

# Particules Elementaires, Gravitation et Cosmologie 2007-2008

## Le Modele Standard et ses extensions

Cours XII: 14 mars 2008

## Neutrino Masses, Mixing and Oscillations part 2: the theory

Ferruccio Feruglio  
Universita' di Padova

# Beyond the Standard Model

a non-vanishing neutrino mass is the **first evidence of the incompleteness of the Standard Model [SM]**

in the SM neutrinos belong to SU(2) doublets with hypercharge  $Y=-1/2$   
they have only two helicities (not four, as the other charged fermions)

$$l = \begin{pmatrix} \nu_e \\ e \end{pmatrix} = (1, 2, -1/2)$$

the requirement of invariance under the gauge group  $G=SU(3)\times SU(2)\times U(1)$  forbids pure fermion mass terms in the lagrangian. Charged fermion masses arise, after electroweak symmetry breaking, through gauge-invariant Yukawa interactions

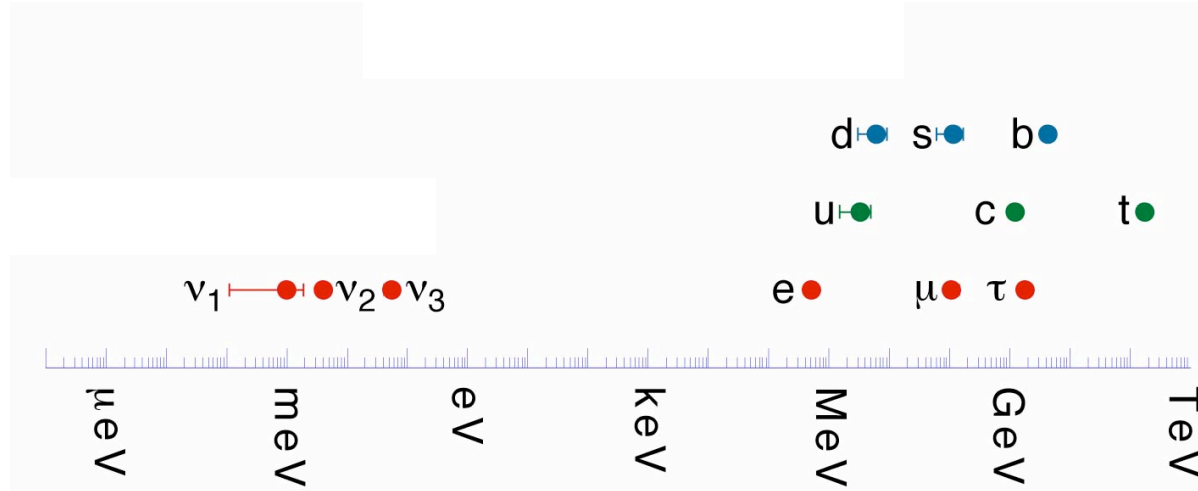
$$\Phi \underbrace{\Psi\Psi'}_{\text{same helicity}}$$

not even this term is allowed for SM neutrinos, by gauge invariance

# Questions

how to extend the SM in order to accommodate neutrino masses?

why neutrino masses are so small, compared with the charged fermion masses?



why lepton mixing angle are so different from those of the quark sector?

$$U_{PMNS} = \begin{pmatrix} \frac{2}{\sqrt{6}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & -\frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix} + \text{corrections}$$

$$V_{CKM} \approx \begin{pmatrix} 1 & O(\lambda) & O(\lambda^4 \div \lambda^3) \\ O(\lambda) & 1 & O(\lambda^2) \\ O(\lambda^4 \div \lambda^3) & O(\lambda^2) & 1 \end{pmatrix}$$

$\lambda \approx 0.22$



# First possibility: modify (1), the particle content

there are several possibilities

one of the simplest one is to mimic the charged fermion sector

**Example 1** {

- add (three copies of) right-handed neutrinos  $\nu^c \equiv (1,1,0)$  full singlet under  $G=SU(3)\times SU(2)\times U(1)$
- ask for (global) invariance under B-L (no more automatically conserved as in the SM)

the neutrino has now four helicities, as the other charged fermions, and we can build gauge invariant Yukawa interactions giving rise, after electroweak symmetry breaking, to neutrino masses

$$L_Y = d^c y_d (\Phi^+ q) + u^c y_u (\tilde{\Phi}^+ q) + e^c y_e (\Phi^+ l) + \nu^c y_\nu (\tilde{\Phi}^+ l) + h.c.$$

$$m_f = \frac{y_f}{\sqrt{2}} v \quad f = u, d, e, \nu$$

with three generations there is an exact replica of the quark sector and, after diagonalization of the charged lepton and neutrino mass matrices, a mixing matrix  $U$  appears in the charged current interactions

$$-\frac{g}{\sqrt{2}} W_\mu^- \bar{e} \sigma^\mu U_{PMNS} \nu + h.c. \quad U_{PMNS} \text{ has three mixing angles and one phase, like } V_{CKM}$$

