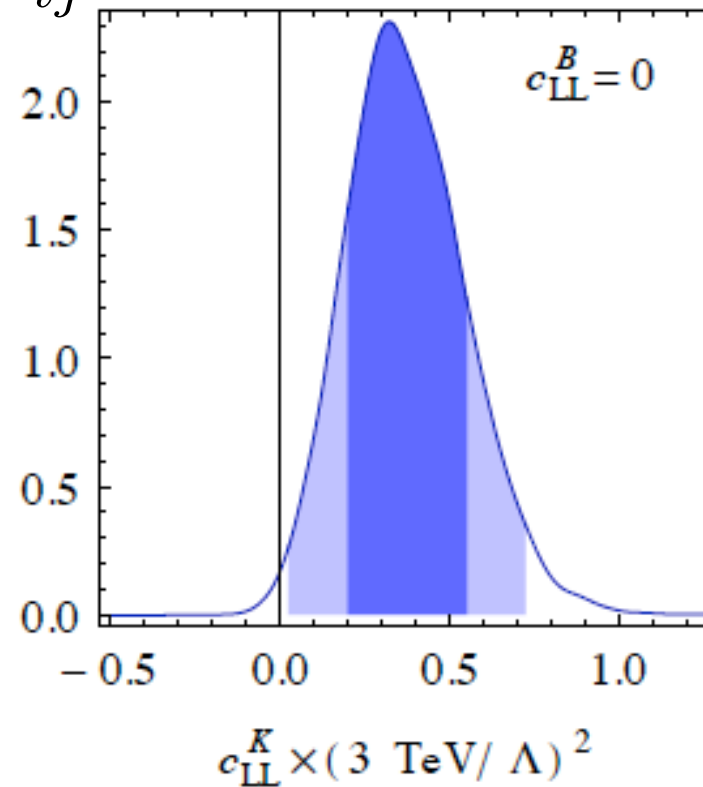
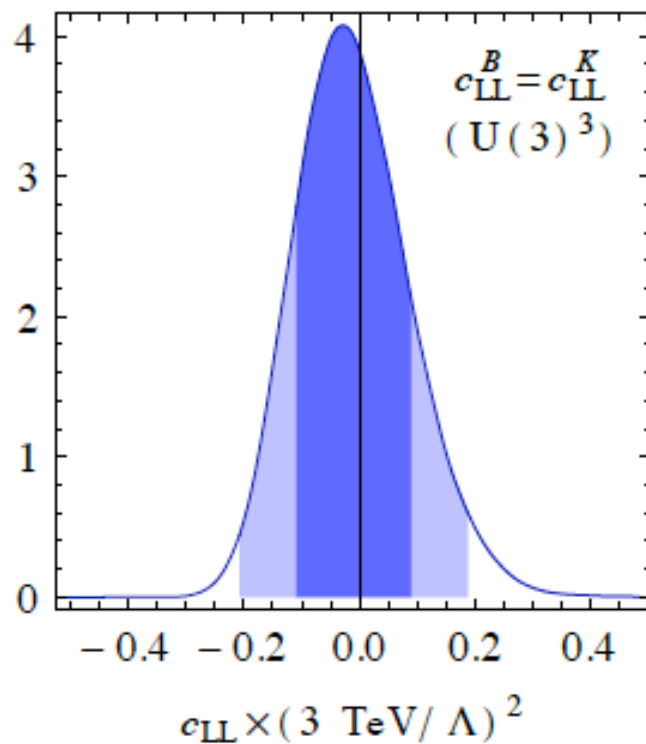
$$\frac{c_{LL}}{\Lambda^2} \xi_{ij}^2 \frac{1}{2} (\bar{d}_{Li} \gamma_\mu d_{Lj})^2$$

$$\xi_{ij} = V_{ti} V_{tj}^*$$

$$\frac{c_{LL}^K}{\Lambda^2} \xi_{ds}^2 \frac{1}{2} (\bar{d}_L \gamma_\mu s_L)^2$$

$$U(2)^3$$

$$\frac{c_{LL}^B e^{i\phi_B}}{\Lambda^2} \xi_{ib}^2 \frac{1}{2} (\bar{d}_{Li} \gamma_\mu b_L)^2$$



(cannot fit the "discrepancy")

B, Buttazzo et al 2012

Flavour tests
versus direct searches
(cum grano salis)

for $c = 1$ $\Lambda \approx 4\pi(m, f)$

E.g. $c \cdot (3 \text{ TeV}/\Lambda)^2 \approx 0.1$ means $m, f \approx 0.8 \text{ TeV}$

$\Delta F = 1$ Summary

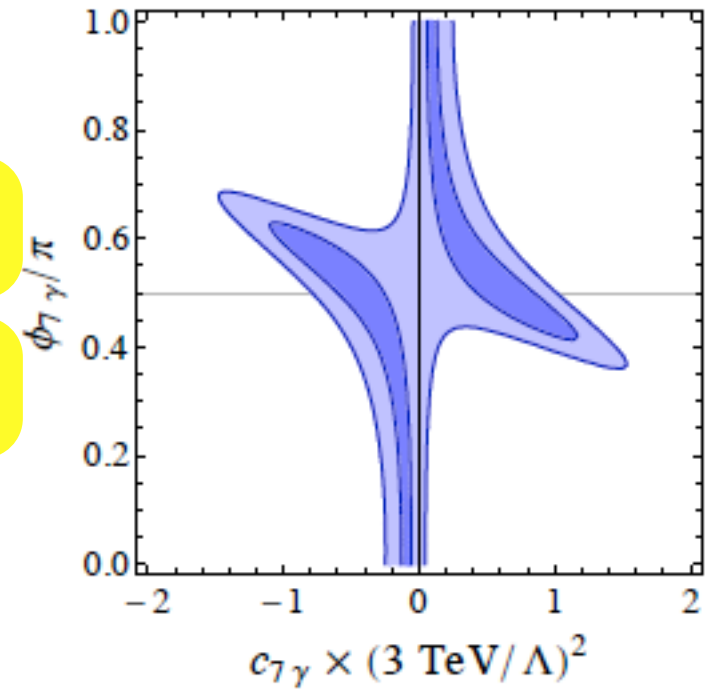
Chirality breaking
(cromo-)magnetic operators

$$B \rightarrow X_{(s,d)} \gamma$$

$$B \rightarrow K(\pi) \mu \mu$$

$U(3)^3$

$U(2)^3$



Chirality conserving op.s

$$B \rightarrow X_{(s,d)} \gamma$$

$$B \rightarrow K(\pi) \mu \mu$$

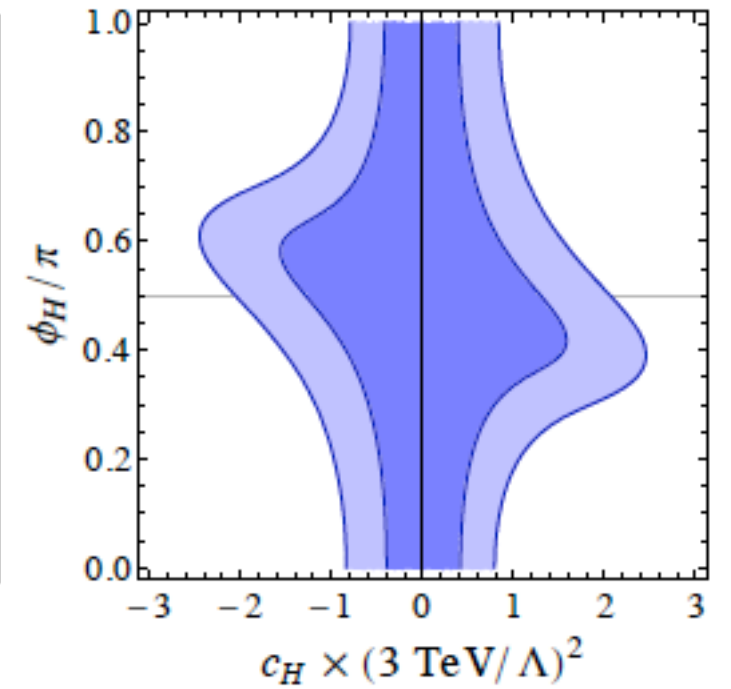
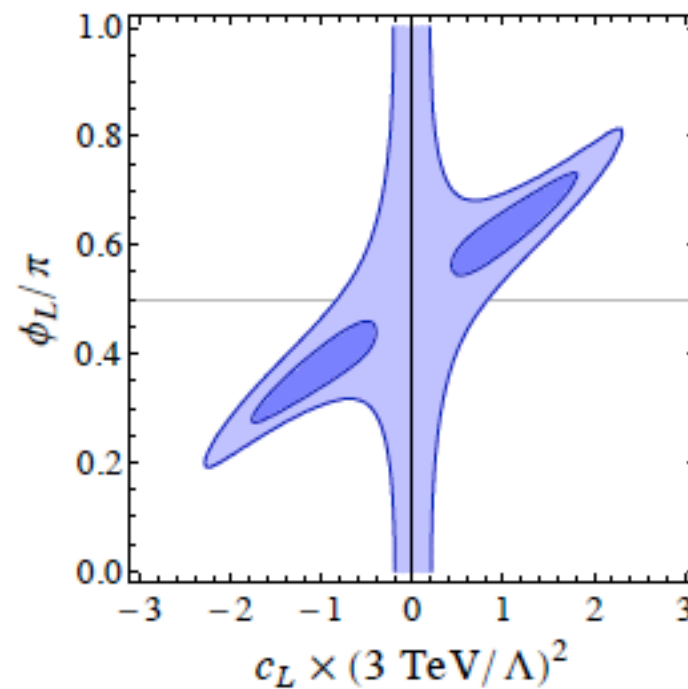
$$B_s \rightarrow \mu \mu$$