## a generic problem of this approach

the particle content can be modified in several different ways in order to account for non-vanishing neutrino masses (additional right-handed neutrinos, new  $SU(2)$  fermion triplets, additional  $SU(2)$  scalar triplet(s), SUSY particles,...). Which is the correct one?

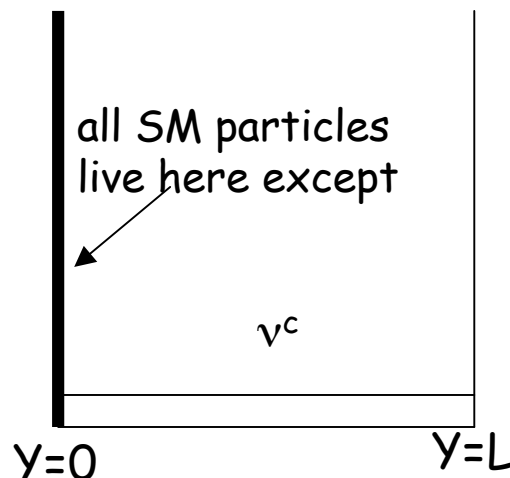
## a problem of the above example

if neutrinos are so similar to the other fermions, why are so light?

$$\frac{y_\nu}{y_{top}} \leq 10^{-12}$$

Quite a speculative answer:

neutrinos are so light, because the right-handed neutrinos have access to an extra (fifth) spatial dimension



neutrino Yukawa coupling

$$\begin{aligned} \nu^c(y=0)(\tilde{\Phi}^+ l) &= \text{Fourier expansion} \\ &= \frac{1}{\sqrt{L}} \nu_0^c(\tilde{\Phi}^+ l) + \dots \quad [\text{higher modes}] \end{aligned}$$

if  $L \ll 1$  (in units of the fundamental scale)  
then neutrino Yukawa coupling is suppressed

# Second possibility: abandon (2) renormalizability

A disaster?

$$L = L_{d \leq 4}^{SM} + \frac{L_5}{\Lambda} + \frac{L_6}{\Lambda^2} + \dots$$

a new scale  $\Lambda$  enters the theory. The new (gauge invariant!) operators  $L_5, L_6, \dots$  contribute to amplitudes for physical processes with terms of the type

$$\frac{L_5}{\Lambda} \rightarrow \frac{E}{\Lambda} \quad \frac{L_6}{\Lambda^2} \rightarrow \left(\frac{E}{\Lambda}\right)^2 \quad \dots$$

the theory cannot be extrapolated beyond a certain energy scale  $E \approx \Lambda$ .  
[at variance with a renormalizable (asymptotically free) QFT]

If  $E \ll \Lambda$  (for example  $E$  close to the electroweak scale,  $10^2 \text{ GeV}$ , and  $\Lambda \approx 10^{15} \text{ GeV}$  not far from the so-called Grand Unified scale), the above effects will be tiny and, the theory will *look like* a renormalizable theory!

$$\frac{E}{\Lambda} \approx \frac{10^2 \text{ GeV}}{10^{15} \text{ GeV}} = 10^{-13}$$

an extremely tiny effect, but exactly what needed to suppress  $m_\nu$  compared to  $m_{\text{top}}$ !