$$[K \rightarrow \pi \nu \nu]$$

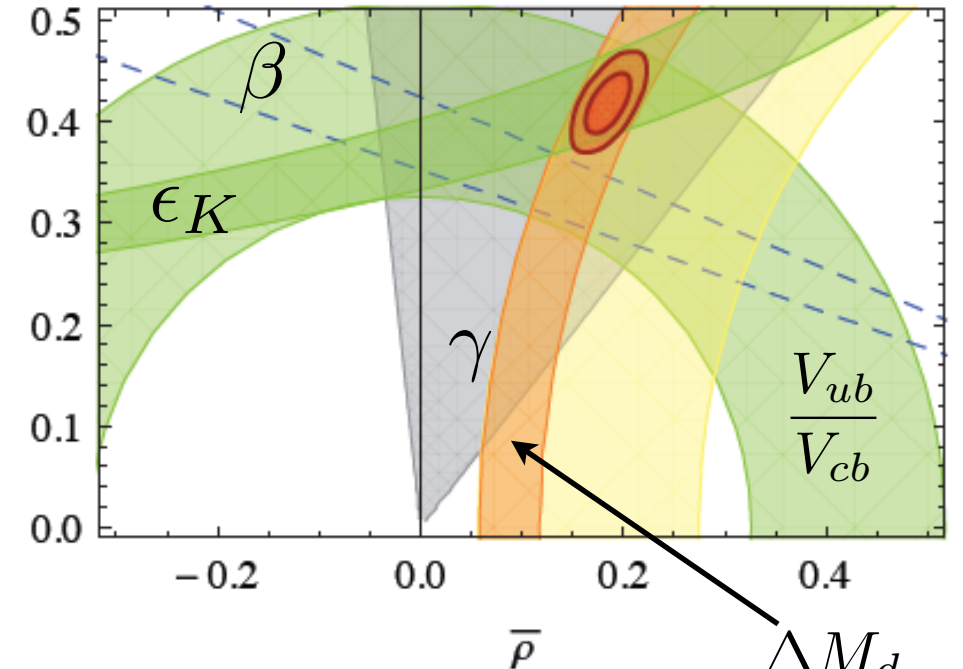
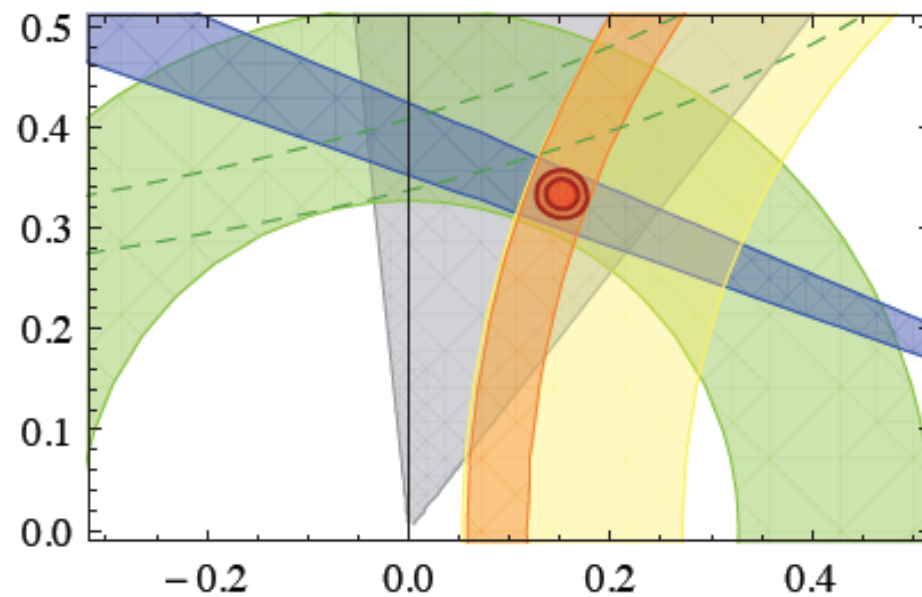
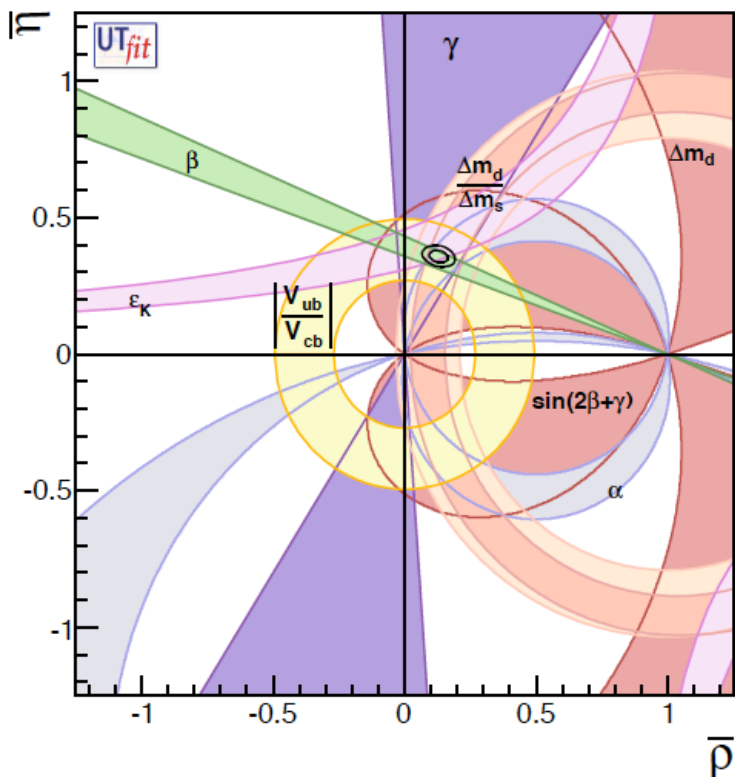
$U(2)^3$

correlated

no phase in $U(3)^3$



$\Delta F = 2$ key measurements



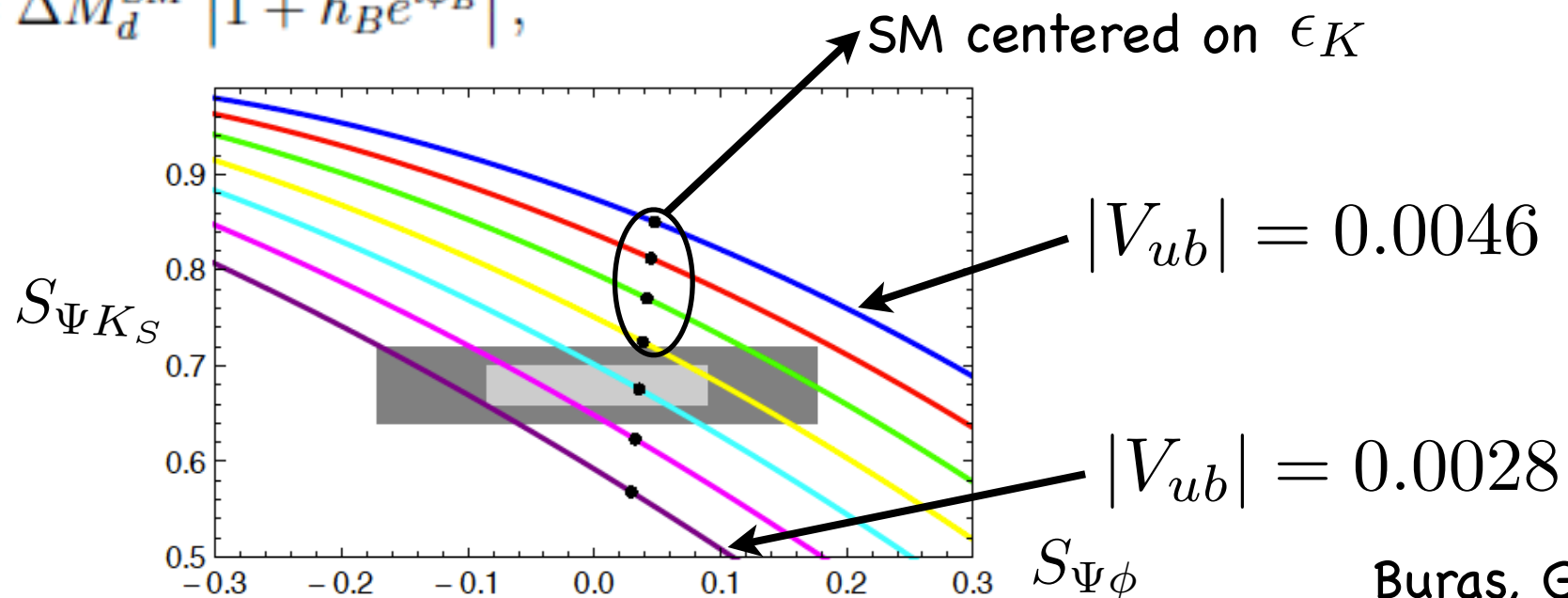
$U(2)^3$

$$\begin{aligned} \epsilon_K &= \epsilon_K^{\text{SM}(tt)} (1 + h_K) + \epsilon_K^{\text{SM}(tc+cc)}, \\ S_{\psi K_S} &= \sin(2\beta + \arg(1 + h_{BE} e^{i\phi_B})), \\ S_{\psi\phi} &= \sin(2|\beta_s| - \arg(1 + h_{BE} e^{i\phi_B})), \\ \Delta M_d &= \Delta M_d^{\text{SM}} |1 + h_{BE} e^{i\phi_B}|, \end{aligned}$$

$$\begin{aligned} \frac{\Delta M_d}{\Delta M_s} &= \frac{\Delta M_d}{\Delta M_s} \Big|_{\text{SM}} = 34.5 \pm 3.0 \\ \frac{\Delta M_d}{\Delta M_s} \Big|_{\text{exp}} &= 35.0 \pm 0.3 \end{aligned}$$