**Worth to explore.** The dominant operators (suppressed by a single power of  $1/\Lambda$ ) beyond  $L_{SM}$  are those of dimension 5. Here is a list of all d=5 gauge invariant operators

$$\frac{L_5}{\Lambda} = \frac{(\tilde{\Phi}^+ l)(\tilde{\Phi}^+ l)}{\Lambda} = \frac{v}{2} \left( \frac{v}{\Lambda} \right) \nu \nu + \dots$$

a unique operator!  
[up to flavour combinations]  
it violates (B-L) by two units

it is suppressed by a factor  $(v/\Lambda)$   
with respect to the neutrino mass term  
of Example 1:

$$\nu^c (\tilde{\Phi}^+ l) = \frac{v}{\sqrt{2}} \nu^c \nu + \dots$$

**it provides an explanation for the smallness of  $m_\nu$ :**  
the neutrino masses are small because the scale  $\Lambda$ , characterizing (B-L) violations, is very large. How large? Up to about  $10^{15}$  GeV

from this point of view neutrinos offer a unique window on physics at very large scales, inaccessible in present (and probably future) man-made experiments.

since this is the dominant operator in the expansion of  $L$  in powers of  $1/\Lambda$ , we could have expected to find the first effect of physics beyond the SM in neutrinos ... and indeed this was the case!



# $L_5$ represents the effective, low-energy description of several extensions of the SM

Example 2:  
see-saw

add (three copies of)  $\nu^c \equiv (1,1,0)$

full singlet under  
 $G = SU(3) \times SU(2) \times U(1)$

this is like Example 1, but without enforcing (B-L) conservation

$$L(\nu^c, l) = \nu^c y_\nu (\tilde{\Phi}^+ l) + \frac{1}{2} \nu^c M \nu^c + h.c.$$

mass term for right-handed neutrinos:  $G$  invariant, violates (B-L) by two units.

the new mass parameter  $M$  is independent from the electroweak breaking scale  $v$ . If  $M \gg v$ , we might be interested in an effective description valid for energies much smaller than  $M$ . This is obtained by "integrating out" the field  $\nu^c$

$$L_{eff}(l) = -\frac{1}{2} (\tilde{\Phi}^+ l) \left[ y_\nu^T M^{-1} y_\nu \right] (\tilde{\Phi}^+ l) + h.c. + \dots$$

terms suppressed by more powers of  $M^{-1}$

this reproduces  $L_5$ , with  $M$  playing the role of  $\Lambda$ . This particular mechanism is called (type I) **see-saw**.

# Theoretical motivations for the see-saw

$\Lambda \approx 10^{15}$  GeV is very close to the so-called unification scale  $M_{GUT}$ .

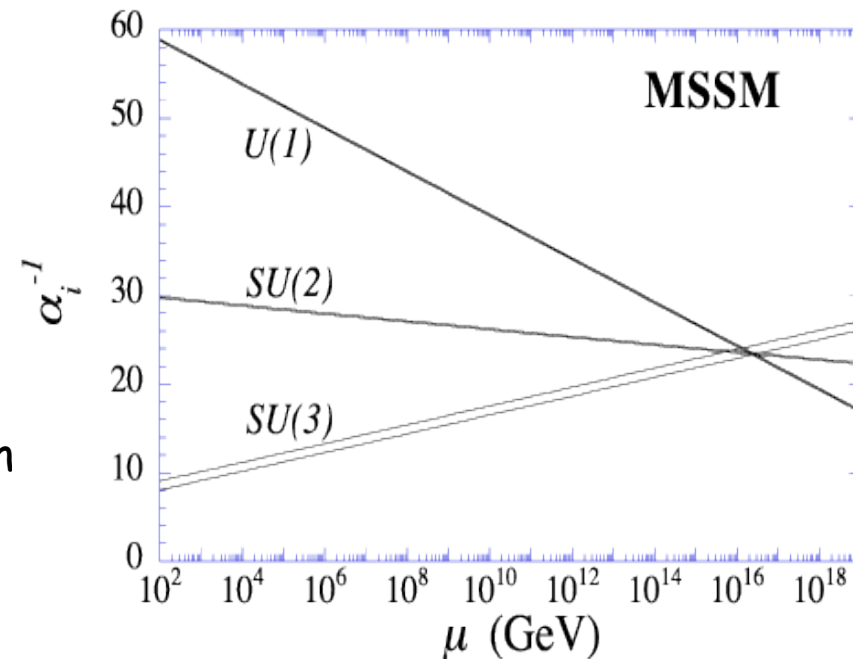
an independent evidence for  $M_{GUT}$  comes from the **unification of the gauge coupling constants** in (SUSY extensions of) the SM.

such unification is a generic prediction of **Grand Unified Theories (GUTs)**: the SM gauge group  $G$  is embedded into a simple group such as  $SU(5)$ ,  $SO(10)$ ,...