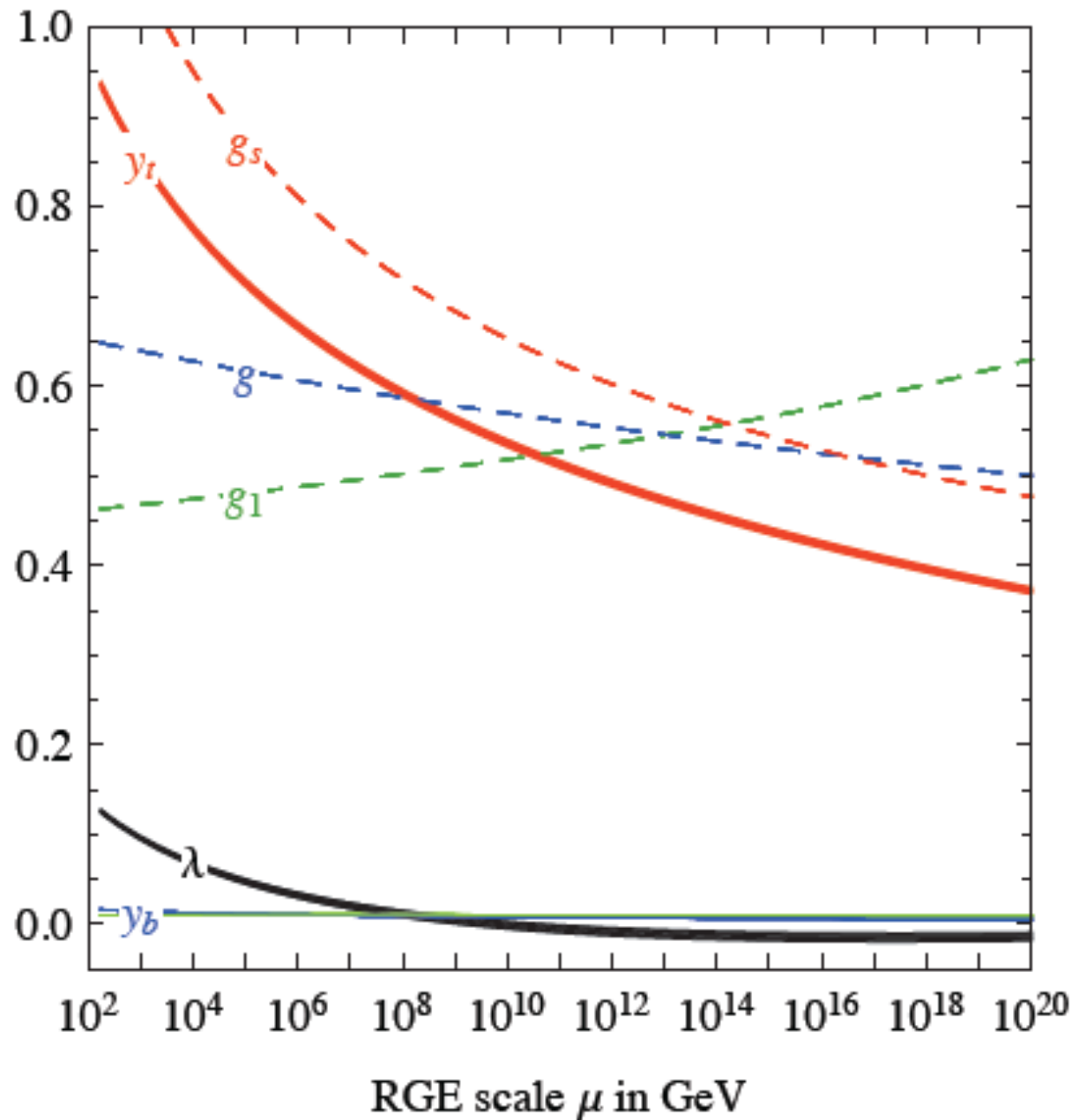
$$\Rightarrow \gamma \approx 70^\circ$$

The key role of V_{ub} and $S_{\Psi\phi}$ as well as of $F_{B_{d,s}} (B_{d,s})^{1/2}$ from the lattice

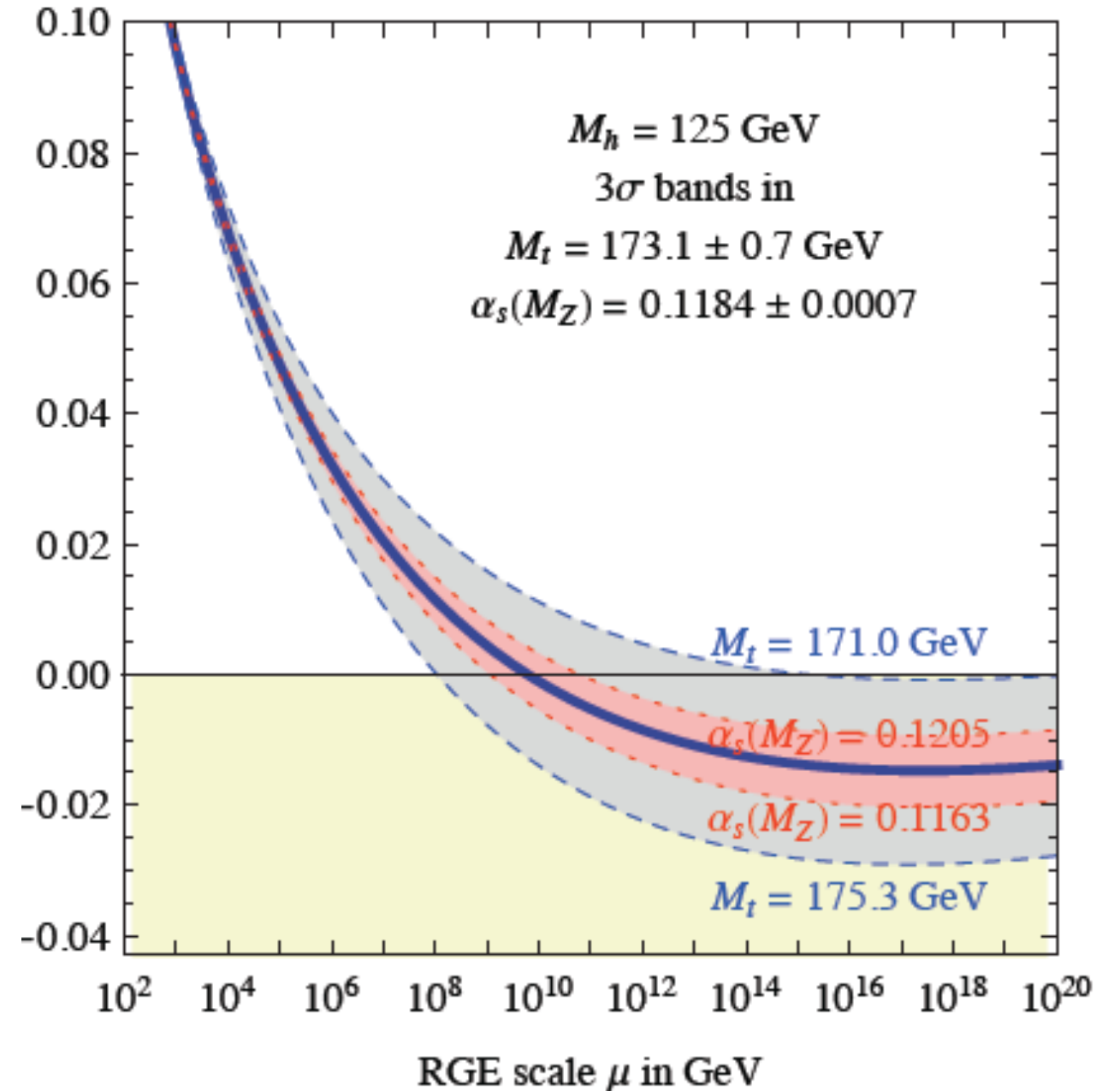


What if the Higgs boson likes to be unnatural and the SM is unchanged up to very high energies?