**Particle classification**: it is possible to unify all SM fermions (1 generation) into a single irreducible representation of the GUT gauge group. Simplest example:  $G_{GUT} = SO(10)$

$$16 = (q, d^c, u^c, l, e^c, \nu^c) \text{ a whole family plus a right-handed neutrino!}$$

quite a fascinating possibility. Unfortunately, it still lacks experimental tests. In GUT new, very heavy, particles can convert quarks into leptons and the **proton is no more a stable particle**. Proton decay rates and decay channels are however model dependent. Experimentally we have only lower bounds on the proton lifetime.



## 2 additional virtues of the see-saw

The see-saw mechanism can enhance small mixing angles into large ones

$$m_\nu = -\left[ y_\nu^T M^{-1} y_\nu \right] \nu^2$$

Example with 2 generations

$$y_\nu = \begin{pmatrix} \delta & \delta \\ 0 & 1 \end{pmatrix} \quad \delta \ll 1 \\ \text{small mixing}$$
$$M = \begin{pmatrix} M_1 & 0 \\ 0 & M_2 \end{pmatrix} \quad \text{no mixing}$$

$$y_\nu^T M^{-1} y_\nu = \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix} \frac{\delta^2}{M_1} + \begin{pmatrix} 0 & 0 \\ 0 & 1 \end{pmatrix} \frac{1}{M_2}$$
$$\approx \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix} \frac{\delta^2}{M_1} \quad \text{for } \frac{M_1}{M_2} \ll \delta^2$$

---

The (out-of equilibrium, CP-violating) decay of heavy right-handed neutrinos in the early universe might generate a net asymmetry between leptons and anti-leptons. Subsequent SM interactions can partially convert it into the observed baryon asymmetry

$$\eta = \frac{(n_B - n_{\bar{B}})}{s} \approx 6 \times 10^{-10}$$

# weak point of the see-saw

full high-energy theory is difficult to test

$$L(\nu^c, l) = \nu^c y_\nu (\tilde{\Phi}^+ l) + \frac{1}{2} \nu^c M \nu^c + h.c.$$

depends on many physical parameters:

3 (small) masses + 3 (large) masses

3 (L) mixing angles + 3 (R) mixing angles

6 physical phases = 18 parameters

the double of those

describing  $(L_{SM}) + L_5$ :

3 masses, 3 mixing angles

and 3 phases, as in lecture 1

few observables to pin down the extra parameters:  $\eta, \dots$

[additional possibilities exist under special conditions, e.g. Lepton Flavor Violation at observable rates]

easier to test the low-energy remnant  $L_5$

[which however is “universal” and does not imply the specific see-saw mechanism of Example 2]

look for a process where B-L is violated by 2 units. The best candidate is

$0\nu\beta\beta$  decay:  $(A, Z) \rightarrow (A, Z+2) + 2e^-$

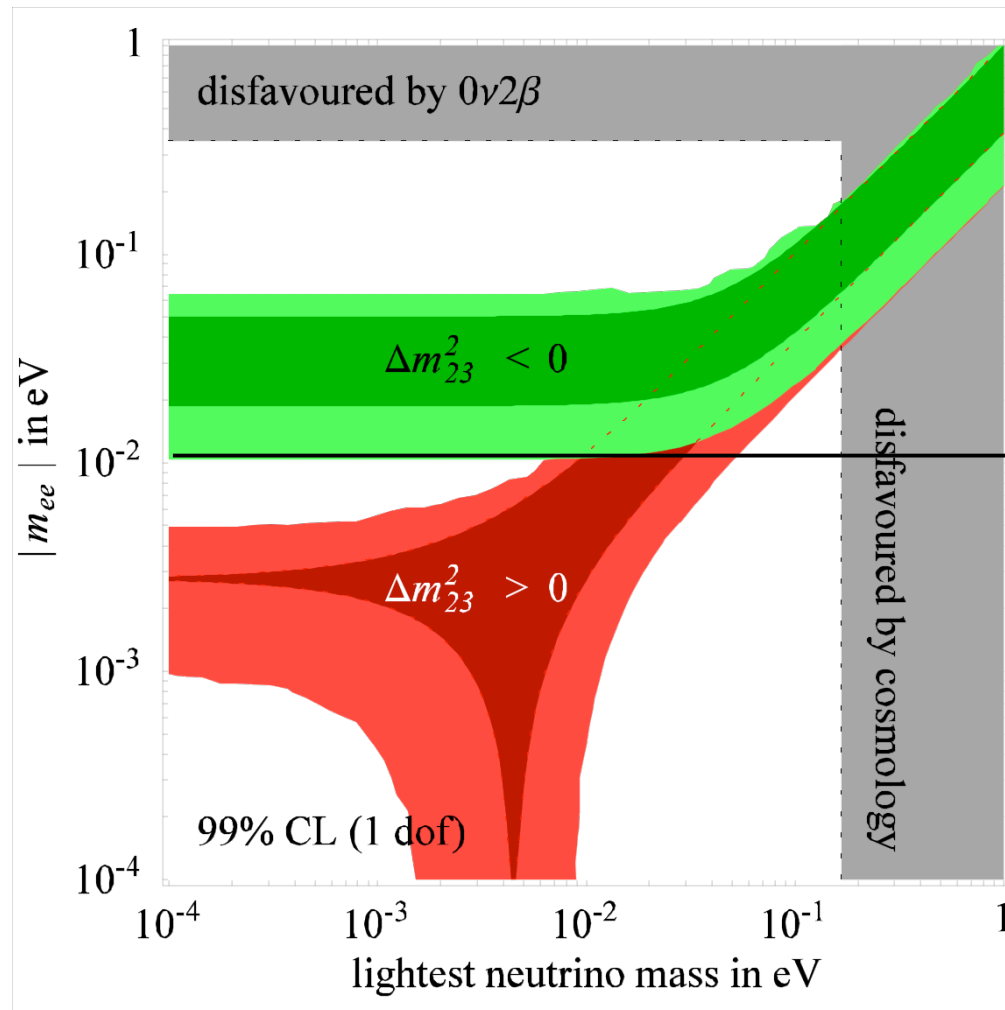
this would discriminate  $L_5$  from other possibilities, such as Example 1.

The decay in  $0\nu\beta\beta$  rates depend on the combination

$$|m_{ee}| = \left| \sum_i U_{ei}^2 m_i \right|$$

$$|m_{ee}| = \left| \cos^2 \vartheta_{13} (\cos^2 \vartheta_{12} m_1 + \sin^2 \vartheta_{12} e^{2i\alpha} m_2) + \sin^2 \vartheta_{13} e^{2i\beta} m_3 \right|$$

[notice the two phases  $\alpha$  and  $\beta$ , not entering neutrino oscillations]



from the current knowledge of  $(\Delta m_{ij}^2, \vartheta_{ij})$  we can estimate the expected range of  $|m_{ee}|$

future expected sensitivity on  $|m_{ee}|$

10 meV

a positive signal would test both  $L_5$  and the absolute mass spectrum at the same time!

# Flavor symmetries I (the hierarchy puzzle)

hierarchies in fermion spectrum

quarks	$\frac{m_u}{m_t} \ll \frac{m_c}{m_t} \ll 1 \quad \frac{m_d}{m_b} \ll \frac{m_s}{m_b} \ll 1$	$ V_{ub}  \ll  V_{cb}  \ll  V_{us}  \equiv \lambda < 1$
leptons	$\frac{m_e}{m_\tau} \ll \frac{m_\mu}{m_\tau} \ll 1$	$\frac{\Delta m_{sol}^2}{\Delta m_{atm}^2} = (0.025 \div 0.049) \approx \lambda^2 \ll 1 \quad (2\sigma)$ $ U_{e3}  < 0.18 \leq \lambda \quad (2\sigma)$

call  $\xi_i$  the generic small parameter. A modern approach to understand why  $\xi_i \ll 1$  consists in regarding  $\xi_i$  as small breaking terms of an approximate flavour symmetry. When  $\xi_i=0$  the theory becomes invariant under a flavour symmetry F

Example: why  $y_e \ll y_{top}$ ? Assume  $F=U(1)_F$

$F(t)=F(t^c)=F(h)=0$	$y_{top}(h+v)t^c t$	allowed
$F(e^c)=p>0 \quad F(e)=q>0$	$y_e(h+v)e^c e$	breaks $U(1)_F$ by $(p+q)$ units
if $\xi = \langle \varphi \rangle / \Lambda < 1$ breaks $U(1)$ by one negative unit		$y_e \approx O(\xi^{p+q}) \ll y_{top} \approx O(1)$

provides a qualitative picture of the existing hierarchies in the fermion spectrum