largest couplings



Higgs self-coupling

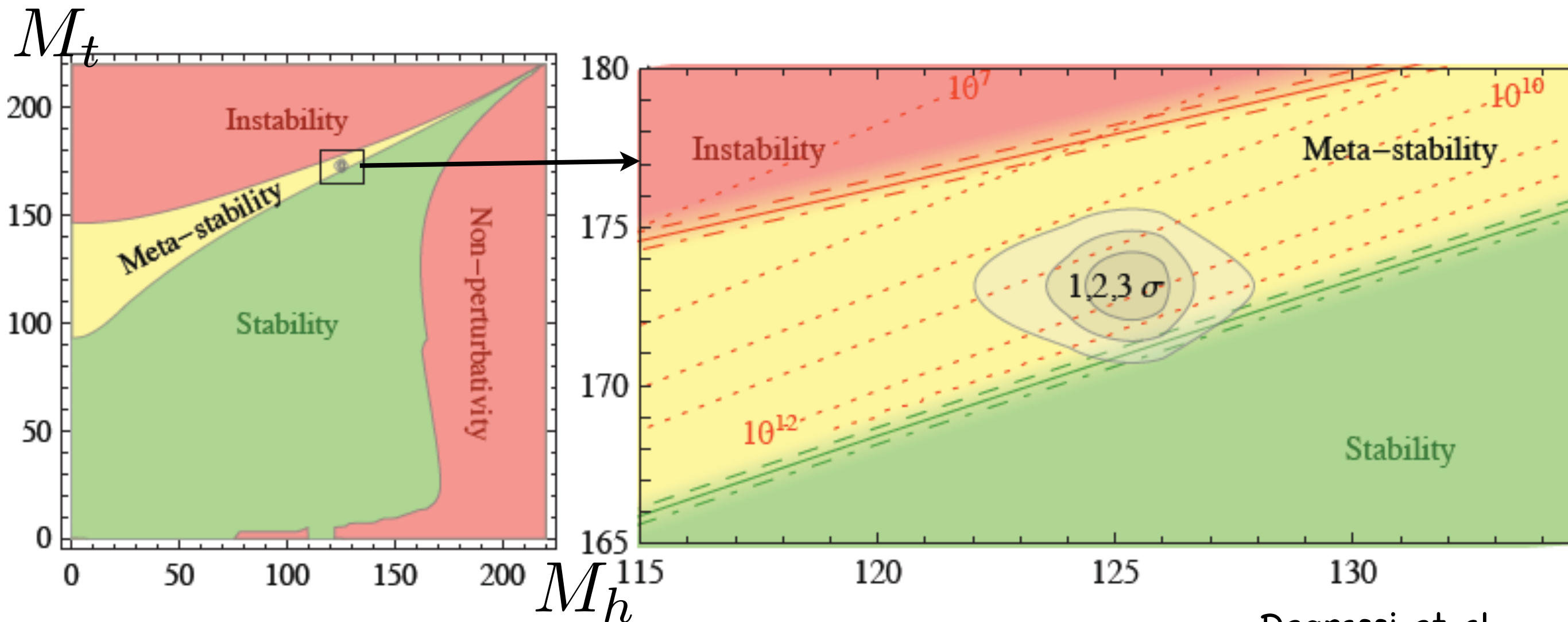


Degrassi et al 2012
 Chetyrkin, Zoller 2005
 Bezrukov et al 2005

A special meaning for $\lambda \approx 0$ at M_{Pl} ?