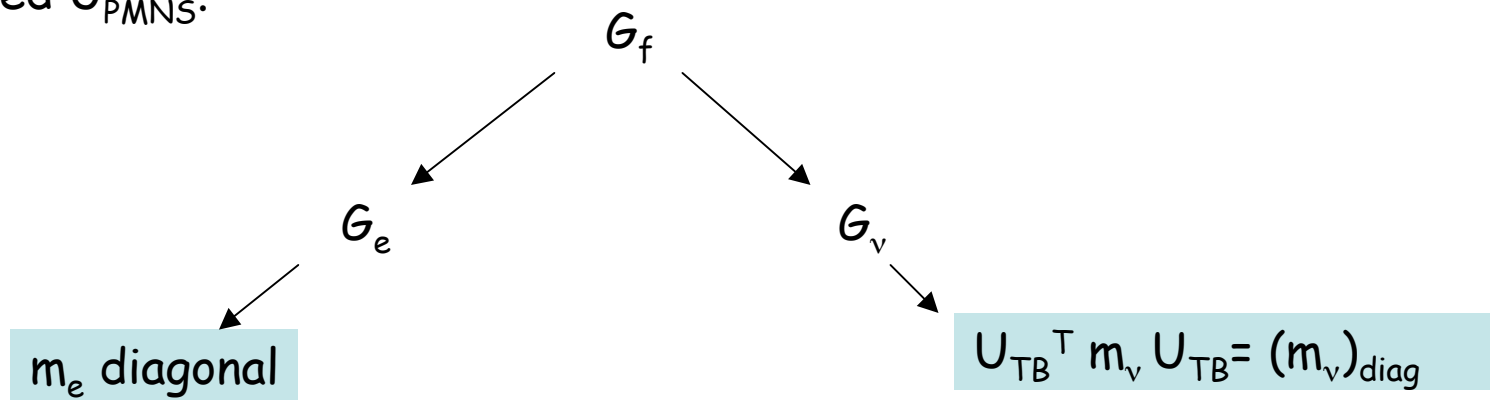
# Flavor symmetries II (the lepton mixing puzzle)

why  $U_{PMNS} \approx U_{TB} \equiv \begin{pmatrix} \frac{2}{\sqrt{6}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & -\frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix} ?$

[TB=TriBimaximal]

$$U_{PMNS} = U_e^\dagger U_\nu$$

Consider a flavor symmetry  $G_f$  such that  $G_f$  is broken into two different subgroups:  $G_e$  in the charged lepton sector, and  $G_\nu$  in the neutrino sector.  $m_e$  is invariant under  $G_e$  and  $m_\nu$  is invariant under  $G_\nu$ . If  $G_e$  and  $G_\nu$  are appropriately chosen, the constraints on  $m_e$  and  $m_\nu$  can give rise to the observed  $U_{PMNS}$ .



The simplest example is based on a small discrete group,  $G_f=A_4$ . It is the subgroup of  $SO(3)$  leaving a regular tetrahedron invariant. The elements of  $A_4$  can all be generated starting from two of them:  $S$  and  $T$  such that

$$S^2 = T^3 = (ST)^3 = 1$$

$S$  generates a subgroup  $Z_2$  of  $A_4$   
 $T$  generates a subgroup  $Z_3$  of  $A_4$

simple models have been constructed where  $G_e=Z_3$  and  $G_\nu=Z_2$  and where the lepton mixing matrix  $U_{PMNS}$  is automatically  $U_{TB}$ , at the leading order in the SB parameters. Small corrections are induced by higher order terms.

the generic predictions of this approach is that  $\theta_{13}$  and  $(\theta_{23}-\pi/4)$  are very small quantities, of the order of few percent: testable in a not-so-far future.



# Conclusion (theory)

theory of neutrino masses

it does not exist! Neither for neutrinos nor for charged fermions. We lack a **unifying principle**.

like weak interactions before the **electroweak theory**

$SU(2)_L \otimes U(1)_Y$   
gauge invariance

all fermion-gauge boson interactions in terms of 2 parameters:  $g$  and  $g'$

?

Yukawa interactions between fermions and spin 0 particles: many free parameters (up to 22 in the SM!)

only few ideas and prejudices about neutrino masses and mixing angles

caveat: several prejudices turned out to be wrong in the past!

- $m_\nu \approx 10$  eV because is the cosmologically relevant range
- solution to solar is MSW Small Angle
- atmospheric neutrino problem will go away because it implies a large angle