Assume SM unchanged up to M_{Pl}

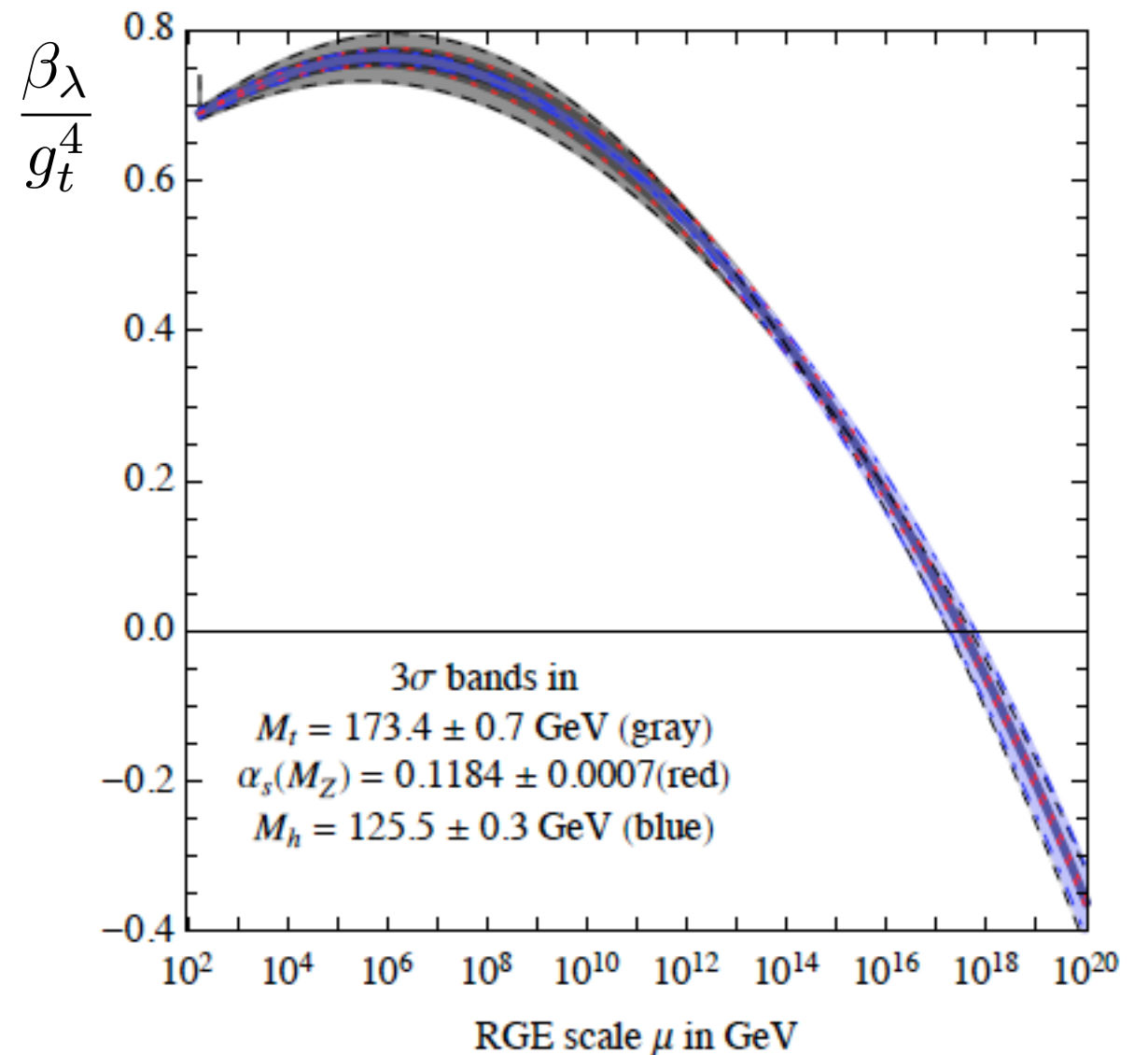
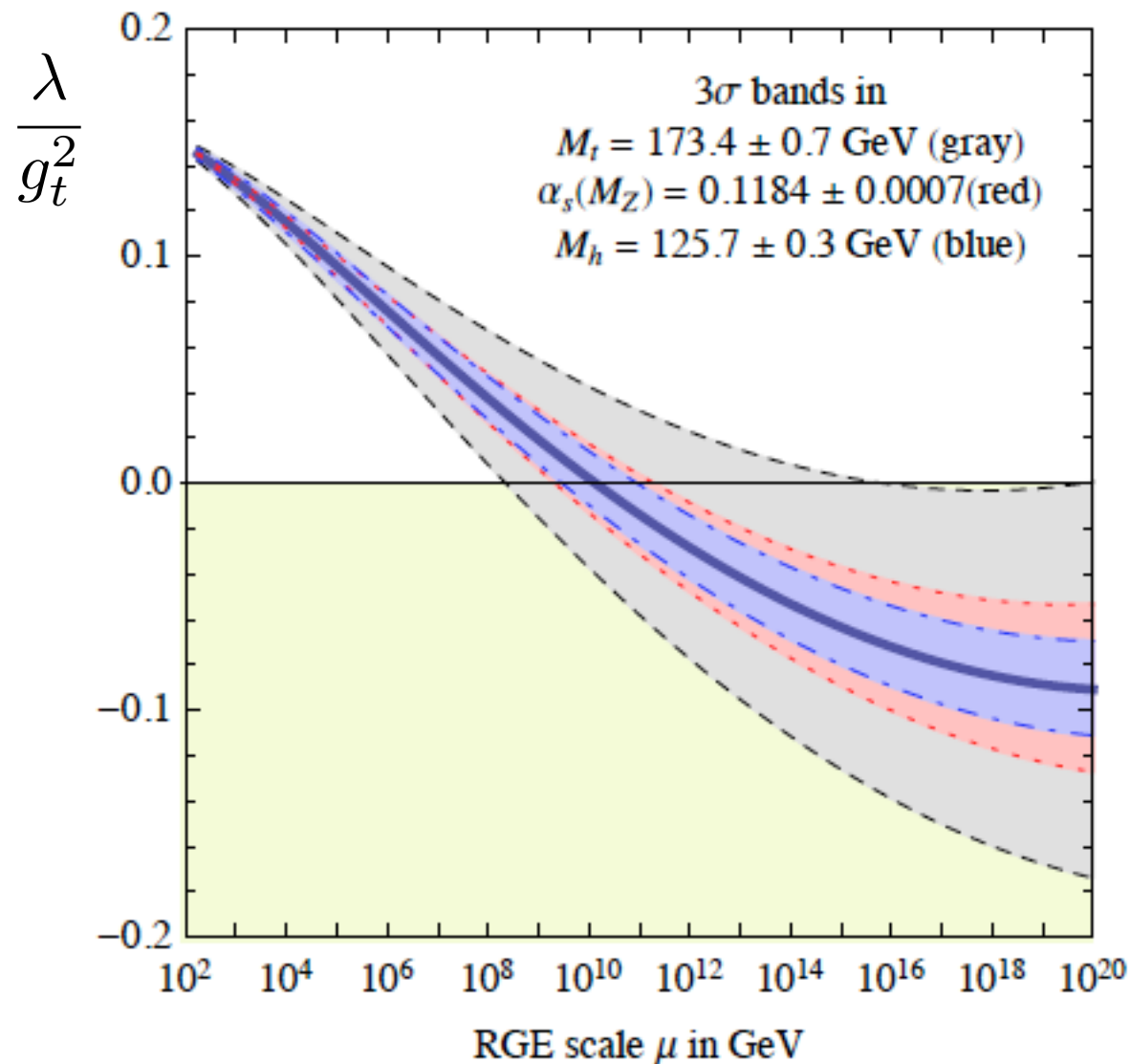
Nielsen et al 1988
Shaposhnikov et al 2009



Absolute stability at $M_{Pl}(\lambda(M_{Pl}) \gtrsim 0)$ not quite achieved for current "best" values of M_t and M_h

Speculations about possible meaning of all this not lacking (anthropic pressure, ...)

What's the real evidence for $\lambda(M_{Pl}) \approx \beta_\lambda(M_{Pl}) = 0$?



Even if this improved (g_t , etc) how shall we know that it is not a coincidence?

Thanks to Rattazzi and Strumia